

Mass estimates of exoplanets based on the topology of interacting networks

Bolyai Seminar, 2024 April

Tamás Kovács, Eötvös University, Hungary

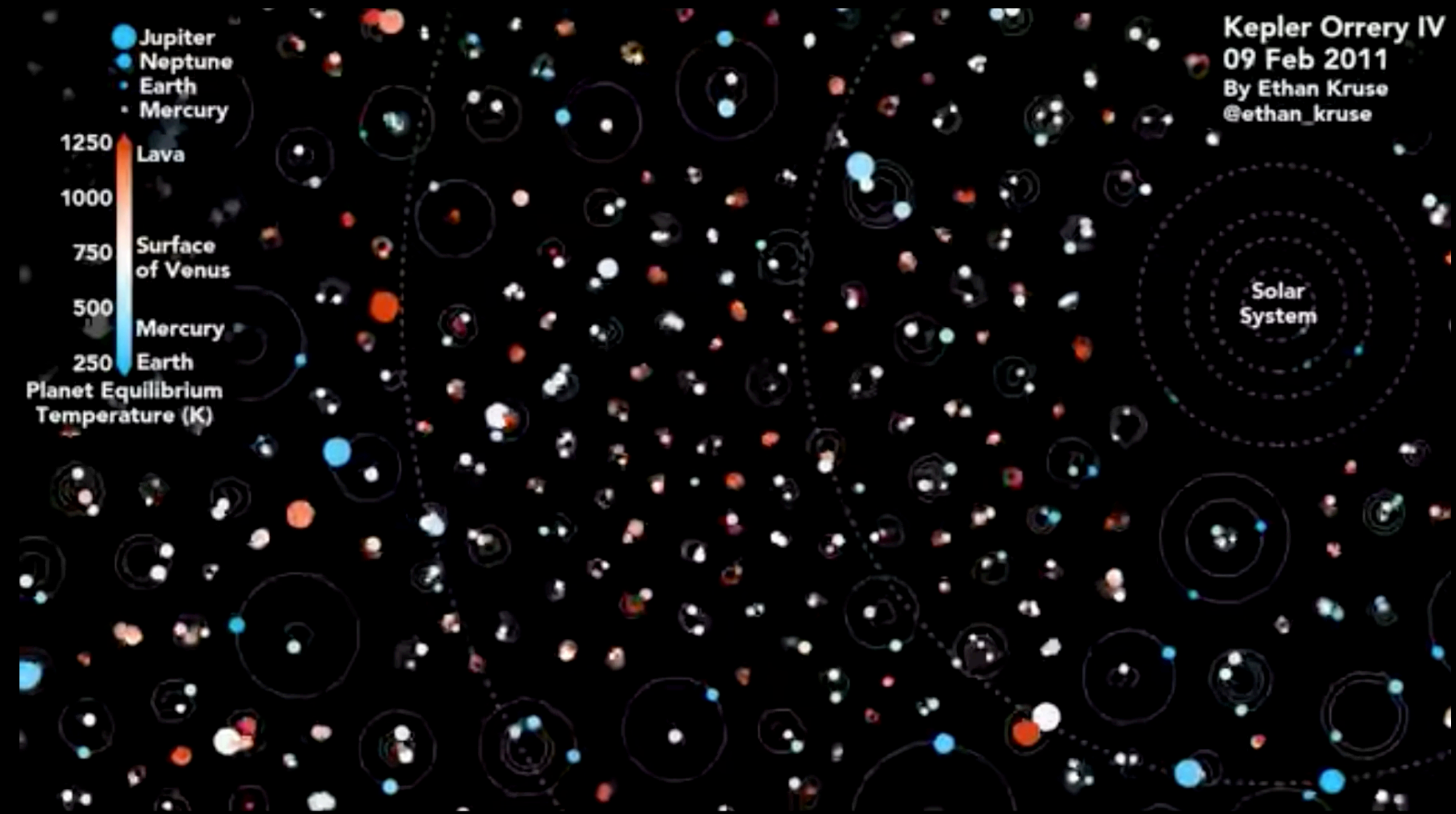
Motivation

Planet masses if we have

- no equations
- no initial conditions
- no numerical integration
- no best fit model & MCMC

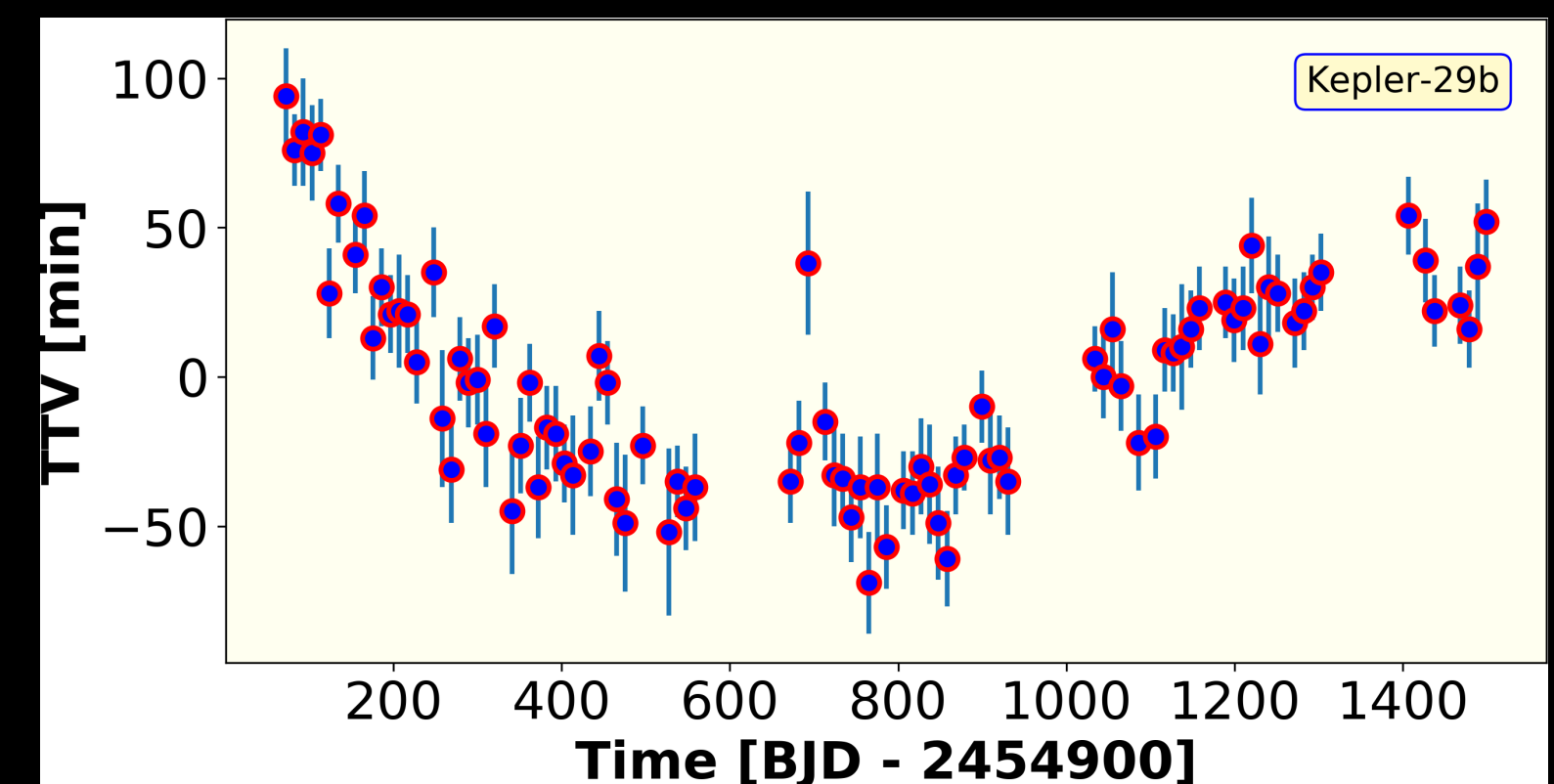
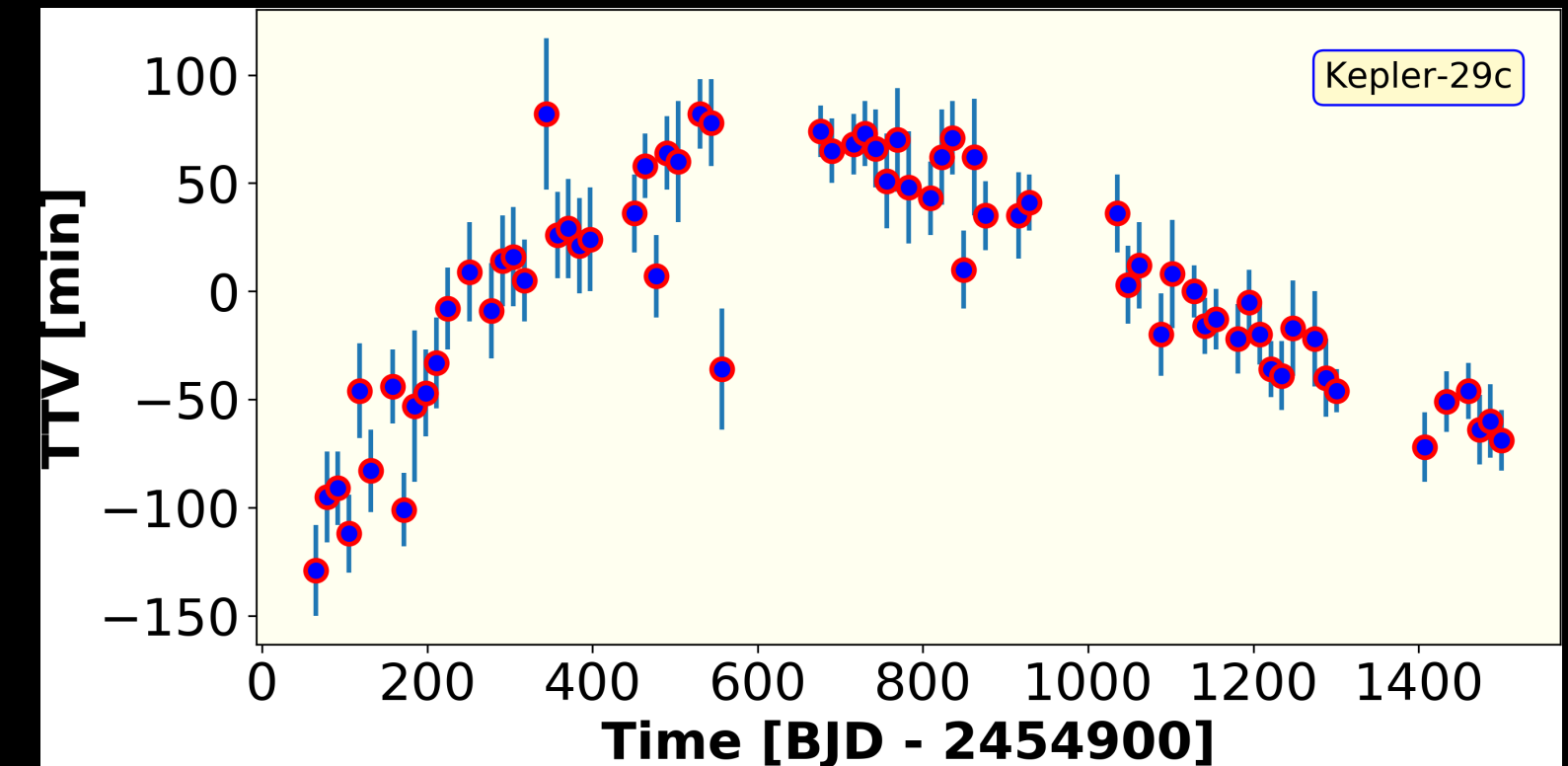
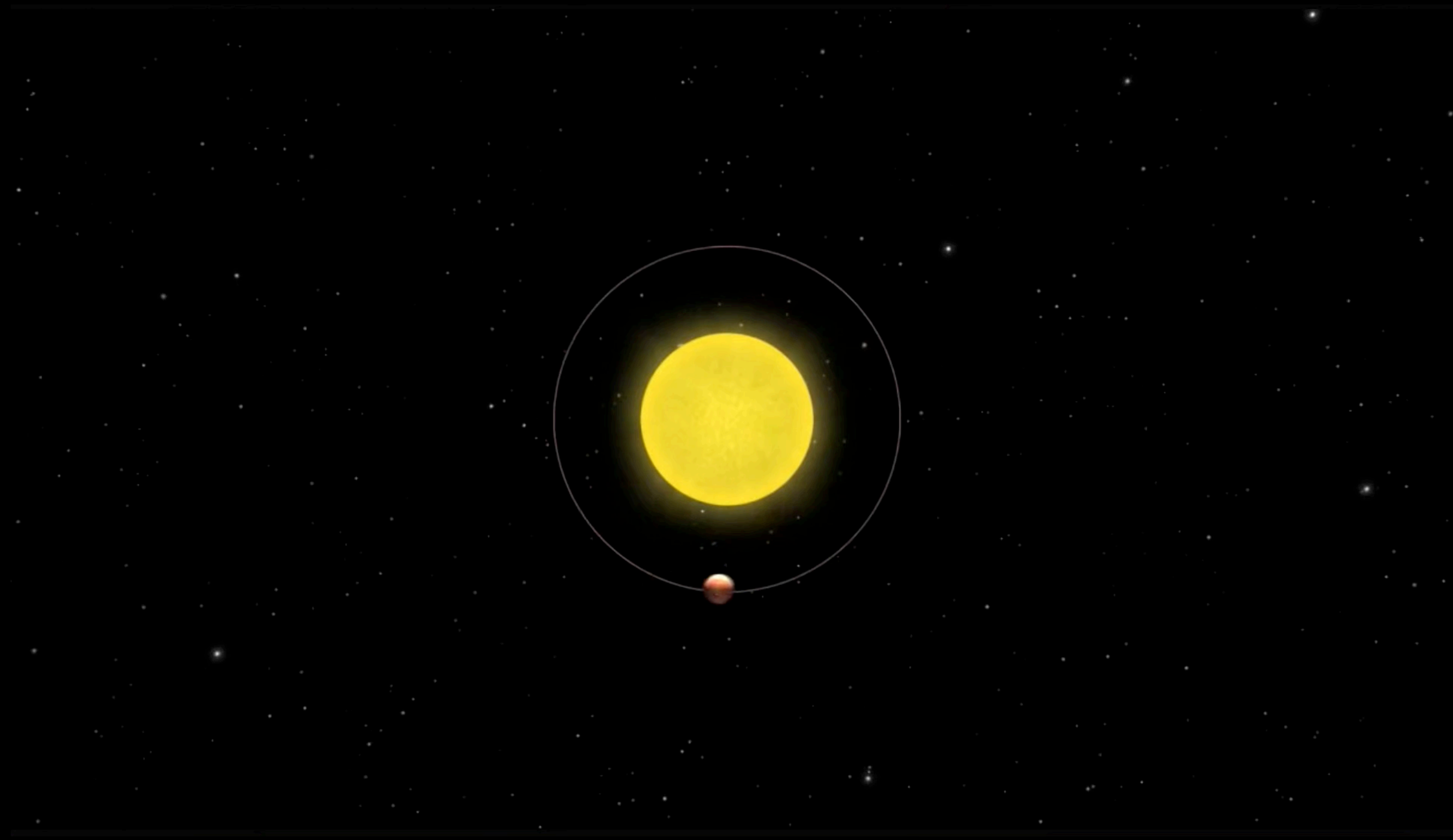
BUT we have

- a lot of data

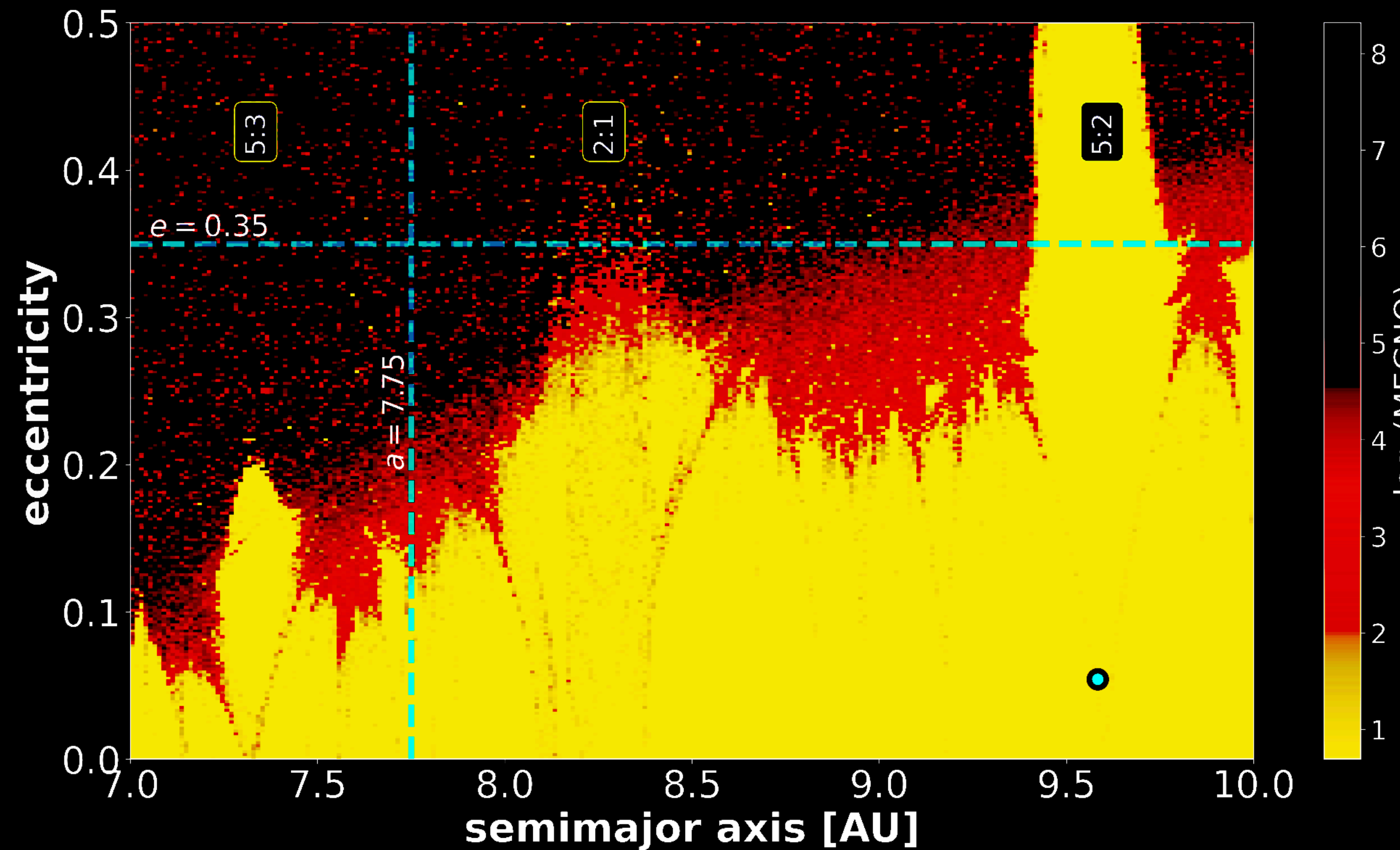
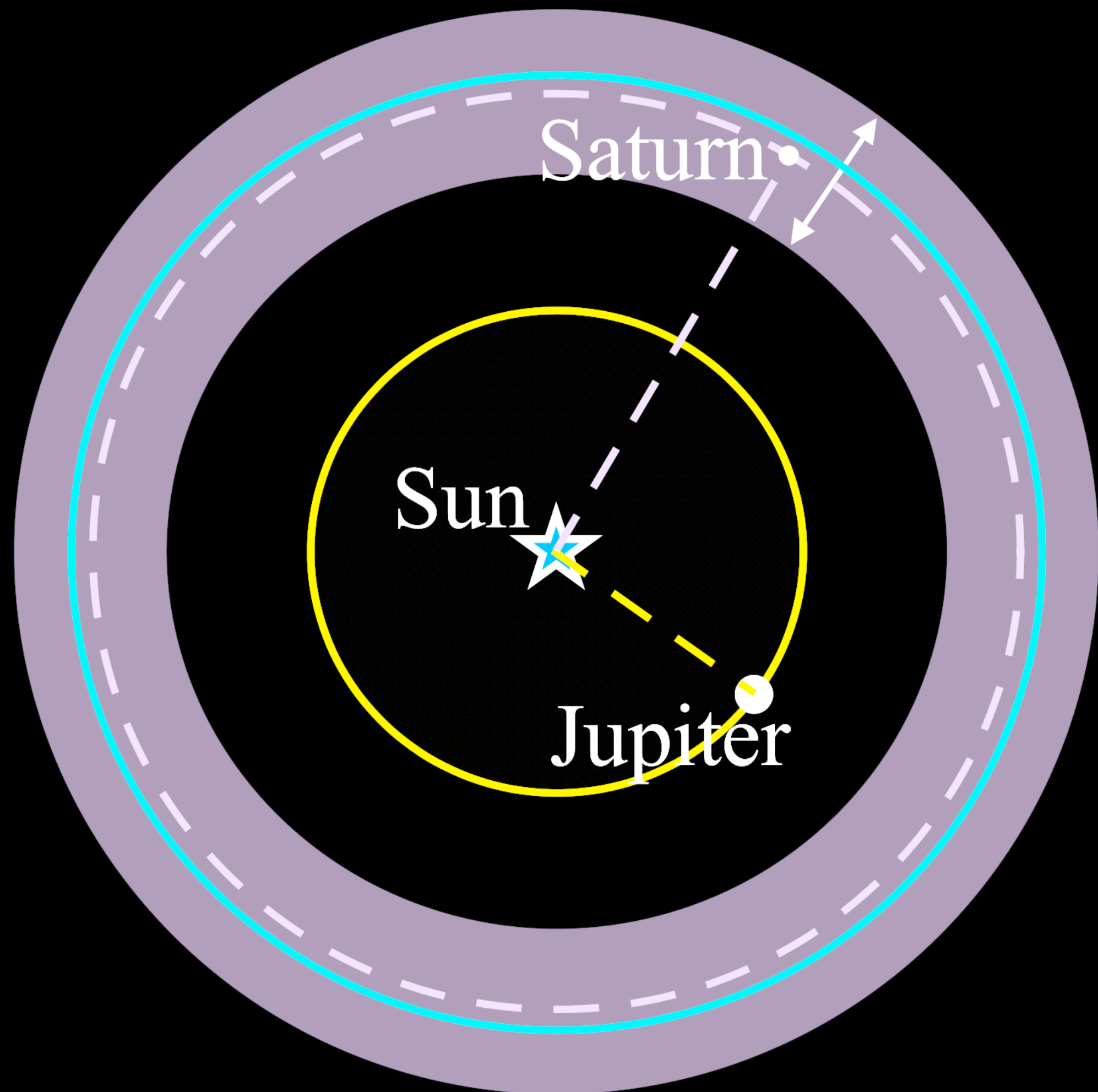


classical parameter estimation

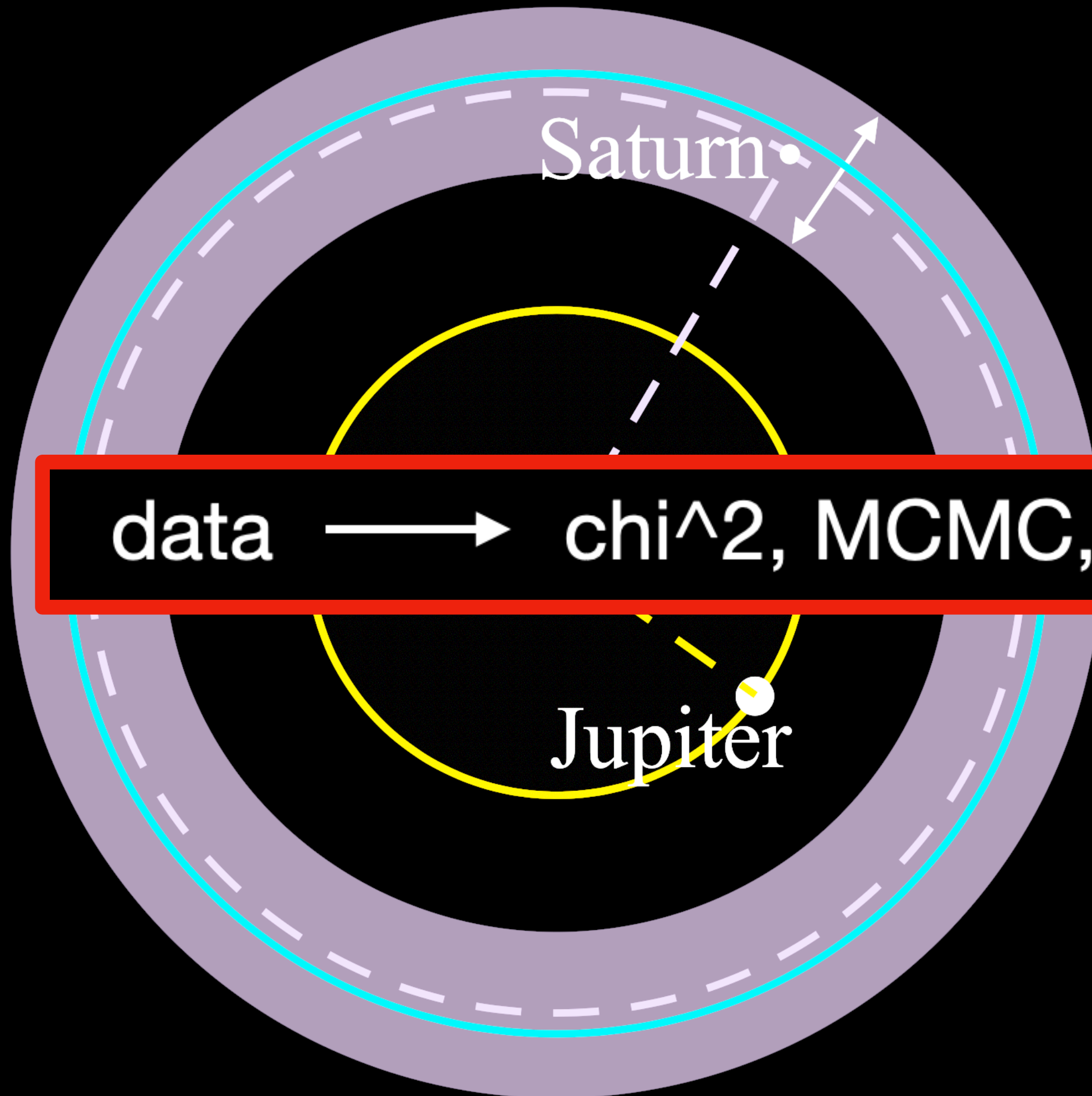
data \longrightarrow χ^2 , MCMC, n-body int. \longrightarrow system parameters



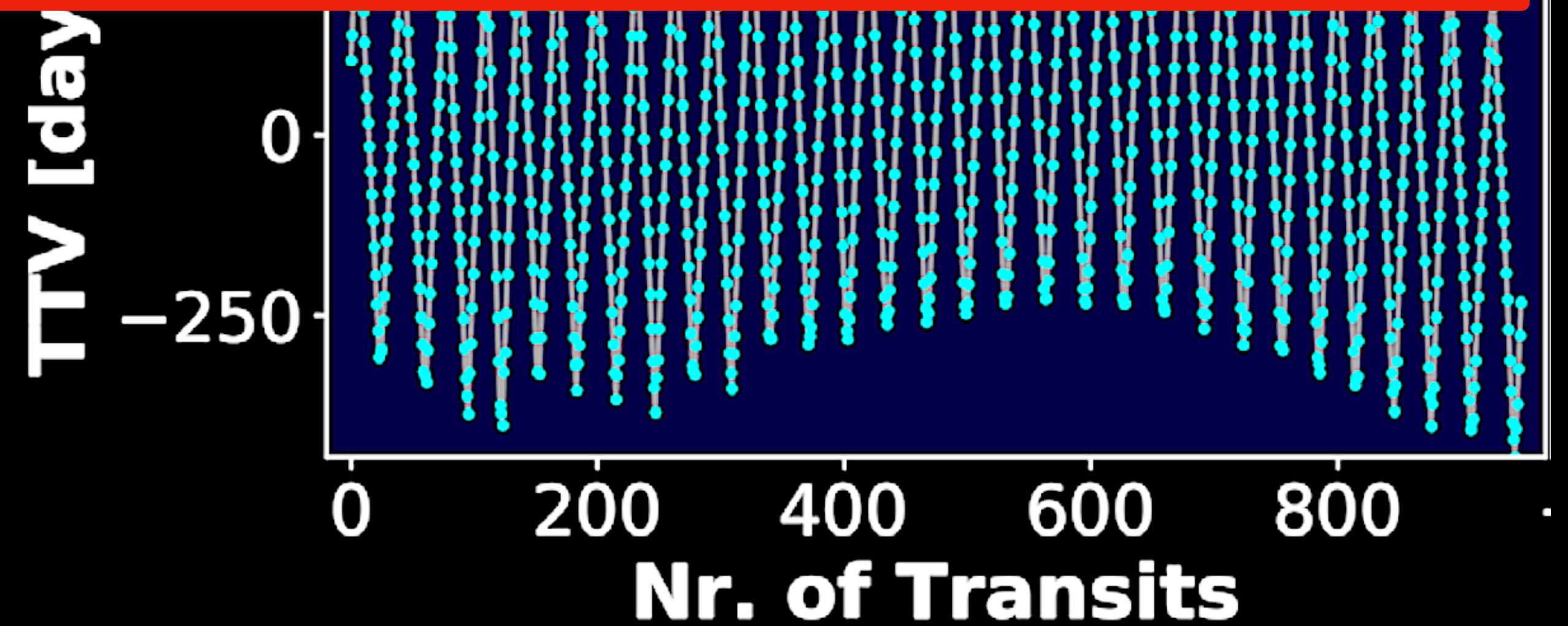
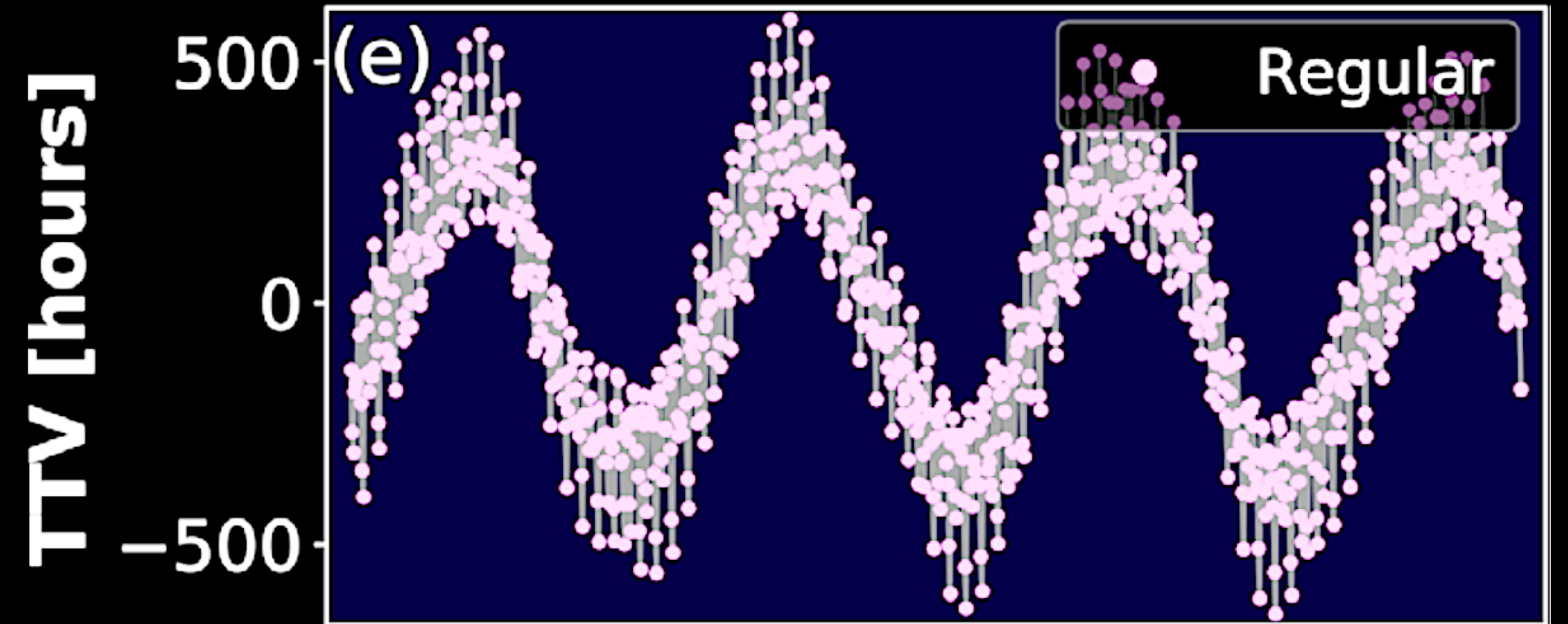
the case study of Jupiter & Saturn



the case study of Jupiter & Saturn



data \longrightarrow χ^2 , MCMC, n-body int. \longrightarrow system parameters



from time series to dynamics

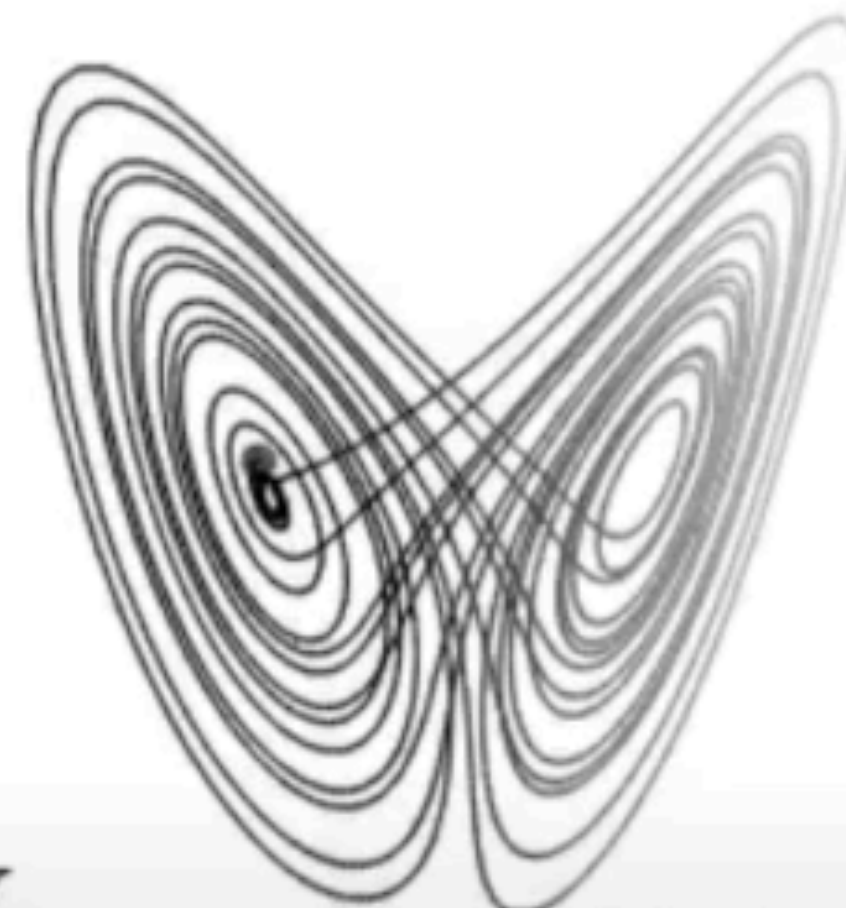
phase space
reconstruction
if no all the
components of
the state vector
are known

Takens' theorem in action for the Lorenz chaotic attractor

Z



Y



Pull up for precise seeking



$$\frac{dX}{dt} = -\sigma Y + \sigma X$$

$$\frac{dY}{dt} = -XZ + \rho X - Y$$

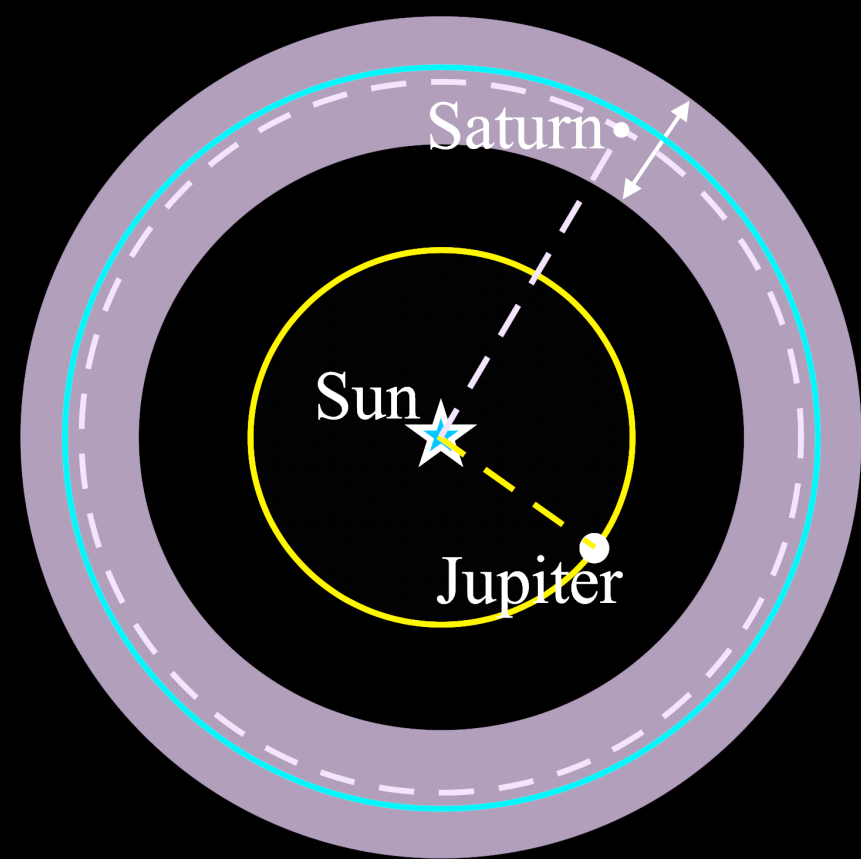
$$\frac{dZ}{dt} = XY - \beta Z$$

X

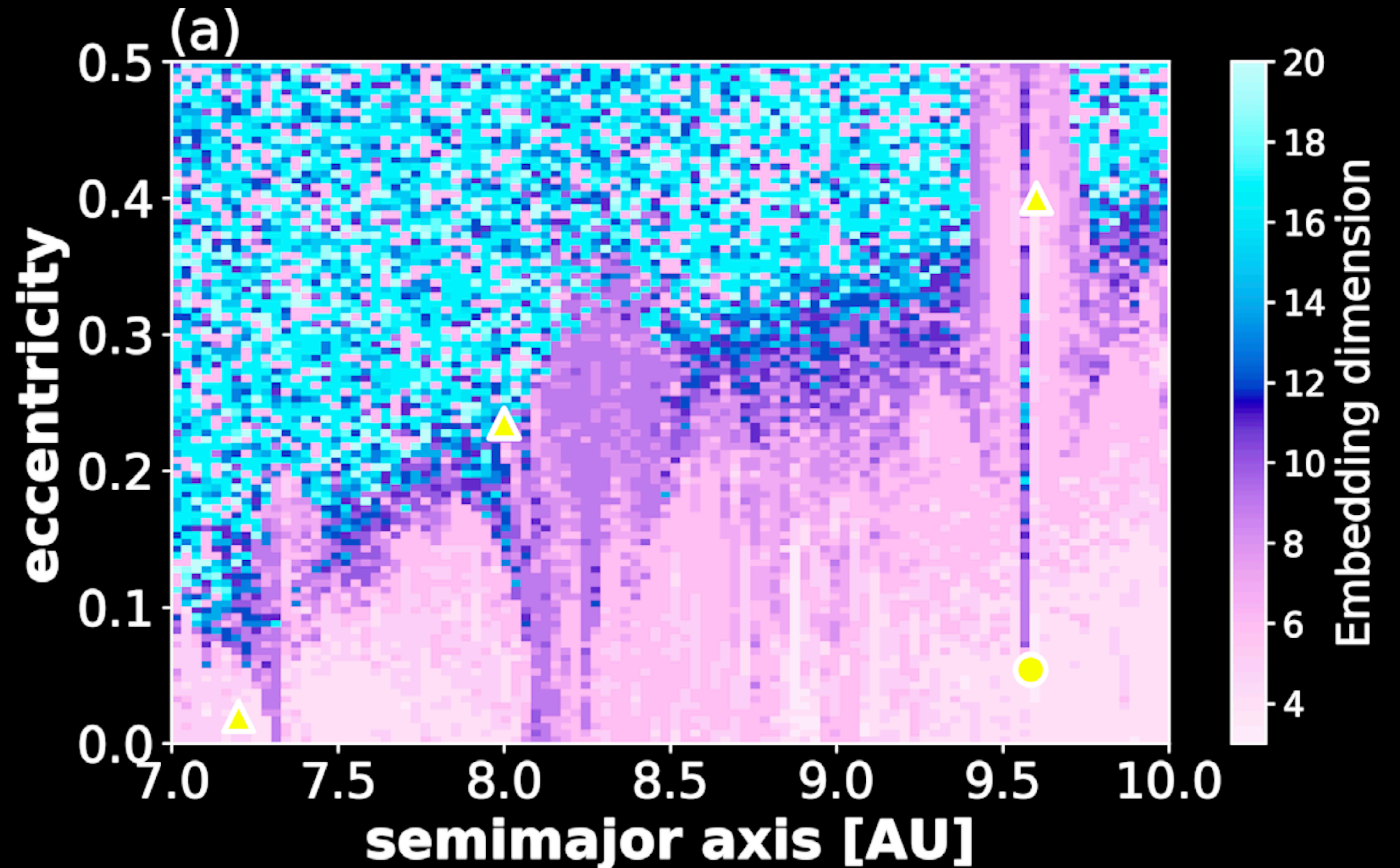
2:51

from time series to dynamics & exoplanets

phase space
reconstruction
from TTV data

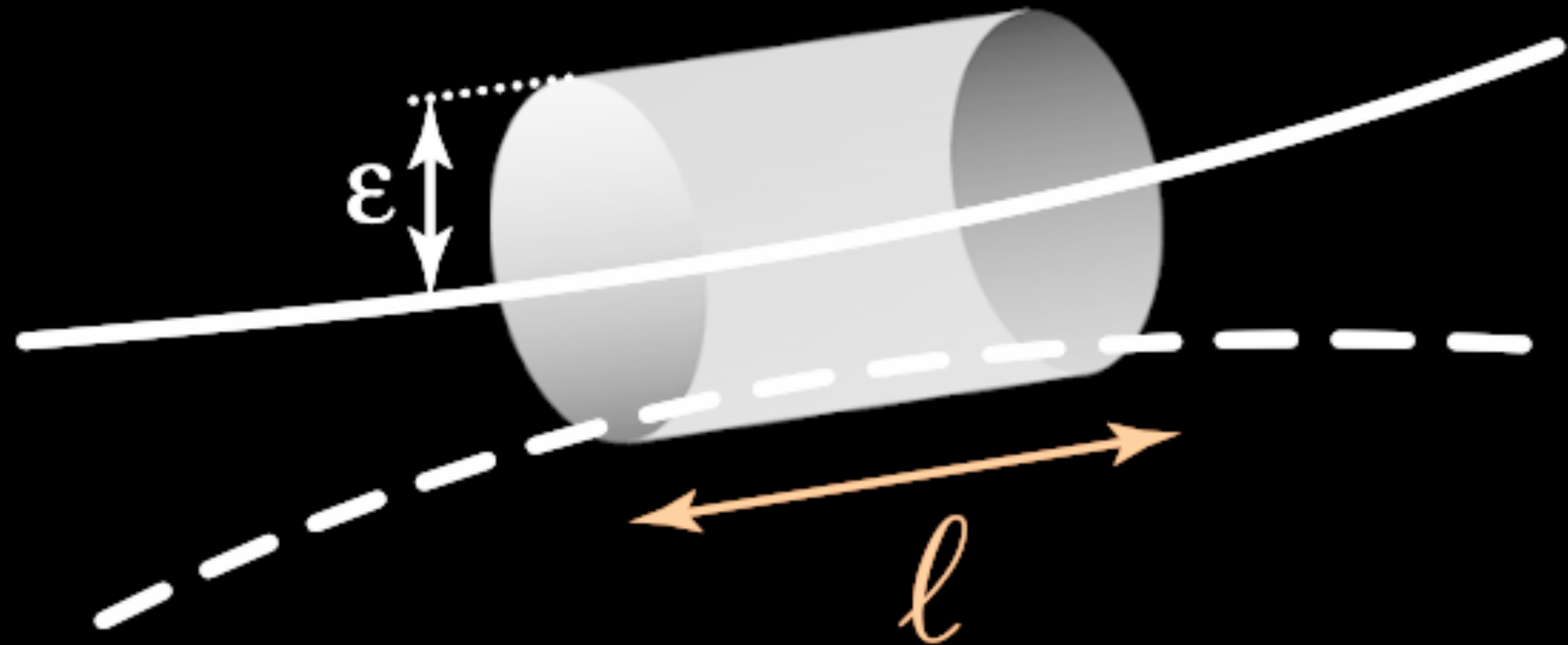


6 DoF system



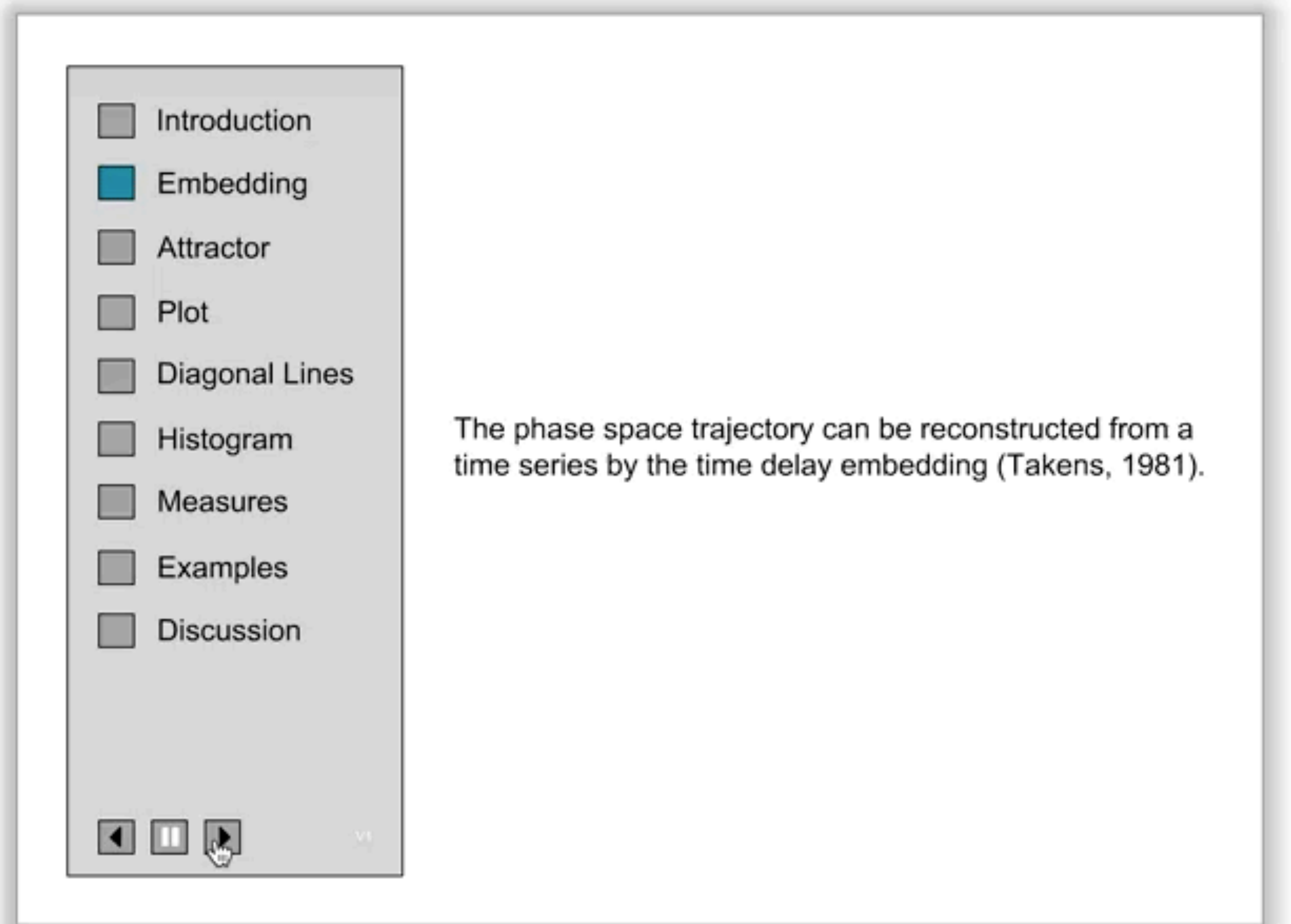
from time series to dynamics & recurrences

use the
reconstructed
trajectories
from the
perspective of
the Poincaré
recurrence
theorem



from time series to dynamics & recurrences

use the
reconstructed
trajectories
from the
perspective of
the Poincaré
recurrence
theorem



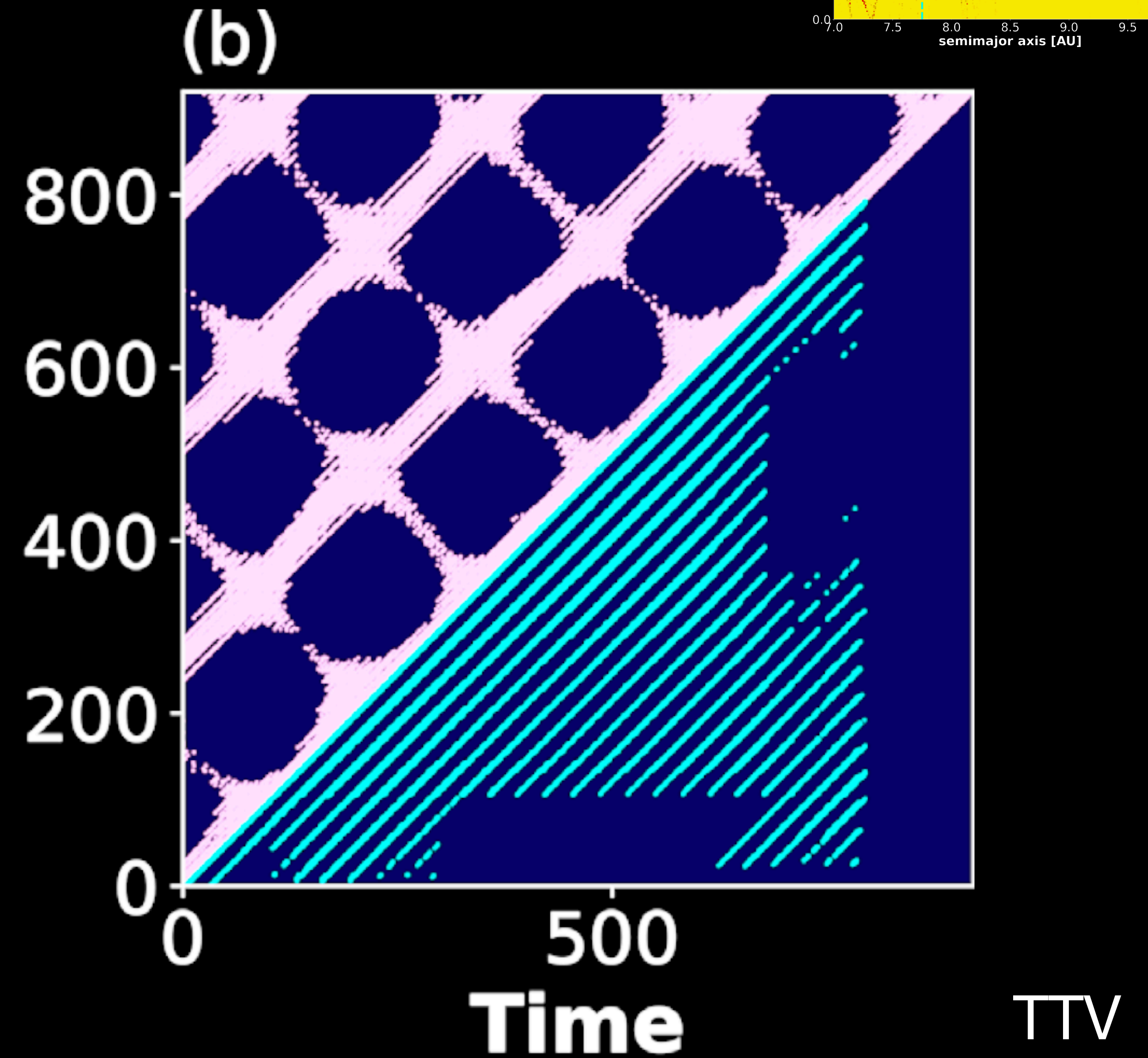
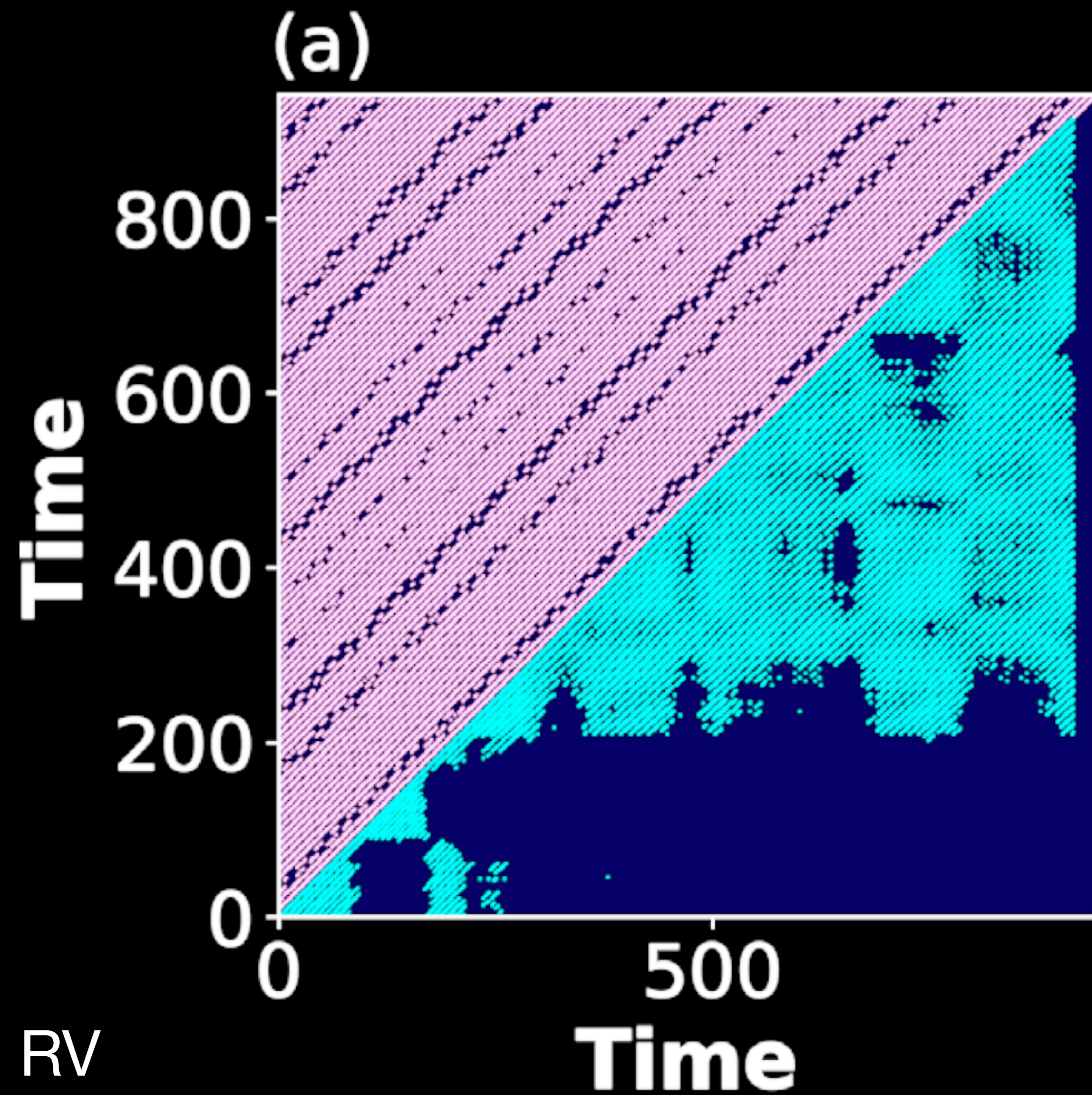
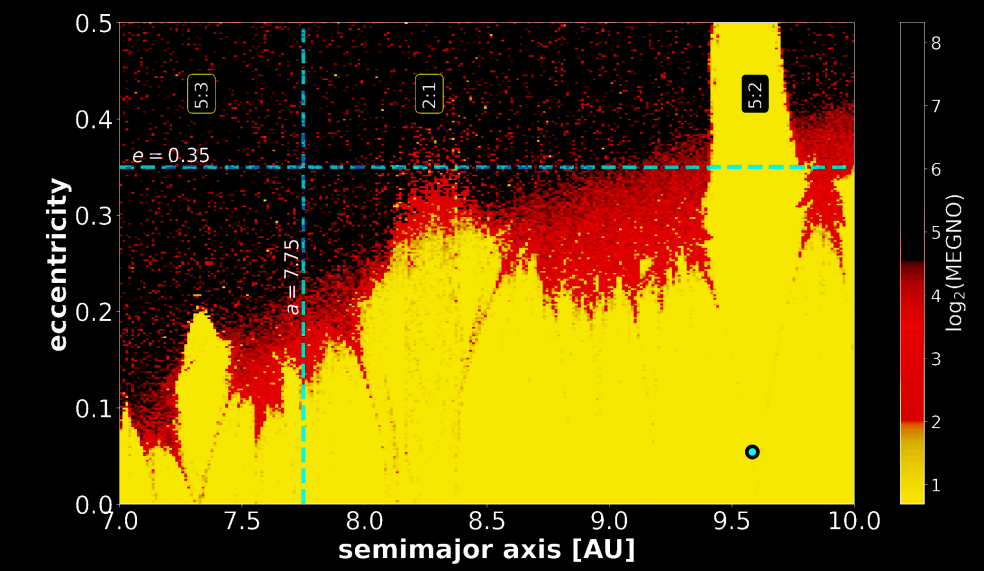
Introduction
 Embedding
 Attractor
 Plot
 Diagonal Lines
 Histogram
 Measures
 Examples
 Discussion

◀ || ▶ v1

The phase space trajectory can be reconstructed from a time series by the time delay embedding (Takens, 1981).

Brought to you by AS-Internetdienst and Norbert Marwan

recurrences & exoplanets



complex networks in brief

a **graph** is a structure amounting to a set of objects in which some of the pairs are in some sense “related”

node/vertices & connections/edges

degree list., clustering, transitivity, ave. path length

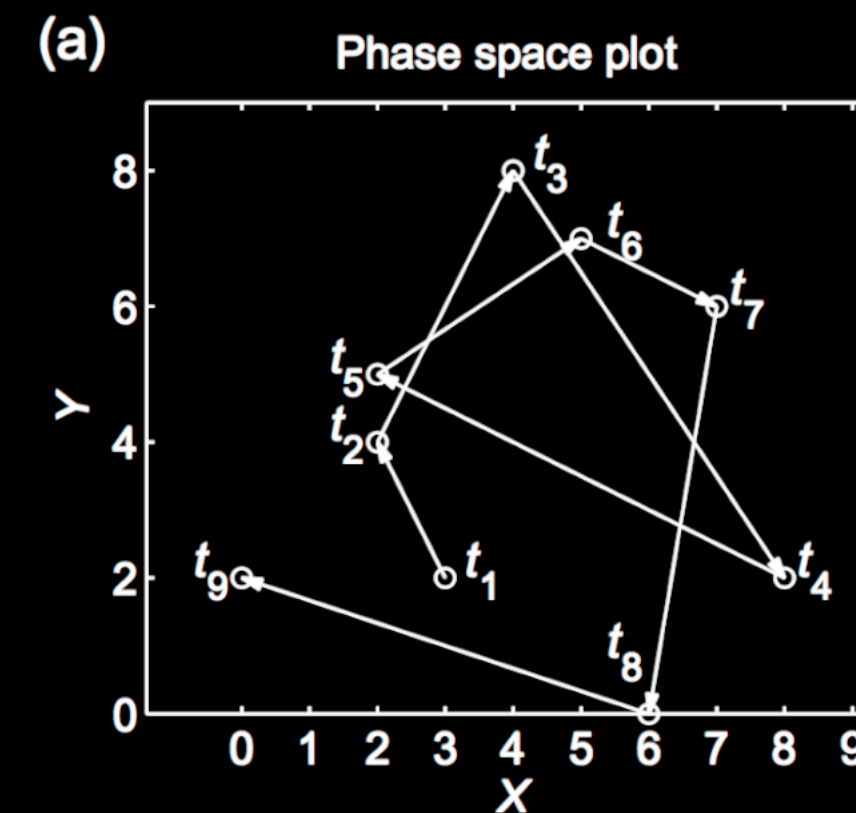
$\mathbf{A}_{17:27,15:25} =$

$$\begin{pmatrix} 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 \\ 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 \\ 1 & 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 \end{pmatrix}$$

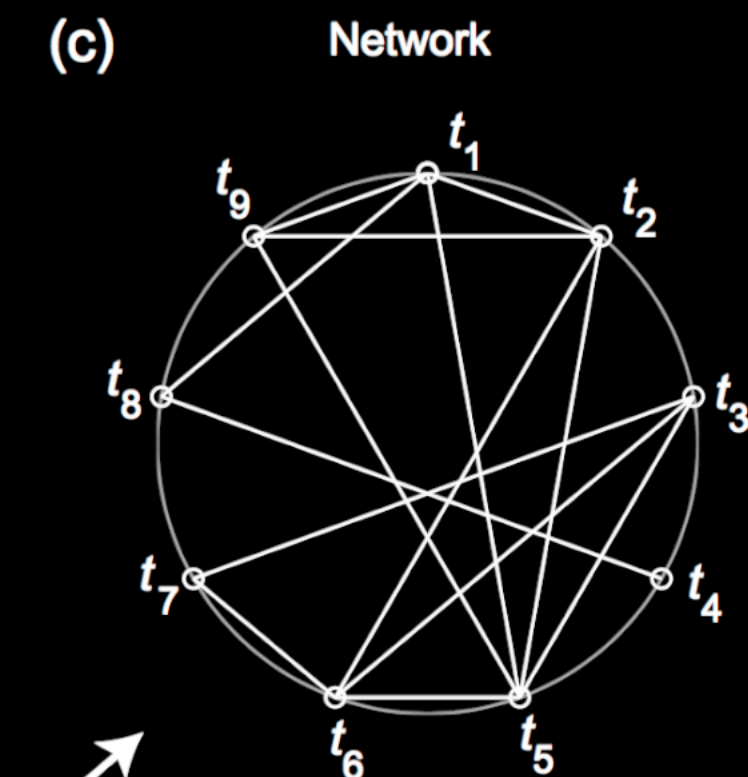
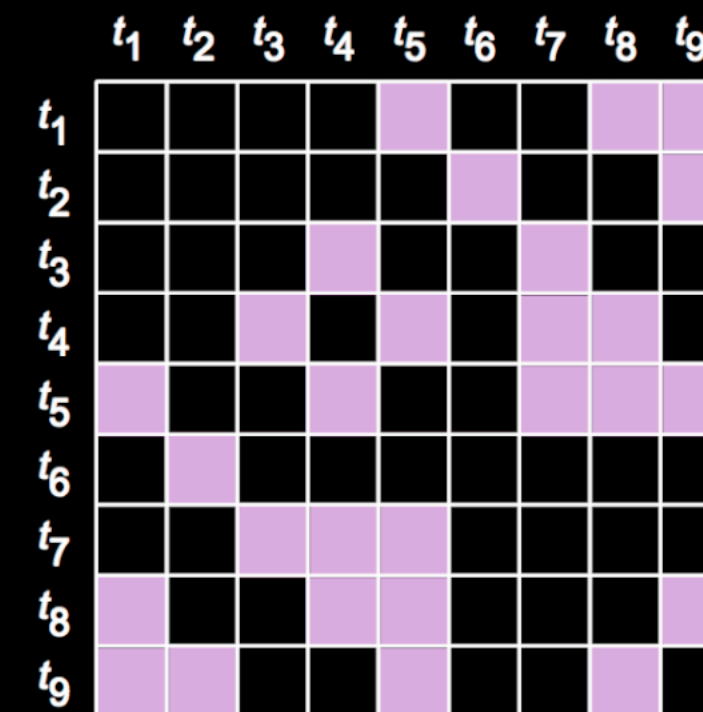
recurrence plots & networks

$$\mathbf{R}_{i,j}(\epsilon) = \Theta(\epsilon - \|\mathbf{x}_i - \mathbf{x}_j\|)$$

$$\mathbf{A}_{i,j}(\epsilon) = \mathbf{R}_{i,j}(\epsilon) - \delta_{ij}$$



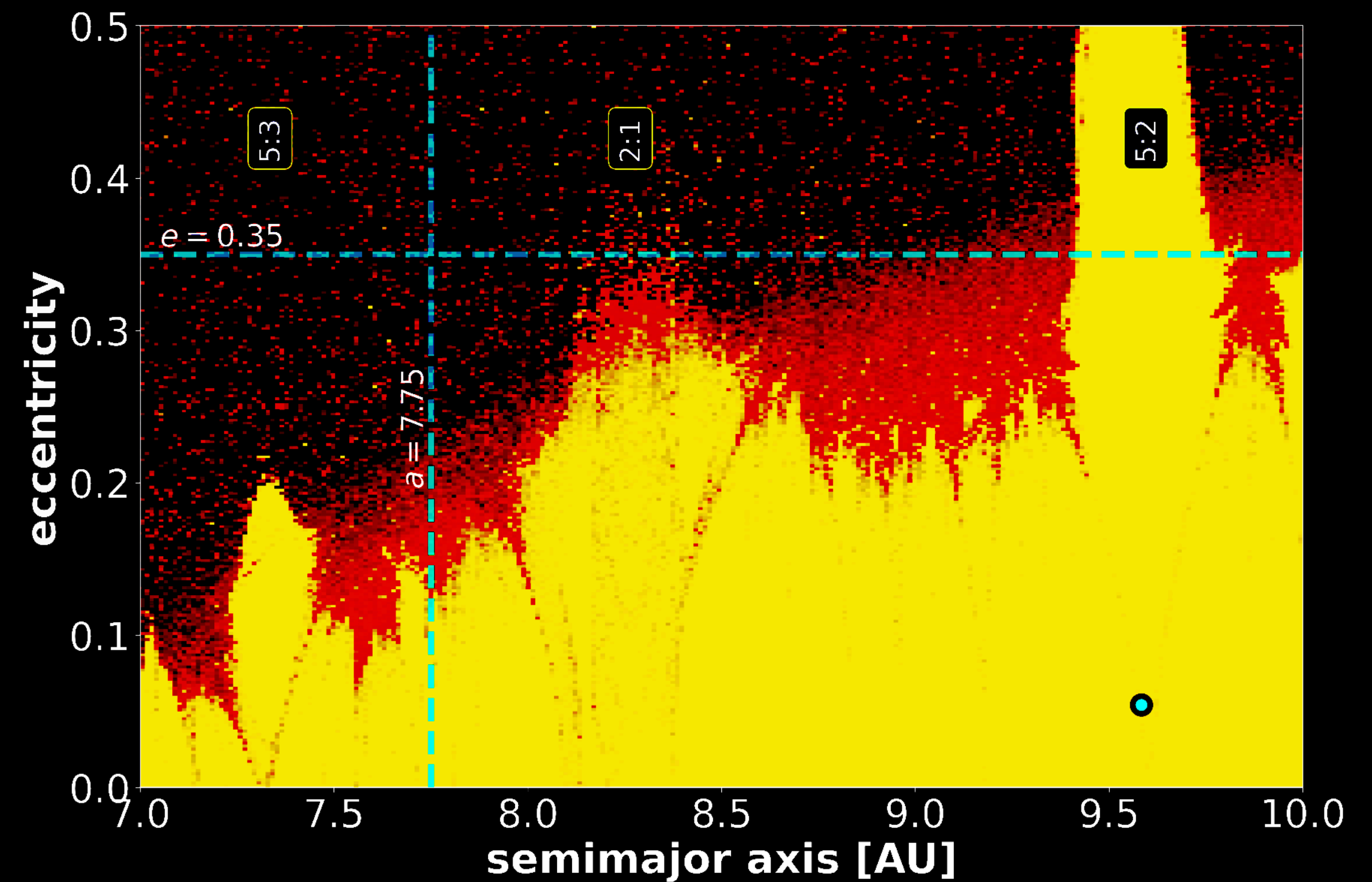
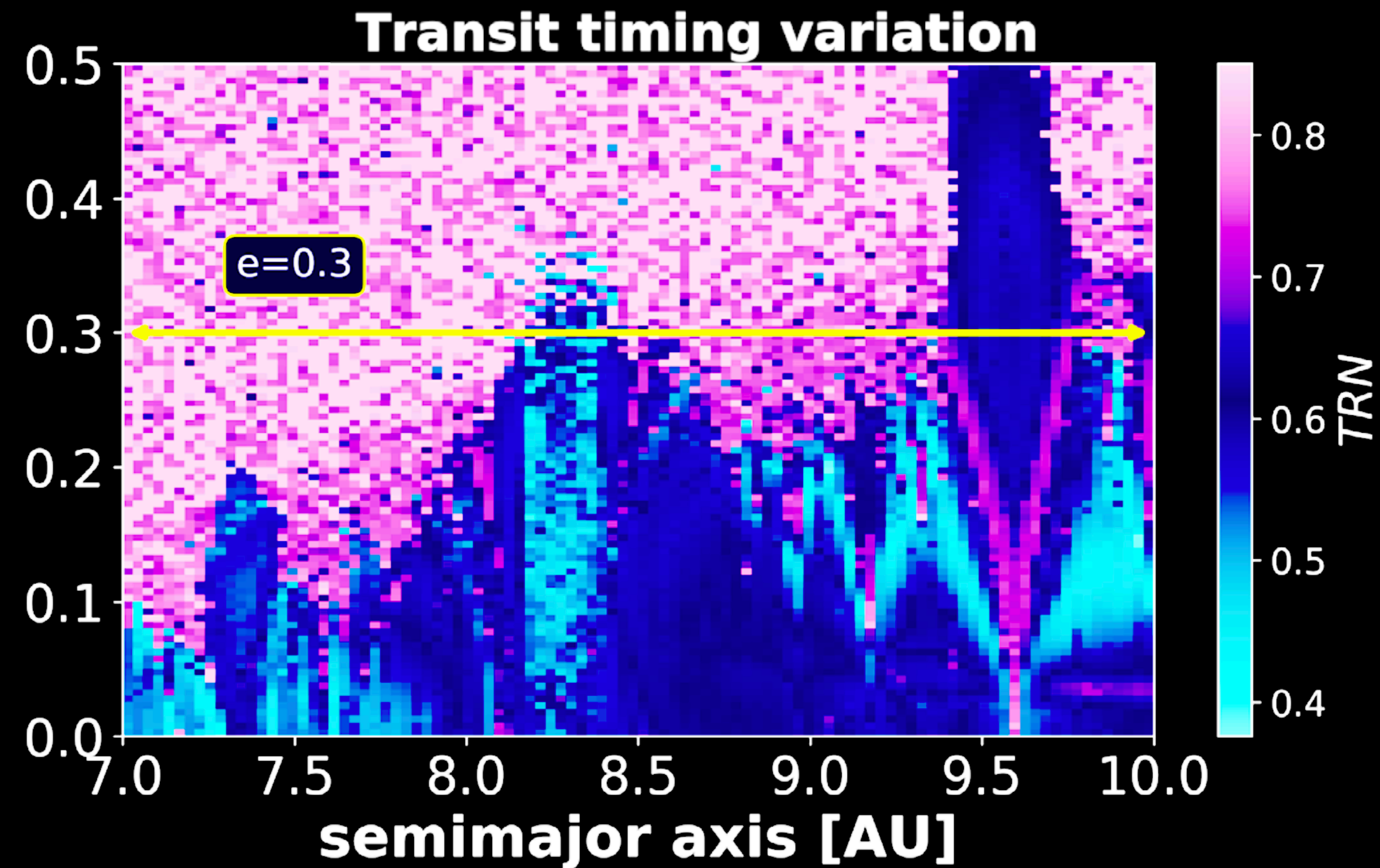
(b) RP and $A_{ij}(\epsilon=3, \text{maximum norm})$



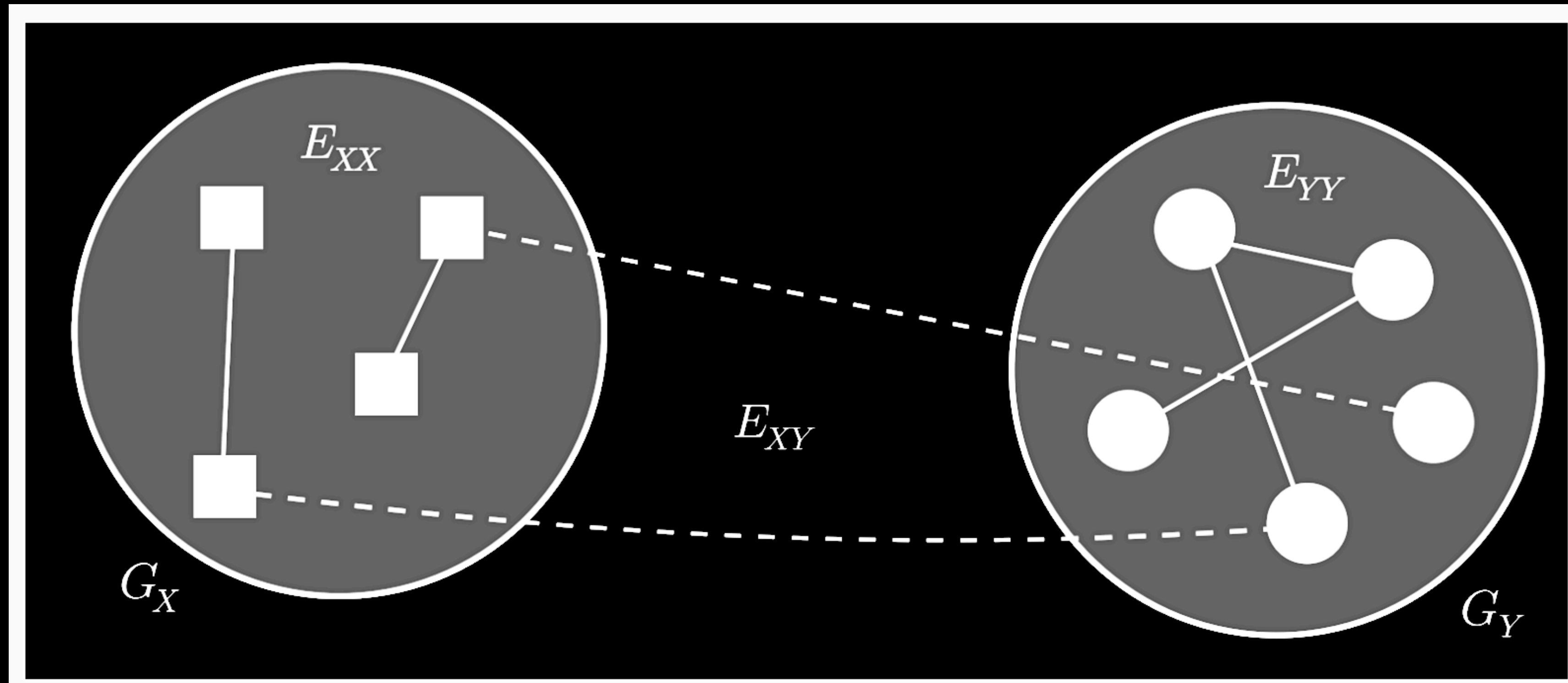
(d) Shortest path length list: l_{ij}

	t_1	t_2	t_3	t_4	t_5	t_6	t_7	t_8	t_9
t_1		1	2	2	1	2	3	1	1
t_2	1		2	3	1	1	2	2	1
t_3	2	2		4	1	1	1	3	2
t_4	2	3	4		3	4	5	1	3
t_5	1	1	1	3		1	2	2	1
t_6	2	1	1	4	1		1	3	2
t_7	3	2	1	5	2	1		4	3
t_8	1	2	3	1	2	3	4		2
t_9	1	1	2	3	1	2	3	2	

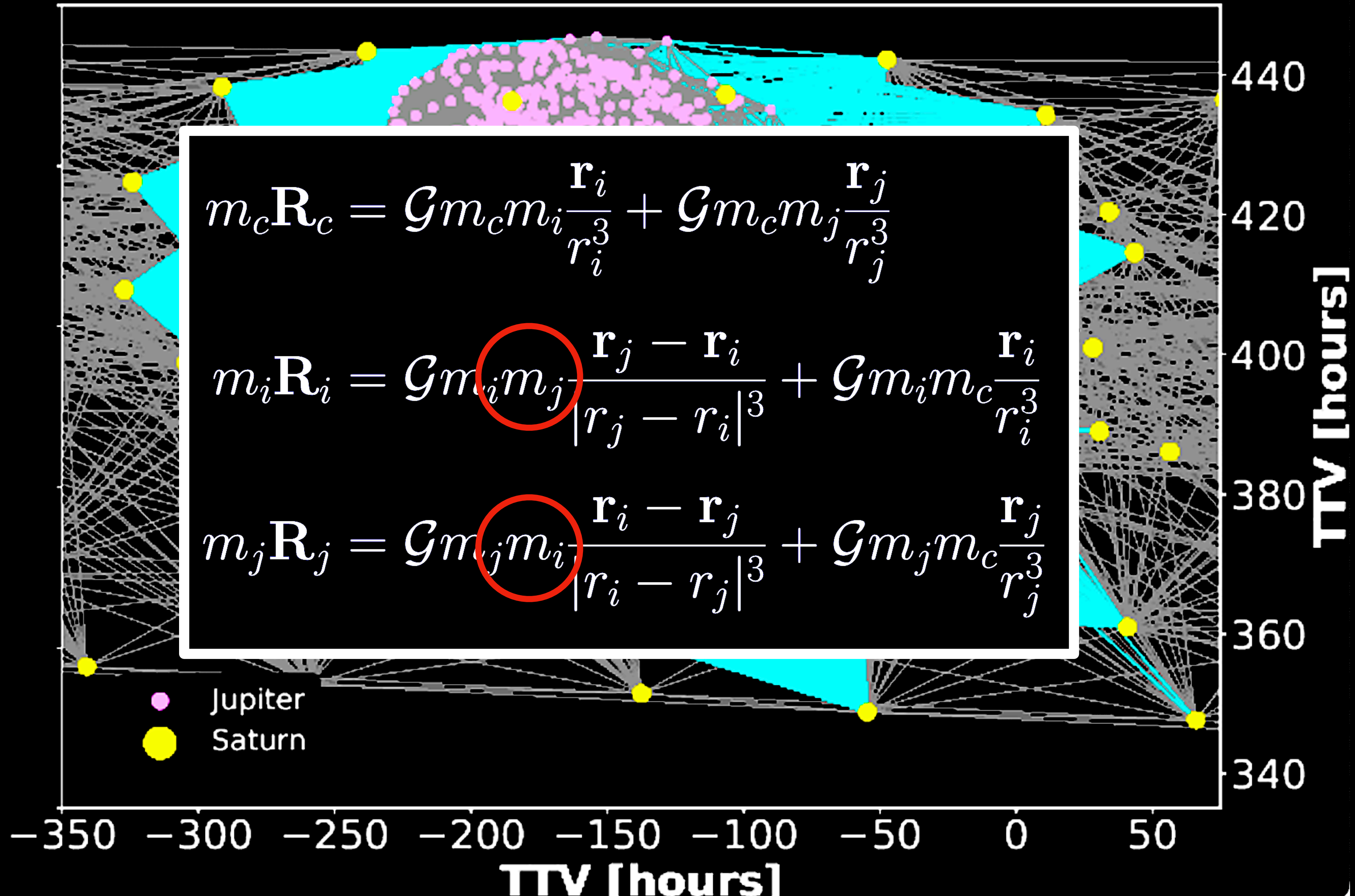
recurrence networks & exoplanet stability



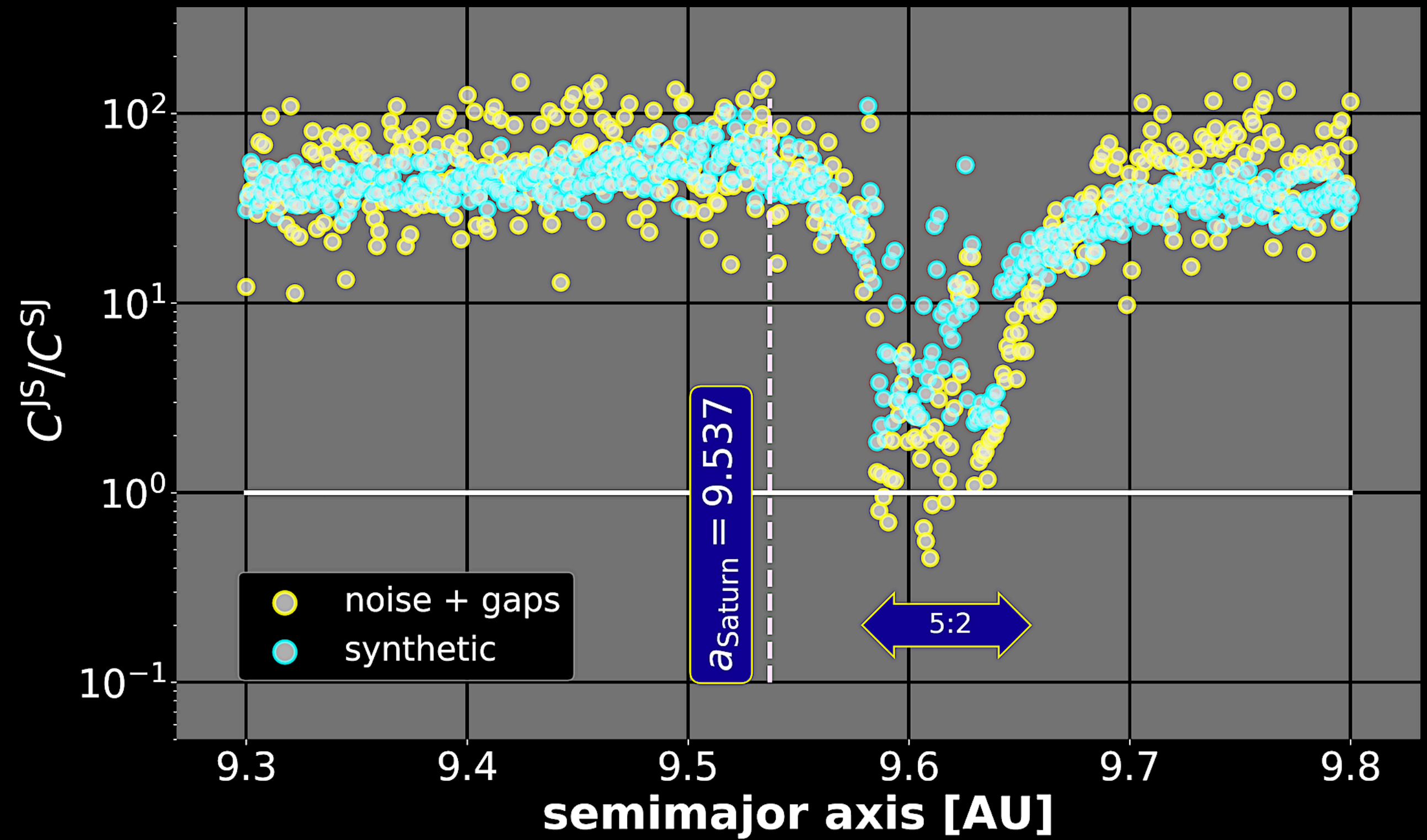
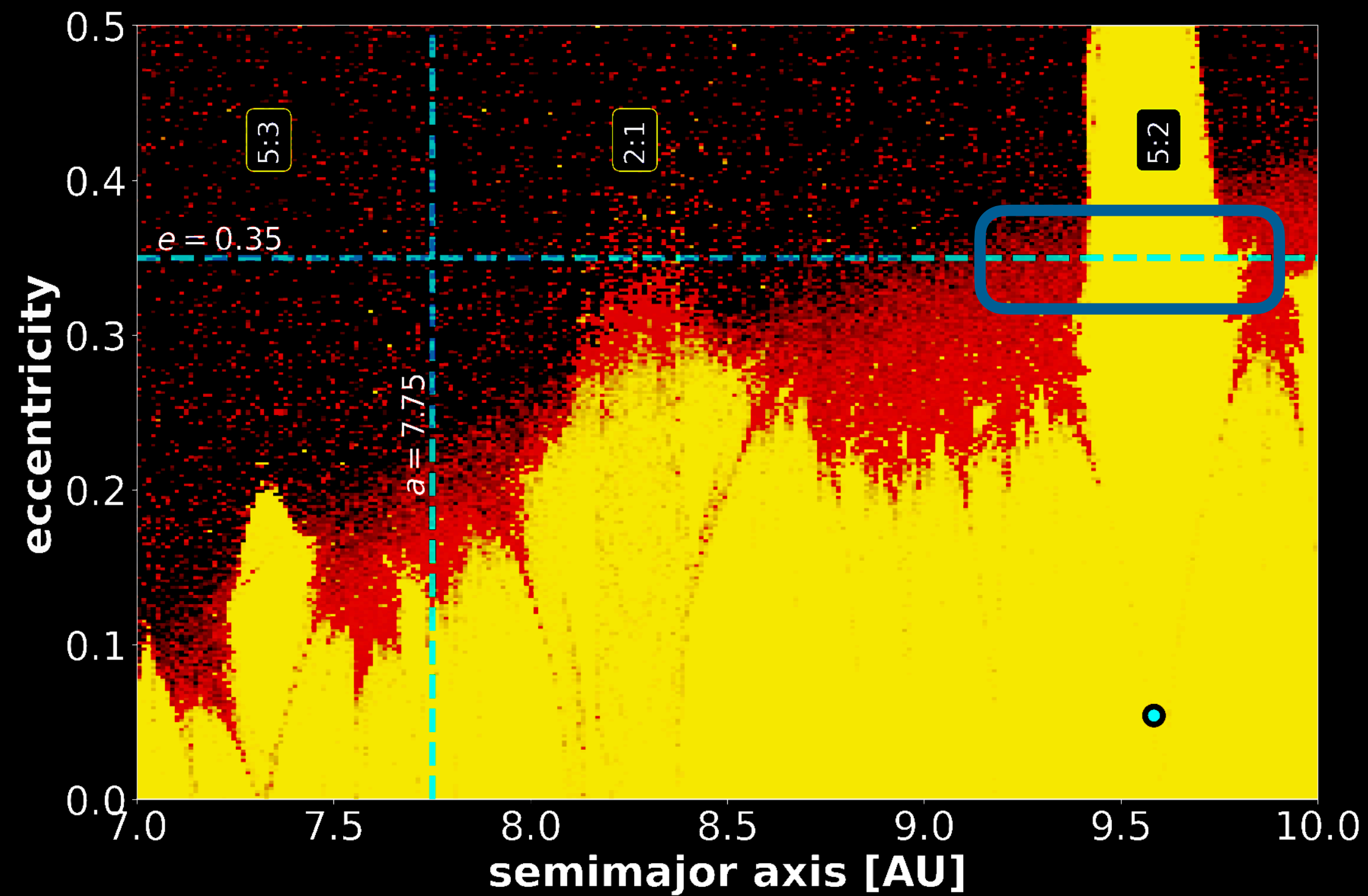
inter-system recurrence networks (isn)



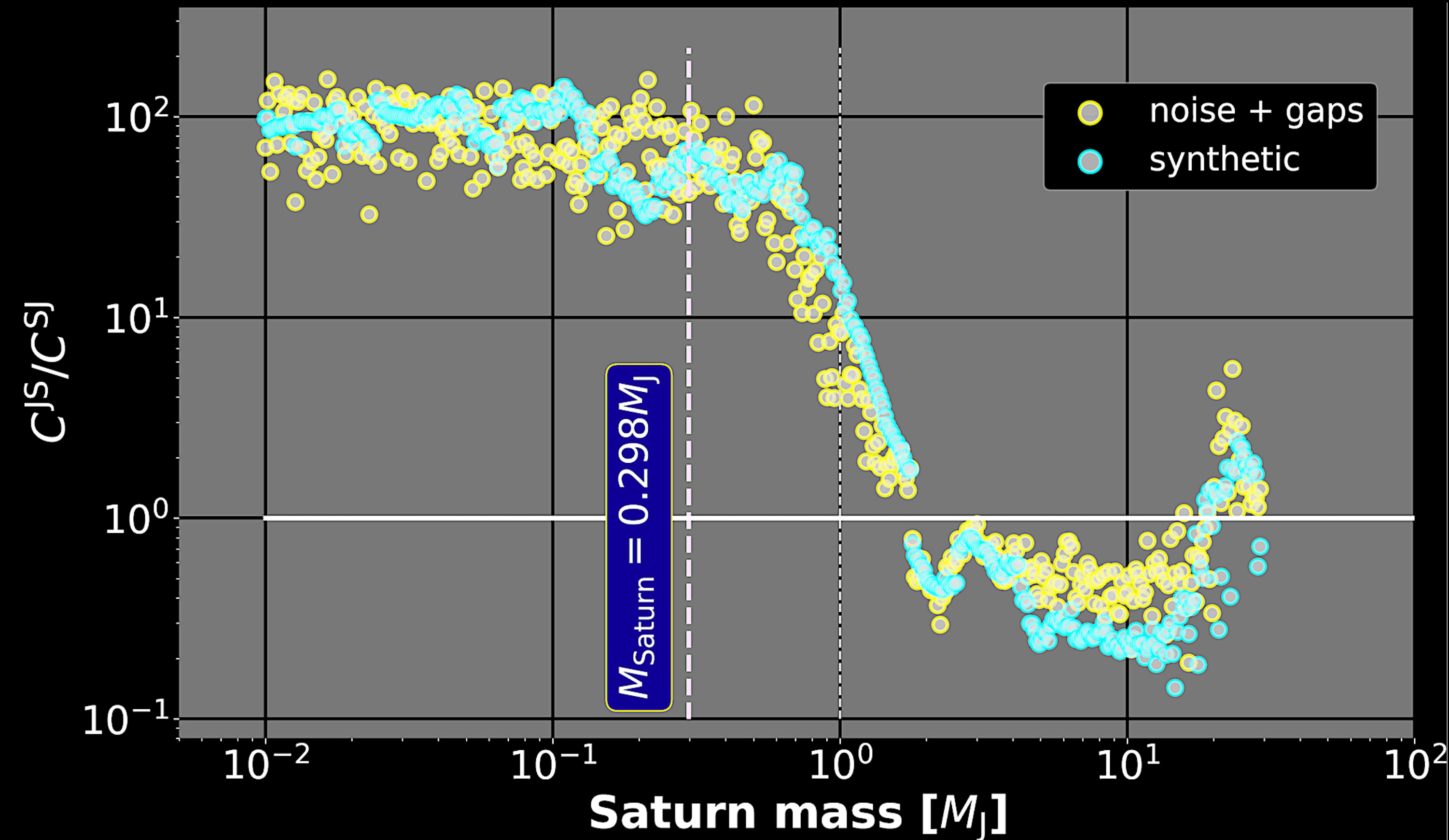
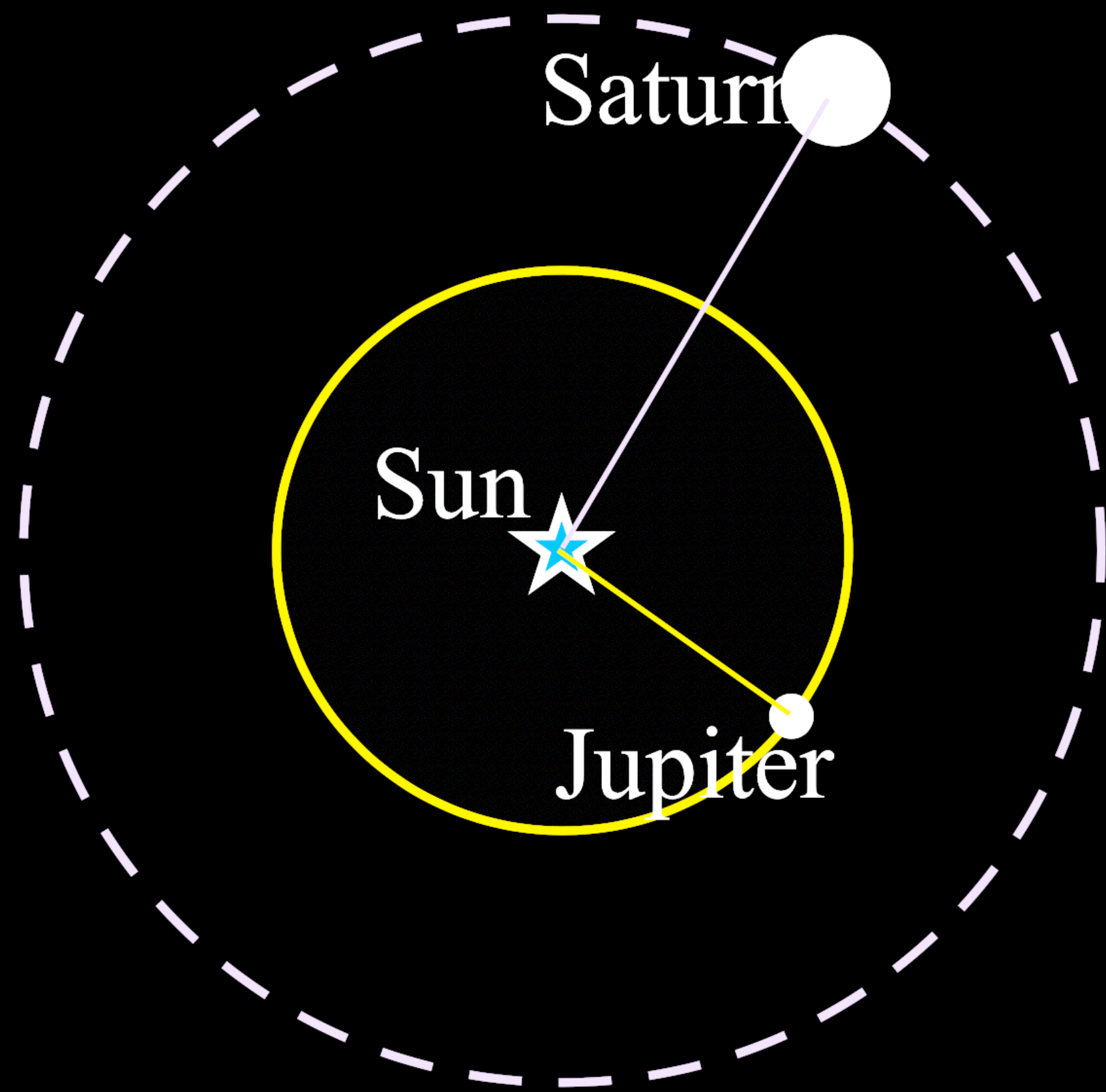
isrn for Jupiter & Saturn



isrn for Jupiter & Saturn



isrn for Jupiter & Saturn



Mass estimates of exoplanets based on the topology of interacting networks

Bolyai Seminar, 2024 April

Tamás Kovács, Eötvös University, Hungary

inter-system recurrence networks

Coupling direction	Expected relation in network measures
no coupling	$\mathcal{T}^{XY} \approx \mathcal{T}^{YX}, \mathcal{C}^{XY} \approx \mathcal{C}^{YX}$
$X \rightarrow Y$	$\mathcal{T}^{XY} < \mathcal{T}^{YX}, \mathcal{C}^{XY} < \mathcal{C}^{YX}$
$Y \rightarrow X$	$\mathcal{T}^{XY} > \mathcal{T}^{YX}, \mathcal{C}^{XY} > \mathcal{C}^{YX}$
$X \leftrightarrow Y$	$\mathcal{T}^{XY} \approx \mathcal{T}^{YX}, \mathcal{C}^{XY} \approx \mathcal{C}^{YX}$


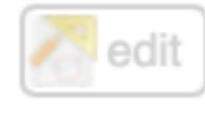
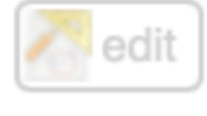

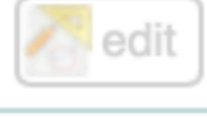



$$\mathbf{A}(\varepsilon) = \mathbf{IR}(\varepsilon) - \mathbb{I}_N$$

$$\begin{aligned} CR_{ij}(\varepsilon^{XY}) &= CR^{XY}(x_i, y_j | \varepsilon^{XY}) \\ &= \Theta(\varepsilon^{XY} - d^{XY}(x_i, y_j)) \end{aligned}$$

$$\mathbf{IR}(\varepsilon) = \begin{pmatrix} \mathbf{R}^X(\varepsilon) & \mathbf{CR}^{XY}(\varepsilon) \\ \mathbf{CR}^{YX}(\varepsilon) & \mathbf{R}^Y(\varepsilon) \end{pmatrix}$$

isrn for real exoplanets

$$\mathcal{T}^{bc} / \mathcal{T}^{cb} \approx 2.28$$

	Kepler-29 b	Kepler-29 c
Alternative planet names	KOI-738.01, KOI-738 b, KIC 10358759 b	KOI-738.02, KOI-738 c, KIC 10358759 c
Lists	Confirmed planets	
Mass [M_{jup}]	0.014 ^{+0.004} _{-0.005} 	0.013±0.004 
Mass [M_{earth}]	4.5 ^{+1.4} _{-1.5}	4.0 ^{+1.2} _{-1.3}
Radius [R_{jup}]	0.328 	0.264 
Radius [R_{earth}]	3.68	2.96
Orbital period [days]	10.33 	13.29 
Semi-major axis [AU]	0.0900 	0.1100 
Eccentricity	N/A	N/A
Equilibrium temperature [K]	N/A	N/A
Discovery method	transit	
Discovery year	2012	