Particle physics: practice 1 Relativistic Kinematics

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#### Eötvös Loránd University (ELTE) September 11, 2023

## Special relativity

• Newton to Einstein: the problem of causality. Information can't be transferred instantaneously.



- "Special": no acceleration. "General": acceleration  $\equiv$  gravity.
- HEP can be thought of as the ultimate testbed for SR, since it's bread and butter for us (all particles relativistic).
- https://pdg.lbl.gov/2019/reviews/rpp2019-rev-kinematics.pdf

## Relativistic kinematics: Minkowski space

Relativistic theories are conveniently formulated in Minkowski space

Minkowski space =  $\mathbb{R}^4$  + Minkowski (pseudo)metric

Euclidean space =  $\mathbb{R}^3$  + Euclidean metric Distance between points in E. space:  $d(\vec{x}, \vec{y}) = (\vec{x} - \vec{y})^2 = (\vec{x} - \vec{y})_i (\vec{x} - \vec{y})_j \delta_{ij}$ Latin indices 1,...,3, sum over repeated indices understood Invariant under translations  $\vec{x} \to \vec{x} + \vec{a}$  and rotations  $\vec{x} \to R\vec{x}$ Point in Minkowski space (=event):  $X^{\mu}, \mu = 0, 1, 2, 3$ 

$$X^{\mu} = (ct, \vec{x}) = (t, \vec{x})$$

In Minkowski space distances replaced by interval

$$\Delta s^2 \equiv (X - Y)^2 \equiv (X - Y)^{\mu} (X - Y)^{\nu} g_{\mu\nu} \equiv (X - Y)^{\mu} (X - Y)_{\mu}$$
$$= (X^0 - Y^0)^2 - (\vec{X} - \vec{Y})^2$$

Minkowski metric tensor:  $g_{\mu\nu} = \text{diag}(1, -1, -1, -1)$ 

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## LORENTZ TRANSFORMATION IN MINKOWSKI SPACE



• Boost will shift the event along the hyperbolas in Minkowski space.

## THE LIGHTCONE

Contravariant vectors:  $X^{\mu} = (X^{0}, \vec{X})$ Covariant vectors:  $X_{\mu} = g_{\mu\nu}X^{\nu} = (X^{0}, -\vec{X})$ Indices lowered by  $g_{\mu\nu}$  and raised by  $g^{\mu\nu} = \text{diag}(1, -1, -1, -1)$   $g^{\mu\nu}$  defined by  $g^{\mu\rho}g_{\rho\nu} = \delta^{\mu}{}_{\nu}$ Minkowski scalar product  $X \cdot Y \equiv X^{\mu}Y^{\nu}g_{\mu\nu} = X^{\mu}Y_{\mu} = X^{0}Y^{0} - \vec{X} \cdot \vec{Y}$  $\vec{X} \cdot \vec{Y}$ : three-dimensional Euclidean scalar product

Interval is not a distance because it is not positive-definite:

- $\Delta s^2 > 0$  timelike interval  $X^2 > 0$  timelike vector
- $\Delta s^2 < 0$  spacelike interval  $X^2 < 0$  spacelike vector
- $\Delta s^2 = 0$  lightlike or null interval  $X^2 = 0$  lightlike or null vector

# THE LIGHTCONE (CONTD.)



• For a fixed event X

(Y − X)<sup>2</sup> = 0, Y<sup>0</sup> − X<sup>0</sup> > 0: forward (future) lightcone of X
(Y − X)<sup>2</sup> = 0, Y<sup>0</sup> − X<sup>0</sup> < 0: backward (past) lightcone of X</li>
(Y − X)<sup>2</sup> > 0, Y<sup>0</sup> − X<sup>0</sup> > 0: future of X (inside future lightcone)
(Y − X)<sup>2</sup> > 0, Y<sup>0</sup> − X<sup>0</sup> < 0: past of X (inside past lightcone)</li>

## TIMELIKE VECTORS, $X^2 > 0$



- Causally connected. Can go to a frame where the two events occur at the same point in space but at two times. Unique time ordering.
- This special frame  $\Delta t$  is called *proper time*. Eg. decay of an unstable particle in the particle's restframe.

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# Spacelike vectors, $X^2 < 0$



• Can go to a frame where two events at the same time occur at two places. But this simultaneity is frame-dependent. No unique time-ordering.

• Eg., virtual particles have spacelike 4-momentum

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#### LORENTZ TRANSFORMATIONS

Principles of special relativity:

- homogeneity and isotropy of space
- equivalence of all inertial reference frames

= travelling at a relative constant speed

• constancy of speed of light

 $\Rightarrow$  equivalent frames are related by a Lorentz transformation  $X' = \Lambda X$ : linear transformation that leaves every interval invariant

$$(X' - Y')^2 = (X - Y)^2 \qquad \forall X, Y$$

$$\Rightarrow X'^2 + Y'^2 - 2X' \cdot Y' = X^2 + Y^2 - 2X \cdot Y \qquad \forall X, Y$$
$$\Rightarrow X' \cdot Y' = X \cdot Y \qquad \forall X, Y$$

In components  $X^{\prime\mu} = \Lambda^{\mu}_{\ \alpha} X^{\alpha}$ 

$$g_{\alpha\beta}X^{\alpha}Y^{\beta} = g_{\mu\nu}X^{\prime\mu}Y^{\prime\nu} = g_{\mu\nu}\Lambda^{\mu}{}_{\alpha}\Lambda^{\nu}{}_{\beta}X^{\alpha}Y^{\beta} \qquad \forall X, Y$$
$$\Longrightarrow g_{\alpha\beta} = g_{\mu\nu}\Lambda^{\mu}{}_{\alpha}\Lambda^{\nu}{}_{\beta}$$

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## LORENTZ TRANSFORMATIONS (CONTD.)

Using matrix notation 
$$\mathbf{\Lambda}_{\mu\alpha} = \Lambda^{\mu}{}_{\alpha}, \ \mathbf{g}_{\mu\nu} = g_{\mu\nu}, \ \mathbf{g}_{\mu\nu}^{-1} = g^{\mu\nu} = \mathbf{g}_{\mu\nu}$$
$$\mathbf{g} = \mathbf{\Lambda}^{T} \mathbf{g} \mathbf{\Lambda}$$

 $(\det \Lambda)^2 = 1 \Rightarrow \det \Lambda = \pm 1, \Lambda$  invertible

- det  $\Lambda = 1$ : proper transformations, leave orientation of space unchanged
- det  $\Lambda = -1$ : *improper* transformations invert the orientation of space

$$\begin{split} \mathbf{\Lambda}^{-1} &= \mathbf{g}^{-1} \mathbf{\Lambda}^T \mathbf{g} \text{ still a Lorentz transformation} \\ \bullet & \mathbf{g} = [\mathbf{\Lambda} \mathbf{\Lambda}^{-1}]^T \mathbf{g} [\mathbf{\Lambda} \mathbf{\Lambda}^{-1}] = \mathbf{\Lambda}^{-1T} [\mathbf{\Lambda}^T \mathbf{g} \mathbf{\Lambda}] \mathbf{\Lambda}^{-1} = \mathbf{\Lambda}^{-1T} \mathbf{g} \mathbf{\Lambda}^{-1} \\ \bullet & \mathbf{\Lambda}_{\alpha\beta}^{-1} = g^{\alpha\mu} \Lambda^{\nu}{}_{\mu} g_{\nu\beta} = \Lambda_{\beta}{}^{\alpha} \end{split}$$

From the  $\alpha = 0, \, \beta = 0$  component of the defining relation

$$1 = \Lambda^{0}{}_{0}\Lambda^{0}{}_{0} - \Lambda^{i}{}_{0}\Lambda^{i}{}_{0} \Longrightarrow \Lambda^{0}{}_{0}\Lambda^{0}{}_{0} = 1 + \Lambda^{i}{}_{0}\Lambda^{i}{}_{0} \ge 1$$
  
•  $\Lambda^{0}{}_{0} \ge 1$ : orthochronous (does not change the sign of time)  
•  $\Lambda^{0}{}_{0} \le -1$ : non-orthochronous (changes the sign of time)

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#### PROPER ORTHOCHRONOUS LORENTZ GROUP

Proper orthochronous Lorentz transformations = three-dimensional rotations (the SO(3) group) and boosts Most general transformation: rotation  $\times$  boost in x direction  $\times$  rotation

Boost along x: 
$$\Lambda^{\mu}{}_{\nu} = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \qquad \beta = \frac{v}{c} = v < 1$$
  
 $\gamma = \frac{1}{\sqrt{1 - \beta^2}}$ 

Boost in general direction  $\vec{n}$ : rotate  $\vec{n}$  to x, boost, rotate back Coordinates in the new frame:

$$ct' = \gamma(ct - \beta x)$$
  $x' = \gamma(x - \beta ct)$   
 $y' = y$   $z' = z$ 

 $\Rightarrow \text{ relates } R \text{ to } R' \text{ moving with speed } \beta \text{ in the } negative x \text{ direction} \\ \text{Nonrelativistic limit } \beta = v/c \ll 1 \Rightarrow \text{Galilei transformations} \\ ct' = ct \qquad x' = x - vt \end{aligned}$ 

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## Full Lorentz group

Most general Lorentz transformation = proper orthochronous transformation times P (parity), T (time reversal), or PT

$$P^{\mu}_{\ \nu} = \text{diag}(1, -1, -1, -1) \qquad T^{\mu}_{\ \nu} = \text{diag}(-1, 1, 1, 1)$$

	$\det \Lambda = 1$		$\det \Lambda = -1$
${\Lambda^0}_0 \geq 1$	proper orthochronous	$\Rightarrow P$	improper orthochronous
	$\Downarrow_T$		$\downarrow_T$
${\Lambda^0}_0 \leq -1$	proper non-orthochronous	$\Rightarrow P$	improper non-orthochronous

#### RAPIDITY AND PSEUDORAPIDITY

$$\begin{pmatrix} ct' \\ x' \end{pmatrix} = \begin{bmatrix} \cosh \zeta & -\sinh \zeta \\ -\sinh \zeta & \cosh \zeta \end{bmatrix} \begin{pmatrix} ct \\ x \end{pmatrix}$$

•  $\tanh \zeta \equiv \beta$ .  $\zeta = \frac{1}{2} \ln \frac{E + |\mathbf{p}|c}{E - |\mathbf{p}|c}$  is called the rapidity for the boost.

- Show that the transformation is  $e^{\mathcal{Z}\zeta}$  where  $\mathcal{Z} = \begin{pmatrix} 0 & -1 \\ -1 & 0 \end{pmatrix}$
- Use this to show that rapidities are additive for two subsequent boosts.
- At colliders, rapidity is  $y = \frac{1}{2} \ln \frac{E+p_L}{E-p_L}$ , while pseudorapidity  $\eta = \frac{1}{2} \ln \frac{|\mathbf{p}|+p_L}{|\mathbf{p}|-p_L}$

## RAPIDITY AND PSEUDORAPIDITY (CONTD.)



- . Pseudorapidity  $\eta = \frac{1}{2} \ln \frac{|\mathbf{p}| + p_L}{|\mathbf{p}| p_L} = -\ln [\tan(\theta/2)]$  is purely angular term. Agrees with the usual rapidity definition in the limit  $p_T \gg m$ .
- Colliders typically use  $\Delta R \equiv \sqrt{(\Delta \eta)^2 + (\Delta \phi)^2}$  as 3-d angular separation between particles/jets.
- $p_L = p_T \sinh \eta$  and  $|\mathbf{p}| = p_T \cosh \eta$

#### POINT PARTICLES: KINEMATICS

Trajectory  $X^{\mu}(t)$  of point particle; over infinitesimal dt,  $X^{\mu} \to X^{\mu} + dX^{\mu}$ 

$$\begin{aligned} X^{\mu}(t) &= (ct, \vec{x}(t)) = (t, \vec{x}(t)) \\ dX^{\mu}(t) &= (dt, d\vec{x}(t)) = dt(1, \frac{d\vec{x}}{dt}(t)) = dt(1, \vec{v}(t)) \end{aligned}$$

Empirical fact: for massive particles  $\vec{v}^2 < 1$ , for massless particles  $\vec{v}^2 = 1$ 

$$(dX)^2 = dX^{\mu} dX_{\mu} = dt^2 (1 - \vec{v}^2) \ge 0$$
 (timelike)  
 $\frac{dX^{\mu}}{dt}(t) = (1, \vec{v}(t))$ 

 $\frac{dX^{\mu}}{dt}$  not a Lorentz vector:  $dX^{\mu} = \text{vector}, dt \neq \text{scalar}$ Massive particle  $\vec{v}^2 < 1$ :  $\exists$  reference frame in which  $\vec{v} = 0$  (rest frame)

$$X^{\mu}_{\rm rest}(\tau) = (\tau, \vec{0})$$

 $\tau$ : proper time (time measured in the particle's rest frame)

$$(dX_{\text{rest}})^2 = d\tau^2 = (dX)^2 = dt^2(1 - \vec{v}^2) = \frac{dt^2}{\gamma^2}$$

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## POINT PARTICLES: KINEMATICS (CONTD.)

Proper time:

• 
$$d\tau^2 = \frac{dt^2}{\gamma^2} \Rightarrow |dt| > |d\tau|$$
 (time-dilation effect)

 determine the elapsed proper time by going over to the instantaneous rest frame of the particle ⇒ twins' paradox

$$\tau = \int d\tau = \int_{t_0}^t dt' \sqrt{1 - \vec{v}^2(t')} \le t - t_0$$

• true scalar  $\Rightarrow \frac{d^n X^{\mu}}{d\tau^n}$  are true vectors

Four-velocity

$$u^{\mu} \equiv \frac{dX^{\mu}}{d\tau} = \left(\frac{dt}{d\tau}, \frac{d\vec{x}}{d\tau}\right) = \left(\gamma, \gamma \frac{d\vec{x}}{dt}\right) = \left(\gamma, \gamma \vec{v}\right) = \left(\gamma, \gamma \vec{\beta}\right)$$

Four-momentum (vector  $u^{\mu}$  times scalar m)

$$p^{\mu} \equiv m u^{\mu} = (\gamma m, \gamma m \vec{\beta})$$
$$p^{0} = m \gamma = \frac{m}{\sqrt{1 - \vec{v}^{2}}} = E \qquad p^{i} = m \gamma \vec{\beta}^{i} = \frac{m \vec{v}^{i}}{\sqrt{1 - \vec{v}^{2}}} = \vec{p}^{i}$$

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# POINT PARTICLES: FOUR-MOMENTUM IN THE NR LIMIT

Do  $E, \vec{p}$  match their non-relativistic definition when  $\frac{|\vec{v}|}{c} \ll 1$ ? Needs reinstating powers of c

$$p^{0} = mc \frac{1}{\sqrt{1 - \left(\frac{\vec{v}}{c}\right)^{2}}} = mc \left(1 + \frac{1}{2} \left(\frac{\vec{v}}{c}\right)^{2} + \mathcal{O}(\left(\frac{v}{c}\right)^{4})\right)$$
$$\vec{p} = mc \frac{\frac{\vec{v}}{c}}{\sqrt{1 - \left(\frac{\vec{v}}{c}\right)^{2}}} = m\vec{v} \left(1 + \mathcal{O}(\left(\frac{v}{c}\right)^{2})\right)$$

Second line ok, first line times c

$$p^0 c = mc^2 + \frac{1}{2}m\vec{v}^2 + \ldots = E_0 + E_K^{\text{NR}} + \ldots$$

 $\Rightarrow$  NR kinetic energy  $E_K^{\text{NR}}$  of a particle plus rest energy  $E_0 = mc^2$ 

# Point particles: four-momentum for $m \neq 0$ and m = 0

Massive particles:  $p^2 = m^2 > 0$ 

$$p^{\mu} = m \frac{dX^{\mu}}{d\tau} = \left(\frac{E}{c}, \vec{p}\right) = (E, \vec{p}) = (p^0, \vec{p})$$

Mass = relativistic invariant

$$p^2 = m^2 \gamma^2 (1 - \vec{\beta}^2) = m^2 > 0$$
  $u^2 = \gamma^2 (1 - \vec{\beta}^2) = 1$ 

Trajectory always inside the forward lightcone

Any constant would do, but m is the constant such that total momentum  $\sum_i p_i = \sum_i m_i u_i$  of a system of particles is conserved Also: correct NR limit of  $p^{\mu} = m u^{\mu}$ 

Energy-momentum relation is called *dispersion relation* 

$$E^2 = \vec{p}^2 + m^2$$

Massless particles:  $p^2 = 0$ 

$$p^{\mu} = (\omega, \vec{k})$$
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#### KINEMATICS OF 2-PARTICLE SCATTERING

Two particle  $\rightarrow$  two particle scattering process  $a b \rightarrow c d$ Lab frame: one initial particle is at rest (= target)

$$p_a = (E_L, \vec{p}_L) \qquad p_b = (m_b, 0) p_c = (E_c, \vec{p}_c) \qquad p_d = (E_d, \vec{p}_d)$$

Scattering angle  $\theta_L$  in the lab: angle between trajectories of c and a

$$\cos \theta_L = \frac{\vec{p}_L \cdot \vec{p}_c}{|\vec{p}_L||\vec{p}_c|}$$

CM frame: vanishing total spatial momentum

$$p_a = (E_a^*, \vec{p}^*) \qquad p_b = (E_b^*, -\vec{p}^*) p_c = (E_c^*, \vec{p}'^*) \qquad p_d = (E_d^*, -\vec{p}'^*)$$

Scattering angle  $\theta^*$  in the CM: angle formed by the trajectories of a and c

$$\cos\theta^* = \frac{\vec{p}^* \cdot \vec{p}^{\prime*}}{|\vec{p}^*||\vec{p}^{\prime*}|}$$

Total center of mass energy  $\sqrt{s}$  = Lorentz invariant

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## KINEMATICS OF 2-PARTICLE SCATTERING (CONTD.)

LAB

$$p_{a} = (E_{L}, \vec{p}_{L}) \qquad p_{b} = (m_{b}, 0)$$
$$p_{c} = (E_{c}, \vec{p}_{c}) \qquad p_{d} = (E_{d}, \vec{p}_{d})$$

$$p_a = (E_a^*, \vec{p}^*) \qquad p_b = (E_b^*, -\vec{p}^*) p_c = (E_c^*, \vec{p}'^*) \qquad p_d = (E_d^*, -\vec{p}'^*)$$



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## KINEMATICS OF 2-PARTICLE SCATTERING (CONTD.)

$$p_a + p_b = p_c + p_d$$

- Four-momentum conservation implies  $E_{c,d}^*$ ,  $|\vec{p}_{c,d}^*| = |\vec{p}^{**}|$  determined uniquely in the CM, independent of  $\theta^*$
- $E_{c,d}$ ,  $|\vec{p}_{c,d}|$  and  $\theta_L$  in the lab by Lorentz transf., depend on  $\theta^*$   $p_b = p_c + p_d - p_a$   $p_b^2 = (p_c + p_d)^2 + p_a^2 - 2p_a \cdot (p_c + p_d)$   $m_b^2 = s + m_a^2 - 2E_a^*\sqrt{s}$  $E_a^* = \frac{s + m_a^2 - m_b^2}{2\sqrt{s}} \xrightarrow[a \leftrightarrow b]{} E_b^* = \frac{s + m_b^2 - m_a^2}{2\sqrt{s}}$

• CM energy squared s Lorentz invariant  $\Rightarrow E_a^*$  from  $E_L$  in the lab:

$$s = (p_a + p_b)^2 = m_a^2 + m_b^2 + 2p_a \cdot p_b = m_a^2 + m_b^2 + 2E_L m_b \Rightarrow E_L = \frac{s - m_a^2 - m_b^2}{2m_b}$$

• Exchanging 
$$a, b \leftrightarrow c, d$$
  
 $E_c^* = \frac{s + m_c^2 - m_d^2}{2\sqrt{s}}$   $E_d^* = \frac{s + m_d^2 - m_c^2}{2\sqrt{s}}$ 

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#### KINEMATICS OF 2-PARTICLE SCATTERING: CM

Center of mass energies:

$$E_a^* = \frac{s + m_a^2 - m_b^2}{2\sqrt{s}} \qquad E_b^* = \frac{s + m_b^2 - m_a^2}{2\sqrt{s}}$$
$$E_c^* = \frac{s + m_c^2 - m_d^2}{2\sqrt{s}} \qquad E_d^* = \frac{s + m_d^2 - m_c^2}{2\sqrt{s}}$$

Center of mass momentum magnitude  $|\vec{p}^*|$ :

$$\begin{split} |\vec{p}^{*}|^{2} &= E_{a}^{*2} - m_{a}^{2} = \frac{(s + m_{a}^{2} - m_{b}^{2})^{2} - 4sm_{a}^{2}}{4s} = \frac{s^{2} + (m_{a}^{2} - m_{b}^{2})^{2} - 2s(m_{a}^{2} + m_{b}^{2})}{4s} \\ &= \frac{(s - m_{a}^{2} - m_{b}^{2})^{2} - 4m_{a}^{2}m_{b}^{2}}{4s} = \frac{[s - (m_{a} + m_{b})^{2}][s - (m_{a} - m_{b})^{2}]}{4s} = \frac{\lambda(s, m_{a}^{2}, m_{b}^{2})}{4s} \\ |\vec{p}^{\prime *}|^{2} &= E_{c}^{*2} - m_{c}^{2} = \frac{(s + m_{c}^{2} - m_{d}^{2})^{2} - 4sm_{c}^{2}}{4s} = \frac{s^{2} + (m_{c}^{2} - m_{d}^{2})^{2} - 2s(m_{c}^{2} + m_{d}^{2})}{4s} \\ &= \frac{(s - m_{c}^{2} - m_{d}^{2})^{2} - 4m_{c}^{2}m_{d}^{2}}{4s} = \frac{[s - (m_{c} + m_{d})^{2}][s - (m_{c} - m_{d})^{2}]}{4s} = \frac{\lambda(s, m_{c}^{2}, m_{d}^{2})}{4s} \end{split}$$

Källén function:  $\lambda(x,y,z)=x^2+y^2+z^2-2xy-2yz-2zx$ 

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#### KINEMATICS OF 2-PARTICLE SCATTERING: LAB

Lab kinematics recovered from CM kinematics Given  $\vec{p}_{\text{lab,CM}}$ ,  $E_{\text{lab,CM}}$  total spatial momentum/total energy in lab/CM

$$\begin{aligned} |\vec{p}_{\rm CM}| &= 0 = \gamma_{\rm CM}(|\vec{p}_{\rm lab}| - \beta_{\rm CM}E_{\rm lab}) = \gamma_{\rm CM}(|\vec{p}_L| - \beta_{\rm CM}(m_b + E_L)) \\ \implies \beta_{\rm CM} = \frac{|\vec{p}_L|}{E_L + m_b} \end{aligned}$$

Inverse Lorentz transformation from CM to lab

$$E_{c,\text{lab}} = \gamma_{\text{CM}} (E_c^* + \beta_{\text{CM}} |\vec{p}'^*| \cos \theta^*) ,$$
$$|\vec{p}_{c,\text{lab}}| \cos \theta_L = \gamma_{\text{CM}} (|\vec{p}'^*| \cos \theta^* + \beta_{\text{CM}} E_c^*) ,$$
$$|\vec{p}_{c,\text{lab}}| \sin \theta_L = |\vec{p}'^*| \sin \theta^* ,$$

Transverse directions unaffected by Lorentz transformation, azimuthal angle transforms trivially

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#### EXAMPLE: PROTON-ANTIPROTON SCATTERING

For  $p\bar{p}$  scattering in circular collider,  $E_p = E_{\bar{p}} = 270 \text{ GeV}$ 

$$\Rightarrow \sqrt{s} = 540 \text{ GeV}$$

Let now p be at rest in the lab.

**Q.** What should be the energy  $E_L$  of  $\bar{p}$  in the lab to obtain the same s? **A.** CM energy square s is a relativistic invariant, can be evaluated in any reference frame; in the lab

$$s = (p_p + p_{\bar{p}})^2 = 2(m_p^2 + E_L m_p) = 2m_p(m_p + E_L)$$

Solve for  $E_L$  and impose  $\sqrt{s} = 540 \text{ GeV} (\gg m_p)$ 

$$E_L = \frac{s - 2m_p^2}{2m_p} \simeq \frac{s}{2m_p} \simeq \frac{(540)^2}{2} \text{ GeV} \simeq \frac{30}{2} \cdot 10^4 \text{ GeV} = 150 \text{ TeV} \quad (!!!)$$

In general total CM energy  $E_{\rm CM} \simeq \sqrt{2m_p E_L}$ 

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## MANDELSTAM VARIABLES (CONT.)

Convenient set of relativistic invariant variables for  $2 \rightarrow 2$  scattering

$$s \equiv (p_a + p_b)^2 = (p_c + p_d)^2$$
$$t \equiv (p_a - p_c)^2 = (p_b - p_d)^2$$
$$u \equiv (p_a - p_d)^2 = (p_b - p_c)^2$$



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## MANDELSTAM VARIABLES (CONT.)

Convenient set of relativistic invariant variables for  $2 \rightarrow 2$  scattering

$$s \equiv (p_a + p_b)^2 = (p_c + p_d)^2$$
  

$$t \equiv (p_a - p_c)^2 = (p_b - p_d)^2$$
  

$$u \equiv (p_a - p_d)^2 = (p_b - p_c)^2$$

• s = total CM energy squared

• t = square of four-momentum transfer from a to c

$$t = p_a^2 + p_c^2 - 2p_a \cdot p_c = m_a^2 + m_c^2 - 2(E_a^* E_c^* - |\vec{p}^*| |\vec{p}'^*| \cos \theta^*)$$

• u = square of four-momentum transfer from a to d

$$u = p_a^2 + p_d^2 - 2p_a \cdot p_d = m_a^2 + m_d^2 - 2(E_a^* E_d^* + |\vec{p}^*| |\vec{p}'^*| \cos \theta^*)$$
  
u obtained from t after  $m_c \to m_d$  and  $\cos \theta^* \to -\cos \theta^*$ 

Energies and magnitudes of momenta entirely determined by s and particle masses  $\Rightarrow t = t(s, \theta^*)$ , or instead  $\theta^* = \theta^*(s, t)$  and use s, t

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## MANDELSTAM VARIABLES (CONTD.)

Only two independent Mandelstam variables:

$$s + t + u = (p_a + p_b)^2 + (p_a - p_c)^2 + (p_a - p_d)^2$$
  
=  $m_a^2 + m_b^2 + m_c^2 + m_d^2 + 2p_a \cdot (p_a + p_b - p_c - p_d)$   
=  $m_a^2 + m_b^2 + m_c^2 + m_d^2$ 

Bounds on Mandelstam variables determine physical region for s, t, u

$$s \ge \max((m_a + m_b)^2, (m_c + m_d)^2)$$
  
$$t = (p_a - p_c)^2 = m_a^2 + m_c^2 - 2p_a \cdot p_c = 2(m_a^2 + m_c^2) - (p_a + p_c)^2$$
  
$$\le 2(m_a^2 + m_c^2) - (m_a + m_c)^2 = (m_a - m_c)^2$$

Similarly using  $p_b$  and  $p_d$ ; same approach for u

 $t \le \min((m_a - m_c)^2, (m_b - m_d)^2) \qquad u \le \min((m_a - m_d)^2, (m_b - m_c)^2)$ Lower bound from this and  $t|u = m_a^2 + m_b^2 + m_c^2 + m_d^2 - s - u|t$ 

$$t \ge \max(m_b^2 + m_c^2 + 2m_a m_d, m_a^2 + m_d^2 + 2m_b m_c) - s$$
$$u \ge \max(m_a^2 + m_c^2 + 2m_b m_d, m_b^2 + m_d^2 + 2m_a m_c) - s$$

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## MANDELSTAM VARIABLES (CONTD.)

Simplification if  $m_a = m_b$ ,  $m_c = m_d \Rightarrow E_a^* = E_b^* = E_c^* = E_d^* = \frac{\sqrt{s}}{2}$ 

$$t = m_a^2 + m_c^2 - \frac{s}{2} \left( 1 - \cos \theta^* \sqrt{1 - \frac{4m_a^2}{s}} \sqrt{1 - \frac{4m_c^2}{s}} \right)$$

If also  $m_a = m_c \equiv m$ 

$$t = 2m^2 - \frac{s}{2}\left(1 - \cos\theta^* \left(1 - \frac{4m^2}{s}\right)\right) = -\left(s - 4m^2\right)\sin^2\frac{\theta^*}{2}$$

$$s \ge 4m^2 \qquad -\left(s - 4m^2\right) \le t \le 0$$

- Upper limit: at threshold  $s = 4m^2$  or when  $\theta^* = 0$  (fwd scatter)
- Lower limit: when  $\theta^* = \pi$  (backscattering)
- In this case  $u(s, \theta^*) = t(s, \pi \theta^*) \Rightarrow$  same bound applies to u; role of  $\theta^* = 0$  and  $\theta^* = \pi$  exchanged

Relevant for

- elastic processes involving only one type of particles/antiparticles
- very high energy limit (masses negligible, particles  $\approx$  massless)

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#### EXAMPLE: PROTON-PROTON SCATTERING

Elastic pp scattering,  $\sqrt{s} = 53$  GeV

Differential cross section  $\frac{d\sigma}{dt}(t)$  has a peak at  $-t = t_0 = 1.81 \text{ GeV}^2$ E. Nagy *et al.*, Nucl. Phys. **B150** (1979) 221

 $\mathbf{Q}.$  What is the corresponding scattering angle in the CM?

**A.** Elastic scattering of identical particles,  $s/m_p^2 \gg 1$ 

$$-t = (s - 4m_p^2)\sin^2\frac{\theta^*}{2} \simeq s\sin^2\frac{\theta^*}{2}$$

$$\sin^2 \frac{\theta^*}{2} = -\frac{t}{s - 4m_p^2} = \frac{1.81}{53^2 - 4 \cdot 0.938^2} = \frac{1.81}{2805} = 6.45 \cdot 10^{-4}$$
$$\sin^2 \frac{\theta^*}{2} \simeq \frac{(\theta^*)^2}{4} \implies \theta^* \simeq 2\sqrt{5} \cdot 10^{-2} \simeq 5 \cdot 10^{-2}$$

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## MANDELSTAM PLANE



- Sides of eq. triangle: s = 0, t = 0 and u = 0 axes
- For appropriate side length  $s + t + u = m_a^2 + m_b^2 + m_c^2 + m_d^2$
- Physical region for the a + b → c + d process (equal masses) = wedge defined by the prolongation of the u and t axes

#### CROSSING SYMMETRY

QFT result: scattering amplitudes for  $a + b \rightarrow c + d$ ,  $a + \bar{c} \rightarrow \bar{b} + d$ ,  $a + \bar{d} \rightarrow c + \bar{b}$  are part of a single analytic function extending beyond physical momenta, and related to each other

$$\begin{split} &A_{ab\rightarrow cd}(p_a,p_b;p_c,p_d)\!=\!A_{a\bar{c}\rightarrow\bar{b}d}(p_a,\!-p_c;\!-p_b,p_d)\!=\!A_{a\bar{d}\rightarrow c\bar{b}}(p_a,\!-p_d;p_c,\!-p_b)\\ &\text{Use Mandelstam variables} \end{split}$$

 $\begin{aligned} a+b \rightarrow c+d & \mathcal{A}_{s}(s,t,u) = A_{ab \rightarrow cd}(p_{a},p_{b};p_{c},p_{d}) & s\text{-channel} \\ a+\bar{c} \rightarrow \bar{b}+d & \mathcal{A}_{t}(s_{t},t_{t},u_{t}) = A_{a\bar{c} \rightarrow \bar{b}d}(p_{a},p_{\bar{c}};p_{\bar{b}},p_{d}) & t\text{-channel} \\ a+\bar{d} \rightarrow c+\bar{b} & \mathcal{A}_{u}(s_{u},t_{u},u_{u}) = A_{a\bar{d} \rightarrow c\bar{b}}(p_{a},p_{\bar{d}};p_{c},p_{\bar{b}}) & u\text{-channel} \\ s = (p_{a}+p_{b})^{2} & t = (p_{a}-p_{c})^{2} & u = (p_{a}-p_{d})^{2} \\ s_{t} = (p_{a}+p_{\bar{c}})^{2} & t_{t} = (p_{a}-p_{\bar{b}})^{2} & u_{t} = (p_{a}-p_{d})^{2} \\ s_{u} = (p_{a}+p_{\bar{c}})^{2} & t_{u} = (p_{a}-p_{c})^{2} & u_{u} = (p_{a}-p_{\bar{b}})^{2} \end{aligned}$ 

Crossing-symmetry relations

$$\mathcal{A}_s(s,t,u) = \mathcal{A}_t(t,s,u) = \mathcal{A}_u(u,t,s)$$

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## CROSSING SYMMETRY (CONTD.)



$$\mathcal{A}_s(s,t,u) = \mathcal{A}_t(t,s,u) = \mathcal{A}_u(u,t,s)$$

- If s, t, u take physical values for the s-channel process a b → c d, crossing relations involve A<sub>t</sub> and A<sub>u</sub> at unphysical values of their arguments
- Relations fully meaningful if  $\mathcal{A}_s$  can be analytically continued outside the physical domain
- For equal masses, physical regions of  $\mathcal{A}_t$  and  $\mathcal{A}_u$  are  $s_t \geq 4m^2, t_t \leq 0$  and  $s_u \geq 4m^2, t_u \leq 0$ , but  $t \leq 0$  and  $s \geq 4m^2$

 $Physical \ regions = wedges \ outside \ Mandelstam \ triangle$ 

#### INVARIANT PHASE SPACE

States of spinless particle, mass m are characterised by four-momenta  $p^\mu$  with  $p^2=m^2$  and positive energy  $p^0\geq m>0$ 

One-particle phase space:

$$\{p\in \mathbb{R}^4 | p^2 - m^2 = 0\,, \ p^0 > 0\} \subset \mathbb{R}^4$$

Measure of infinitesimal element of phase space

$$d\Phi^{(1)} = \frac{d^4p}{(2\pi)^4} 2\pi\delta(p^2 - m^2)\theta(p^0)$$

- Manifestly invariant under orthochronous Lorentz transformations:  $p^2$  invariant, sign $(p^0)$  invariant under orthochronous transformations
- Overall scale appropriate for relativistic normalisation of one-particle states:  $\langle \vec{p}' | \vec{p} \rangle = (2\pi)^3 2 p^0 \delta^{(3)} (\vec{p}' \vec{p})$

## INVARIANT PHASE SPACE (CONTD.)

Recast  $d\Phi^{(1)}$  in more convenient form: for any f with simple zeros  $\{x_n\}$ 

$$\delta(f(x)) = \sum_{x_n, f(x_n) = 0} \frac{1}{|f'(x_n)|} \delta(x - x_n)$$

- multiply both sides by some function h(x), integrate over  $\mathbb{R}$ , show that one gets the same result
- divide R = (-∞, +∞) = ∪<sub>k</sub>I<sub>k</sub> with f(x) monotonic in I<sub>k</sub>
  ⇒ f invertible in I<sub>k</sub> and vanishes at most once (|f'| ≠ 0 there)
  set y = f(x) → x = f<sup>-1</sup>(y) in each I<sub>k</sub>

$$\begin{split} &\int_{-\infty}^{+\infty} dx \,\delta(f(x))h(x) = \sum_{k} \int_{I_{k}} dx \,\delta(f(x))h(x) \\ &= \sum_{k} \int_{f(I_{k})} dy \,\frac{1}{|f'(f^{-1}(y))|} \delta(y)h(f^{-1}(y)) \\ &= \sum_{k} \int_{0 \in f(I_{k})} dy \,\frac{1}{|f'(f^{-1}(0))|} \delta(y)h(f^{-1}(0)) = \sum_{n} \frac{1}{|f'(x_{n})|} h(x_{n}) \end{split}$$

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INVARIANT PHASE SPACE (CONTD.)

$$\begin{split} d\Phi^{(1)} &= \frac{d^4p}{(2\pi)^3} \delta(p^2 - m^2) \theta(p^0) = \frac{d^4p}{(2\pi)^3} \delta(p^{0\,2} - \vec{p}^2 - m^2) \theta(p^0) \\ &= \frac{d^4p}{(2\pi)^3} \frac{1}{2|p^0|} \left[ \delta(p^0 - \varepsilon(\vec{p}\,)) + \delta(p^0 + \varepsilon(\vec{p}\,)) \right] \theta(p^0) \\ &= \frac{d^4p}{(2\pi)^3} \frac{1}{2\varepsilon(\vec{p}\,)} \delta(p^0 - \varepsilon(\vec{p}\,)) \theta(p^0) = \frac{d^3p}{(2\pi)^3 2\varepsilon(\vec{p}\,)} \equiv d\Omega_p \\ \varepsilon(\vec{p}\,) \equiv \sqrt{\vec{p}^2 + m^2} \end{split}$$

*n*-particle phase space  $\subset \mathbb{R}^{4n}$  corresponding to four-momenta of *n* particles subjected to a constraint on the total four-momentum Measure of infinitesimal element:

$$d\Phi^{(n)} = \prod_{j=1}^{n} d\Omega_{p_j}(2\pi)^4 \delta^{(4)} \left( p_{\text{tot}} - \sum_{j=1}^{n} p_j \right)$$

Lorentz invariant:  $d\Omega_{p_i}$  Lorentz invariant

$$\delta^{(4)}(\Lambda P) = |\det \Lambda|^{-1} \delta^{(4)}(P) = \delta^{(4)}(P)$$

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#### INVARIANT PHASE SPACE: 2-PARTICLE CASE

 $\begin{aligned} \text{Total momentum } p_{\text{tot}} &= (E_{\text{tot}}, \vec{p}_{\text{tot}}), \text{ particle energies} \\ \varepsilon_i(\vec{p}) &= \sqrt{\vec{p}^2 + m_i^2} \\ d\Phi^{(2)} &= \frac{d^3 p_1}{(2\pi)^3 2\varepsilon_1(\vec{p}_1)} \frac{d^3 p_2}{(2\pi)^3 2\varepsilon_2(\vec{p}_2)} (2\pi)^4 \delta^{(4)}(p_{\text{tot}} - p_1 - p_2) \\ &= \frac{1}{(2\pi)^2} \frac{d^3 p_1}{2\varepsilon_1(\vec{p}_1)} \frac{d^3 p_2}{2\varepsilon_2(\vec{p}_2)} \delta^{(3)}(\vec{p}_{\text{tot}} - \vec{p}_1 - \vec{p}_2) \delta(E_{\text{tot}} - \varepsilon_1(\vec{p}_1) - \varepsilon_2(\vec{p}_2)) \end{aligned}$ 

Integrate trivially over  $\vec{p}_2$ , setting it equal to  $\vec{p}_2 = \vec{p}_{tot} - \vec{p}_1$ 

$$d\Phi^{(2)} = \frac{1}{(2\pi)^2} \frac{d^3 p_1}{2\varepsilon_1(\vec{p_1})} \frac{1}{2\varepsilon_2(\vec{p_{\text{tot}}} - \vec{p_1})} \delta(E_{\text{tot}} - \varepsilon_1(\vec{p_1}) - \varepsilon_2(\vec{p_{\text{tot}}} - \vec{p_1}))$$

To further integrate over  $|\vec{p_1}|$  requires changing variables, most easily done working in the CM

$$\vec{p}_{\rm tot,CM} = 0 \Rightarrow \vec{p}_{1\,CM} = -\vec{p}_{2\,CM}, \quad |\vec{p}_{1\,CM}| = |\vec{p}_{2\,CM}| = p$$

Dropping "CM" in the following

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#### INVARIANT PHASE SPACE: 2-PARTICLE CASE (CONTD.)

Delta function depends on  $E_{\text{tot}} - \varepsilon_1(p) - \varepsilon_2(p)$ 

Dropped vector sign on  $\pm \vec{p}$ 

$$\left|\frac{\Delta E}{\Delta Ep}\left[E_{\rm tot} - \varepsilon_1(p) - \varepsilon_2(p)\right]\right| = \left[\frac{p}{\varepsilon_1(p)} + \frac{p}{\varepsilon_2(p)}\right] = \frac{p}{\varepsilon_1(p)\varepsilon_2(p)}\left[\varepsilon_1(p) + \varepsilon_2(p)\right]$$

Changing variables to  $d^3p_1 = dpp^2 d\cos\theta^* d\phi^* = dpp^2 d\Omega^*$ 

$$d\Phi^{(2)} = \frac{1}{(2\pi)^2} \frac{dpp^2 d\Omega^*}{2\varepsilon_1(p)} \frac{1}{2\varepsilon_2(p)} \underbrace{\frac{\varepsilon_1(p)\varepsilon_2(p)}{p} \left[\varepsilon_1(p) + \varepsilon_2(p)\right]^{-1} \delta(p - p^*)}_{\delta(E_{\text{tot}} - \varepsilon_1(p) - \varepsilon_2(p))}$$
$$= \frac{d\Omega^*}{(2\pi)^2} \frac{p^*}{4(\varepsilon_1(p^*) + \varepsilon_2(p^*))} = \frac{d\Omega^*}{(2\pi)^2} \frac{p^*}{4E_{\text{tot}}^*} = \frac{d\Omega^*}{16\pi^2} \frac{p^*}{\sqrt{s}}$$
$$= \frac{d\Omega^*}{32\pi^2} \frac{\sqrt{\lambda(s, m_1^2, m_2^2)}}{s}$$
For equal masses  $\lambda(s, m^2, m^2) = s(s - 4m^2) \Rightarrow d\Phi^{(2)} = \frac{d\Omega^*}{32\pi^2} \sqrt{\frac{s - 4m^2}{s}}$ 

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