

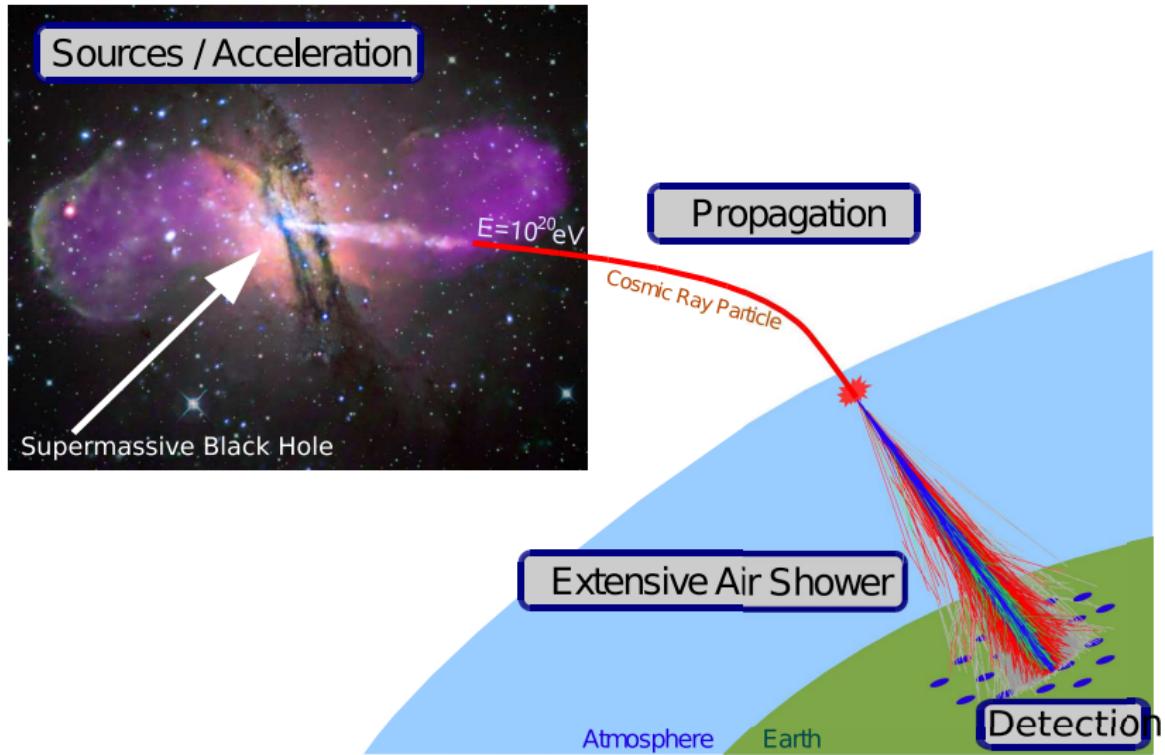
Importance of dedicated LHC measurements for cosmic ray physics

Ralf Ulrich

Karlsruhe Institute of Technology

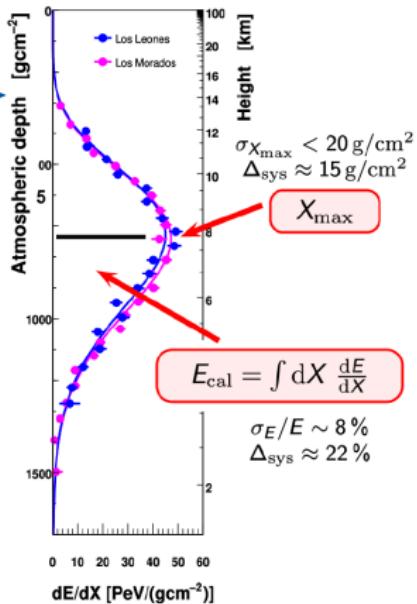
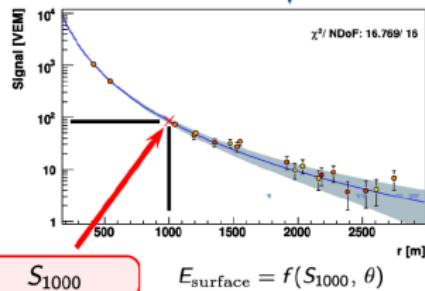
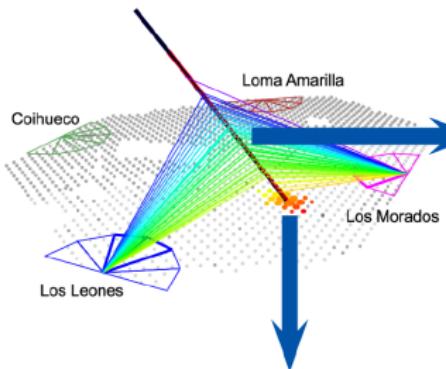
30. October 2019, Budapest

Cosmic rays at Earth



Messengers from violent places in our universe. What are they?
Where do they come from? How do they interact?

Introduction of observables



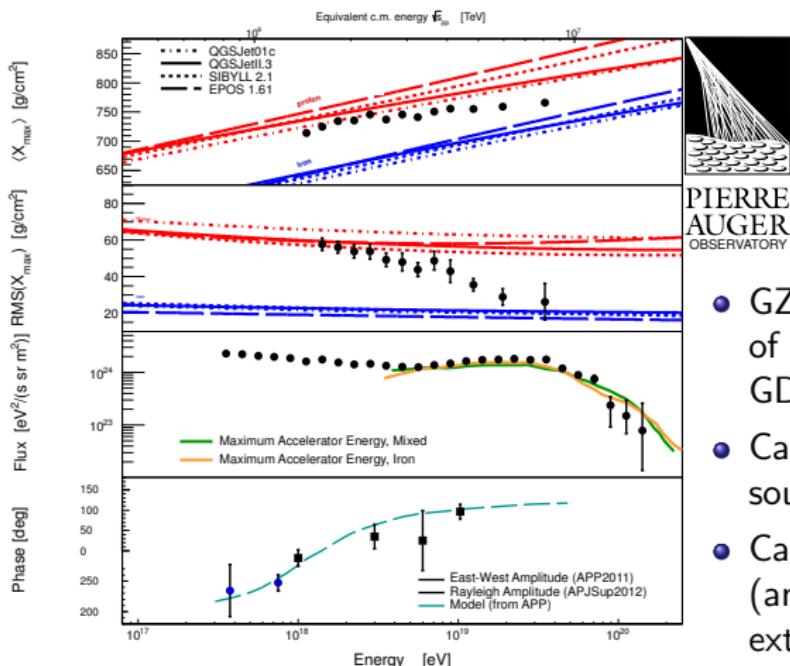
The atmosphere as calorimeter

Telescopes measure dE/dX and timing.

Surface detectors measure particle fluxes and timing.

$$N_\mu \propto \left(\frac{E}{A}\right)^\beta$$
$$X_{\max} \propto \ln\left(\frac{E}{A}\right)$$
$$E_0 = E_{\text{cal}} + E_{\text{invisible}}$$
$$E_0 \propto S_{1000}$$

What is the nature and the sources of the highest energy particles in the universe?



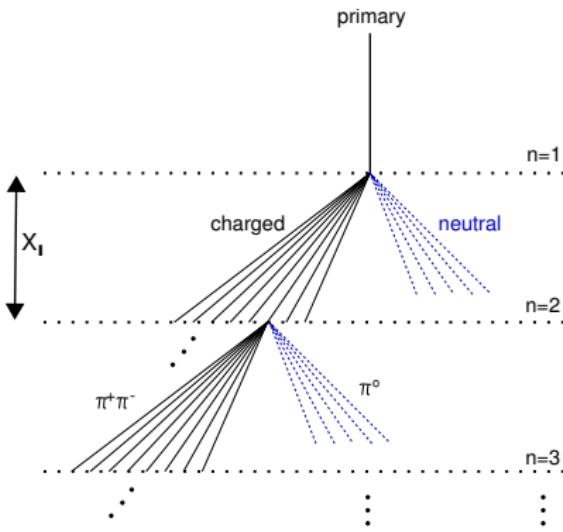
PIERRE
AUGER
OBSERVATORY

- GZK cutoff? Maximum energy of accelerators? GDR/dissociation?
- Can we find individual (point) sources?
- Can we identify a transition (anisotropy) from galactic to extragalactic?

→ Mass composition main challenge to resolve questions!

Air shower cascades

Extended Heitler Model:



Shower maximum

$$X_{\max} \approx \lambda_I + X_0 \ln \frac{E_0}{N_{\text{mult}} E_{\text{crit}}^{\text{e.m.}}}$$

Muon number at observation level

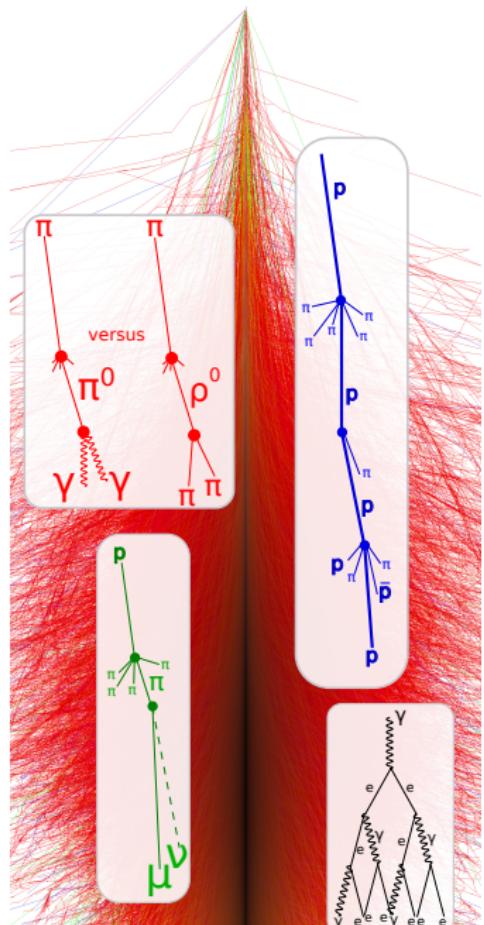
$$N_\mu = N_{\pi^\pm} = \left(\frac{E_0}{E_{\text{crit}}^I} \right)^\beta$$

where

$$\beta = \ln \left(\frac{2}{3} N_{\text{mult}} \right) / \ln (N_{\text{mult}}) \approx 0.9$$

(J. Matthews, APP 22 (2005) 387)

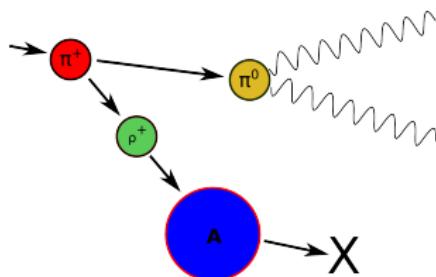
Ultra-high energy cosmic ray extensive air showers



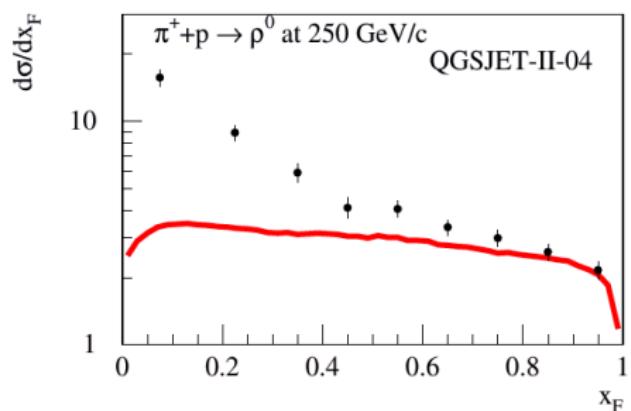
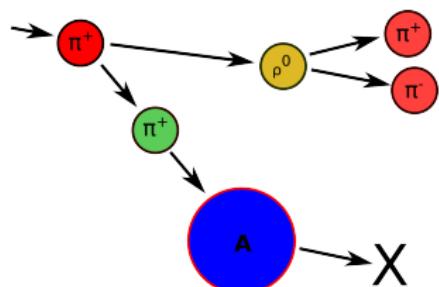
- Huge coupled cascading process
- Conversion of primary energy into
 - Electrons/ photons
 - Muons
 - Low energy hadrons
- Most relevant mass-sensitive observables are:
 - Depth of shower maximum
 - Muon content
- Precise modelling is mandatory for a cosmic ray mass measurement

Forward ρ^0 Production, QGSJetII.3→QGSJetII.4

Charge Exchange, Leading π^0/ρ^0 production:



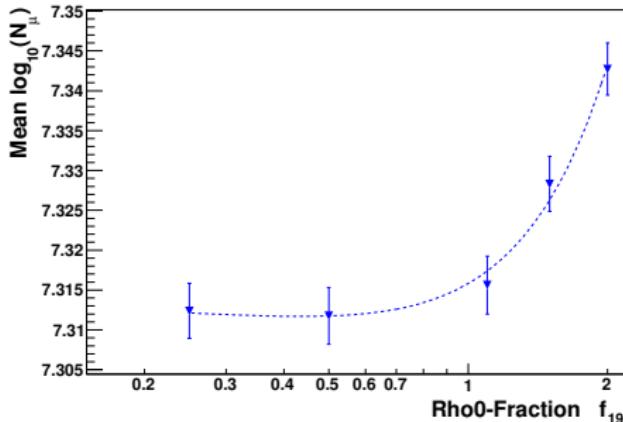
versus



S. Ostapchenko, ISVHECRI 2012

Impact on muons in air showers

Systematically change the leading π^0/ρ^0 ratio:



Ulrich, Engel, Baus, ISVHECRI 2014

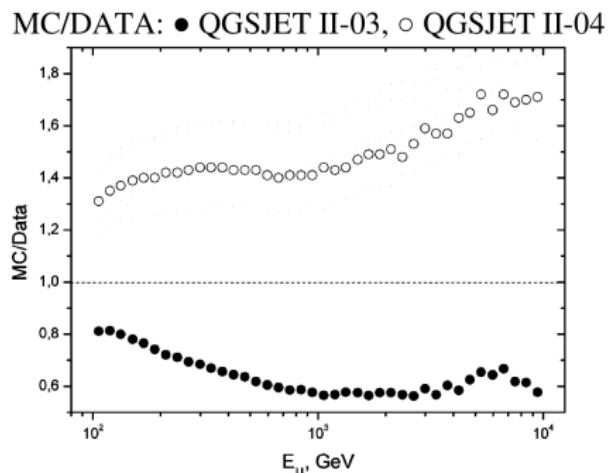
SIBYLL, protons, $10^{19.5}$ eV

Ad-hoc modifications (PRD 83 (2011) 054026): f_{19} is scaling factor for ρ^0 and baryons at 10^{19} eV:

$$\alpha^{\text{modified}}(E) = \alpha^{\text{orig}}(E) \cdot \left(1 + (f_{19} - 1) \cdot \frac{\lg_{10}(E/E_{\text{thr}})}{\lg_{10}(10^{19} \text{ eV}/E_{\text{thr}})} \right) \quad \text{with } E_{\text{thr}} = 10^{15} \text{ eV}$$

Forward ρ^0 production, QGSJetII.4

Prediction of inclusive atmospheric muon fluxes as a test of hadronic interaction models



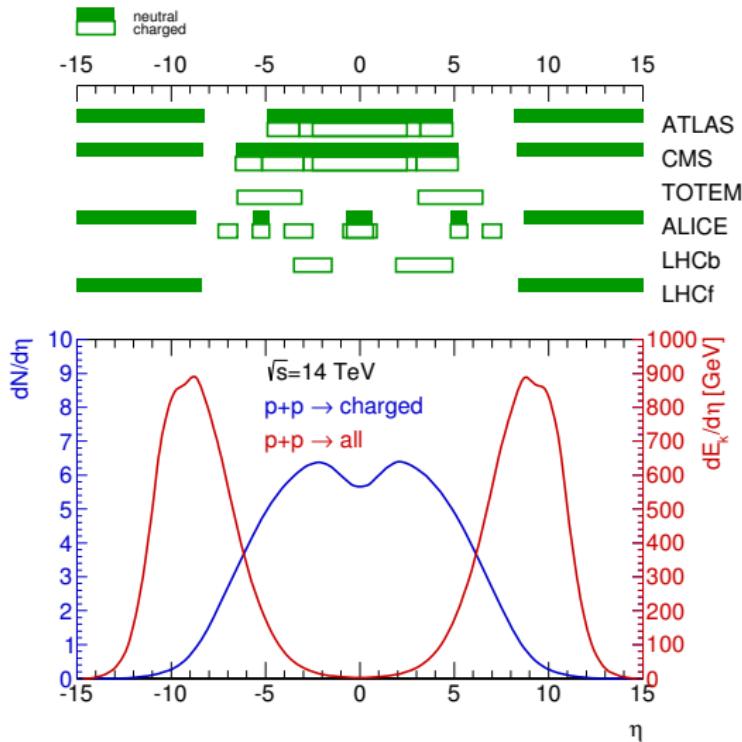
A.V. Lukyashin, ISVHECRI 2014

- ⇒ Too many muons at \sim TeV
- ⇒ But: still not enough muons at \sim GeV

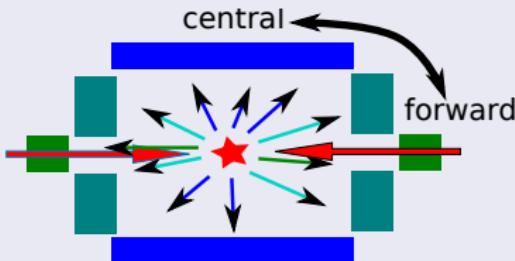
Angular acceptance of LHC experiments

Definition of *pseudorapidity*: $\eta = -\log \tan(\theta/2)$

where θ is the angle wrt. to the beam



Phase-space coverage at LHC

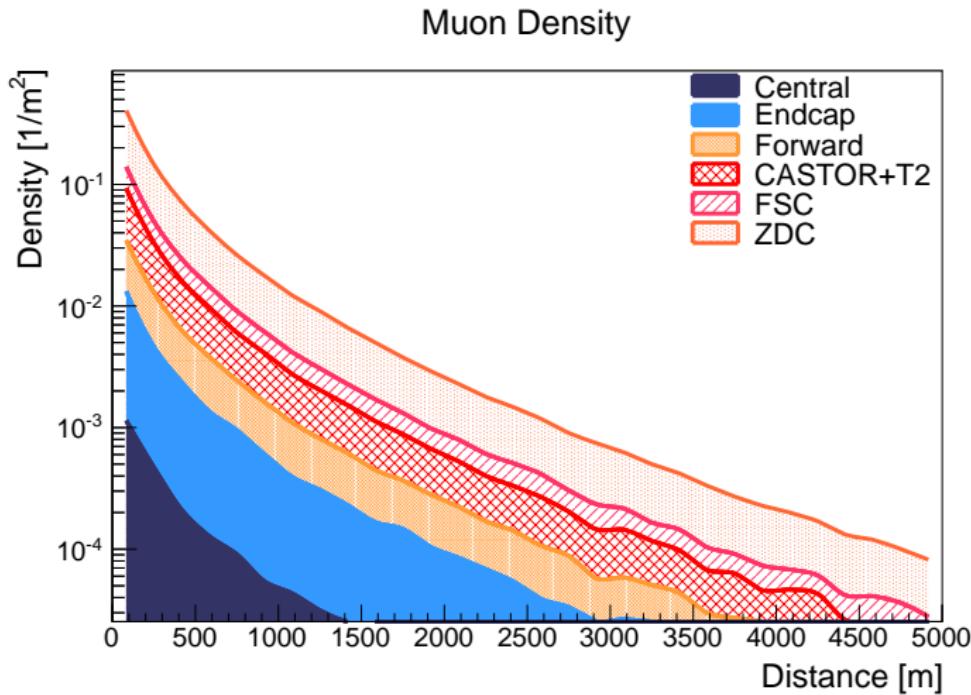


- Central ($|\eta| < 1$)
- Endcap ($1 < |\eta| < 3.5$)
- Forward ($3 < |\eta| < 5$), HF
- CASTOR+T2 ($5 < |\eta| < 6.6$)
- FSC ($6.6 < |\eta| < 8$)
- ZDC ($|\eta| > 8$), LHCf

- How relevant are specific detectors at LHC for air showers?
- Simulate parts of shower individually.

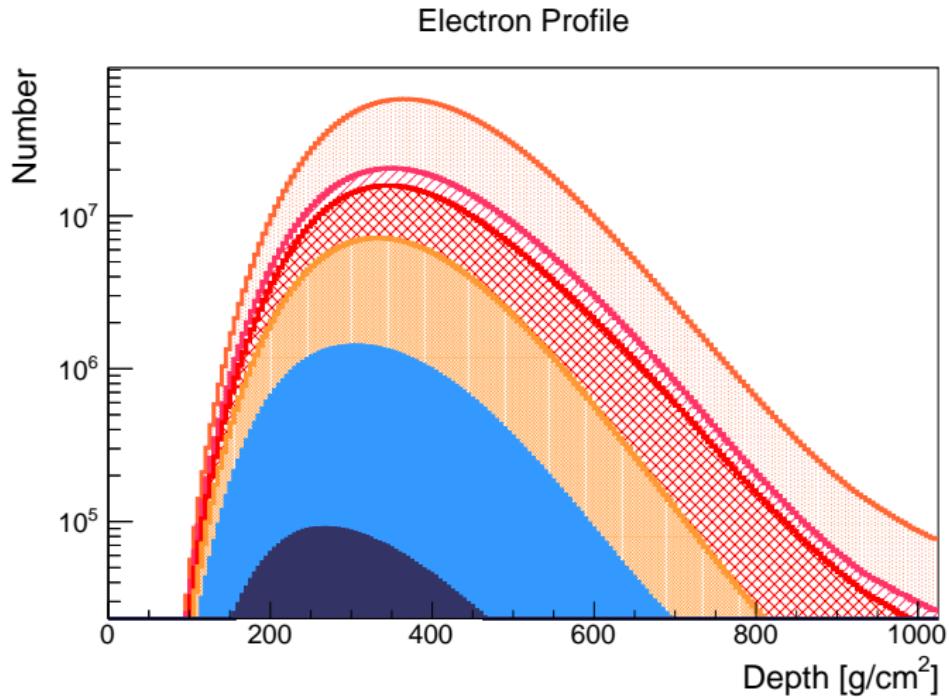


Lateral particle density on ground level



- Air shower models so far only tuned to about 10 % !
- Forward detectors are crucial.

Longitudinal shower development



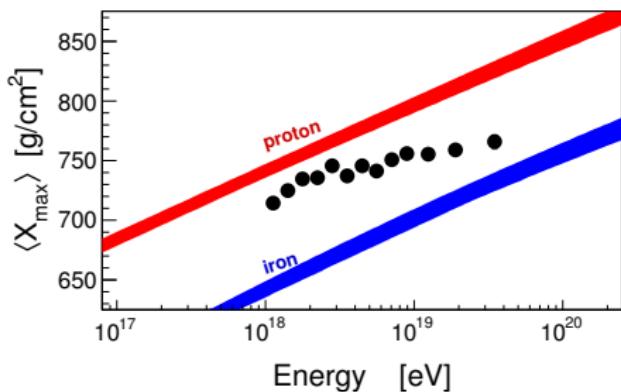
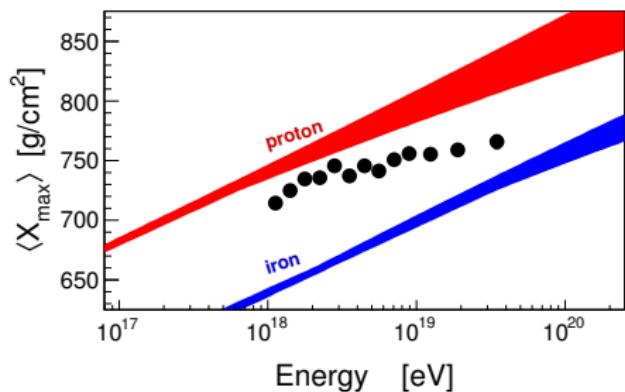
- Air shower models so far only tuned to about 10 % !
- Forward detectors are crucial.

Model tuning to LHC data and $\langle X_{\max} \rangle$ predictions

EPOS 1.99
QGSJetII.3



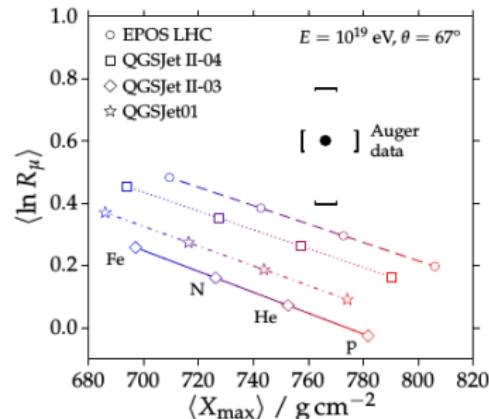
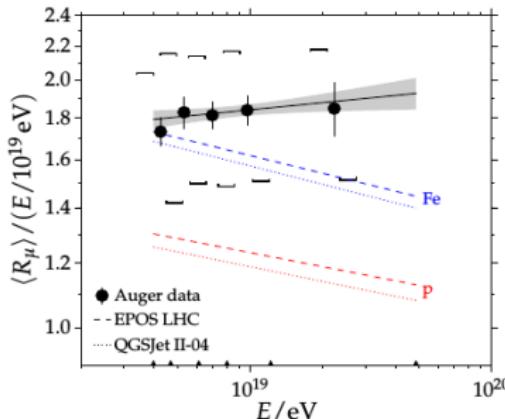
EPOS LHC
QGSJetII.4



Tuning impact:

- Obvious apparent improved model predictions
- But is this really a quantitative indication of a better understanding?

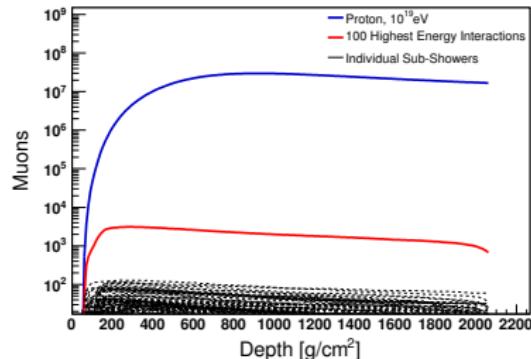
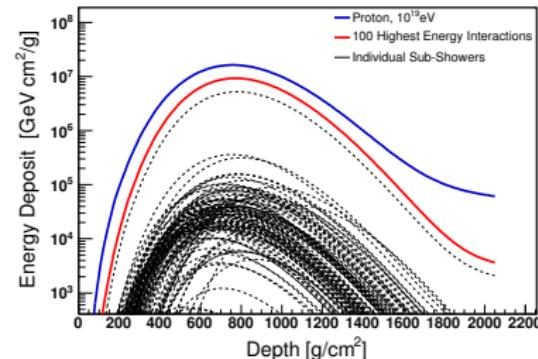
Muon content of air showers at ground level



Auger, PRD 91 (2015) 032003

- More muons in air shower data than expected
- No consistency between different observables can be achieved
- Interaction physics in air shower models still not accurate

Sensitivity to interaction physics



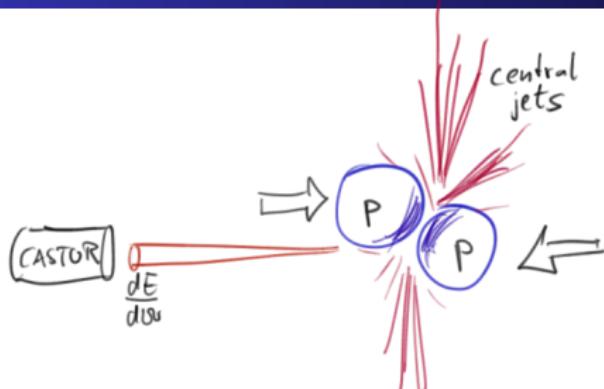
⇒ Very different impact on different EAS observables:

- Global shower properties and the shower maximum X_{\max} are sensitive to the highest energy interactions
 - Muons in air showers are sensitive to the hadronic cascade over all energies
- ⇒ Large problem in predicting the overall muon number is small problem on the level of individual interactions

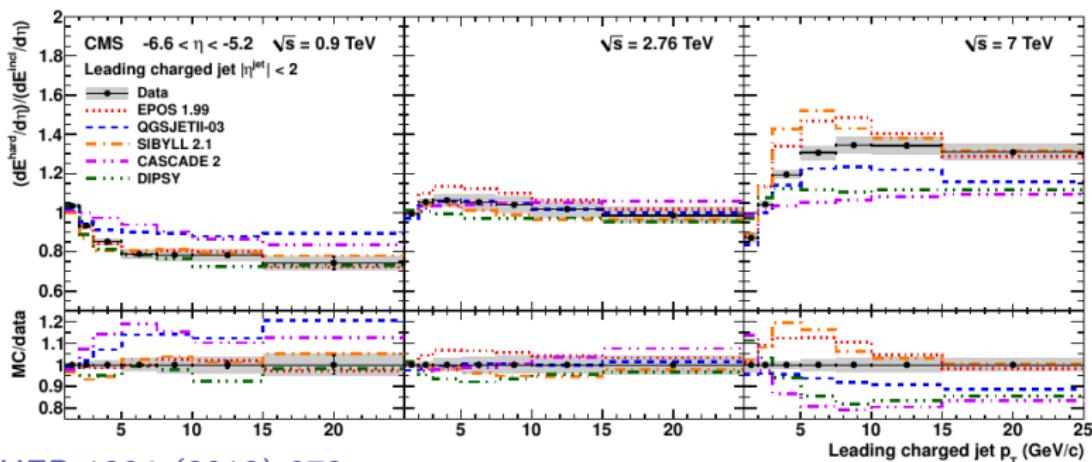
Example:

typical 8 had. generations, 5% effect in each → $1.05^8 \approx 50\%$ total effect

Forward energy as a function of central activity (pp)



- **Forward energy** \sim Remnant fragmentation
- **Central jets** \sim String fragmentation
- “Underlying-Event” study in very forward direction

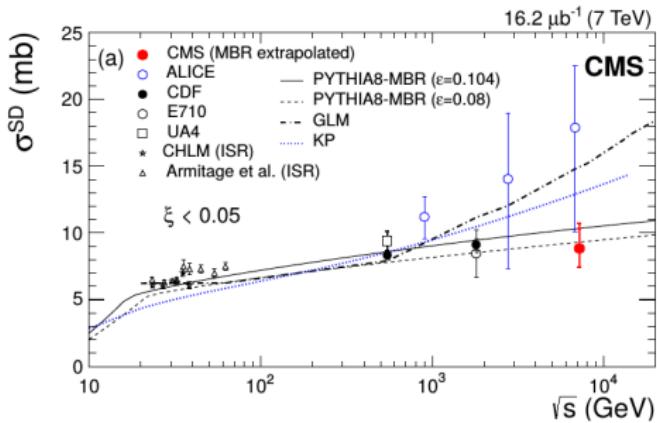
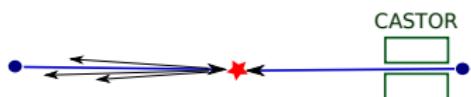


CMS/CASTOR low-mass single diffraction (pp, 7 TeV)

Double Diffraction

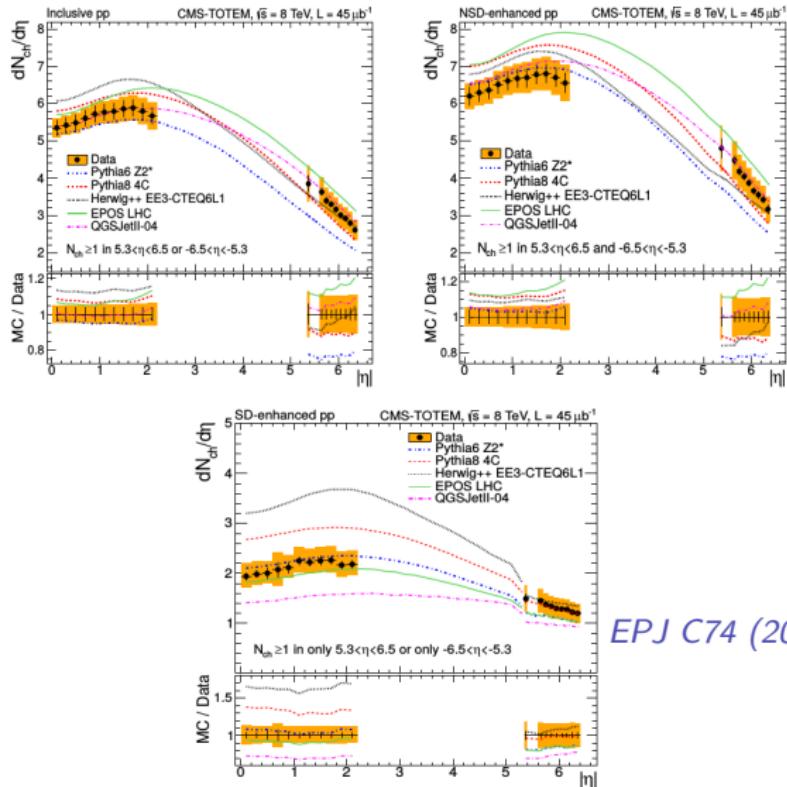


Single Diffraction



Separation of single- and double-diffraction possible with CASTOR detector

CMS + TOTEM combined multiplicity data (pp, 8 TeV)



EPJ C74 (2014) 3053

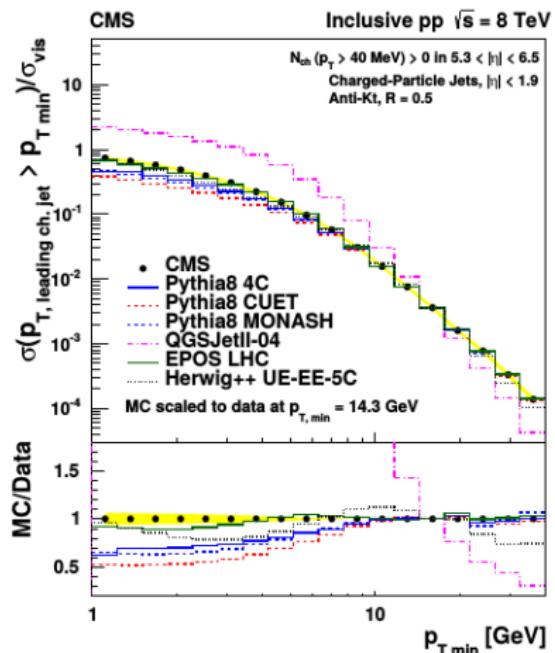
1) Very wide acceptance! 2) Correlations and final state selection!

CMS minijet measurements (pp, 8 TeV)

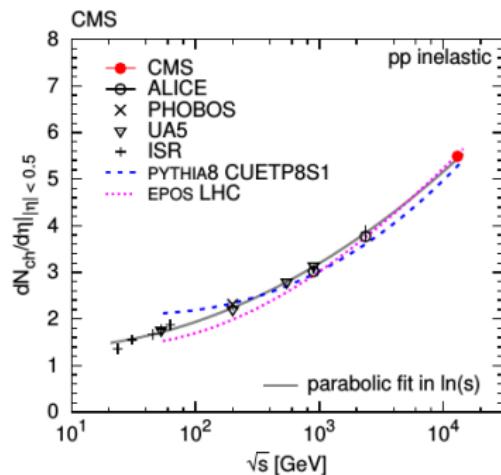
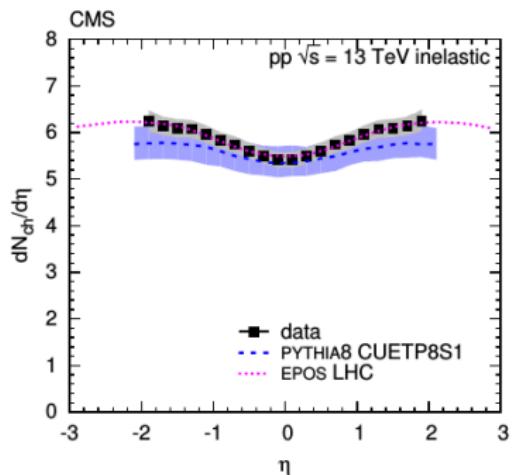
$$\zeta_{\text{QCD}}(s, p_{T,\text{min}}) = \int dP_T \int dx_1 \int dx_2 \sum_{ijkl} f_{iA}(x_1, p_T^2) f_{jB}(x_2, p_T^2) \frac{d\hat{\sigma}_{ij}^{kl}(p)}{dP_T}$$

P_T - Cut off
 Parton distribution function, PDFs
 Minijet Cross section

- Hadronization in string fragmentation, minijet production
- p_T threshold



$dN_{\text{ch, had}}/d\eta$ at 13 TeV



Phys.Lett. B 751 (2015) 143

- First LHC paper at 13 TeV (without CMS magnet \rightarrow no p_T -cutoff)
- EPOS-LHC makes an excellent first impression
- And: analysis performed in Budapest!

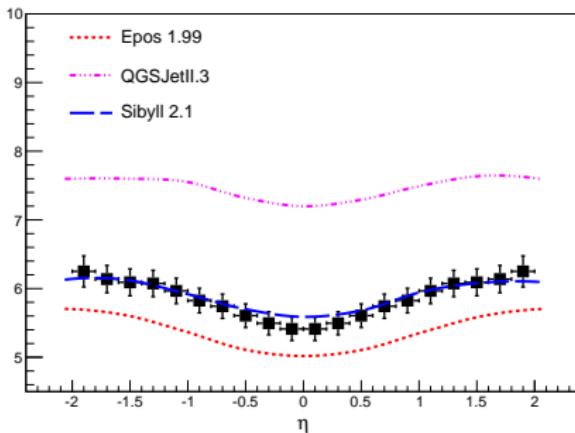
Multiplicity measurements at 13 TeV

pre-LHC models

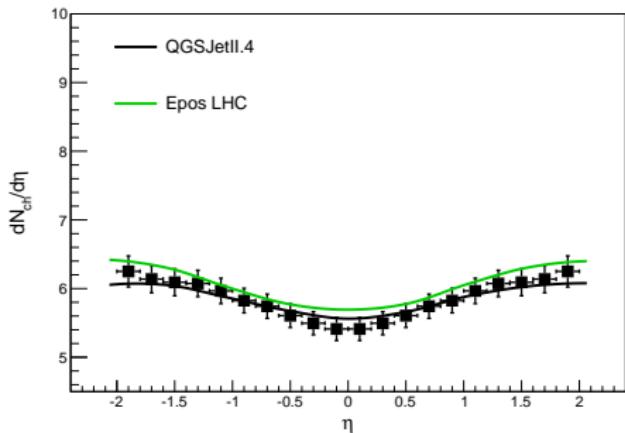


Models tuned at 7 TeV

CMS 13TeV, Inelastic Events



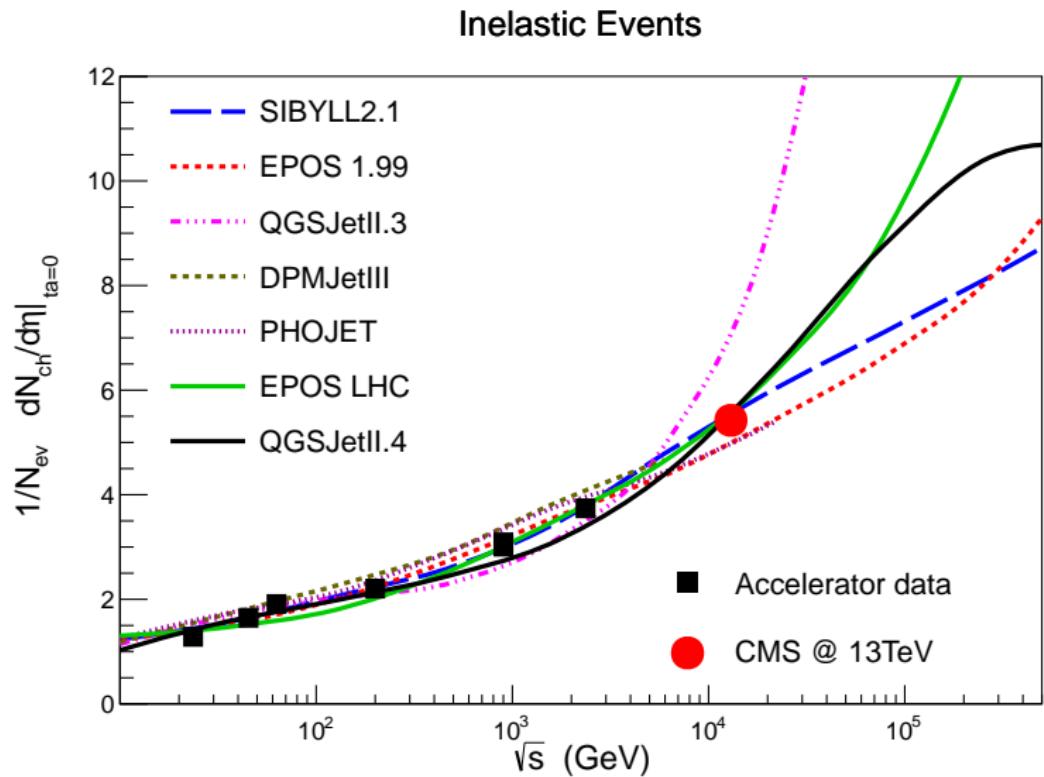
CMS 13TeV, Inelastic Events



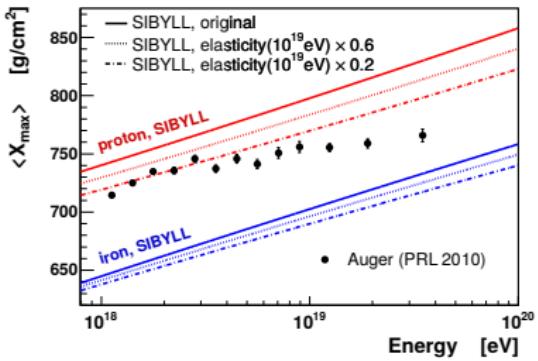
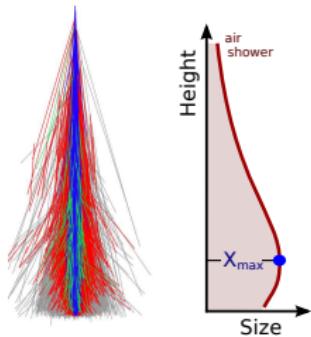
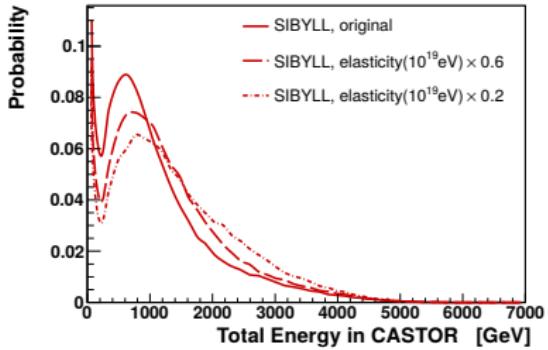
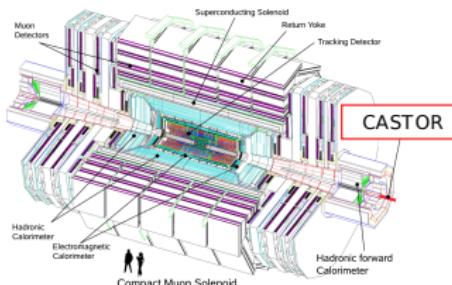
CMS: PLB 751 (2015) 143

- Good extrapolation over factor ≈ 2 in \sqrt{s}
- Poorly constraint beyond factor > 10 in \sqrt{s}

Extrapolation to ultra-high energies



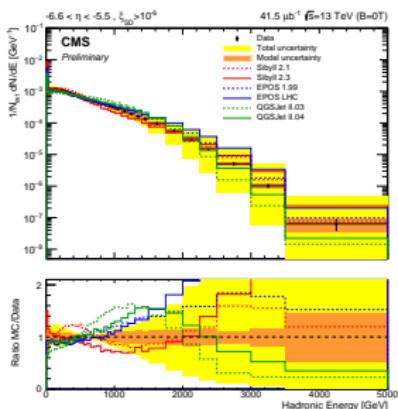
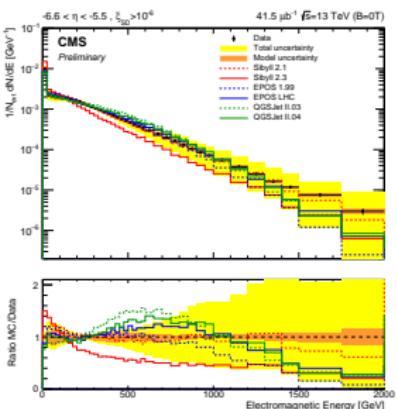
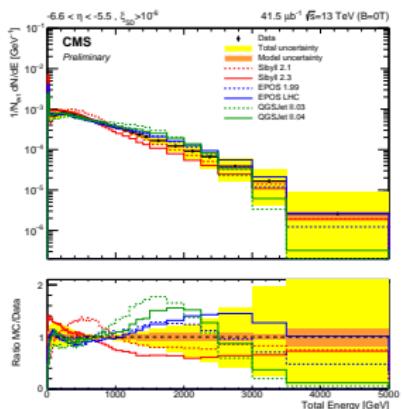
CASTOR and cosmic ray interpretation



Total energy

Electromagnetic energy

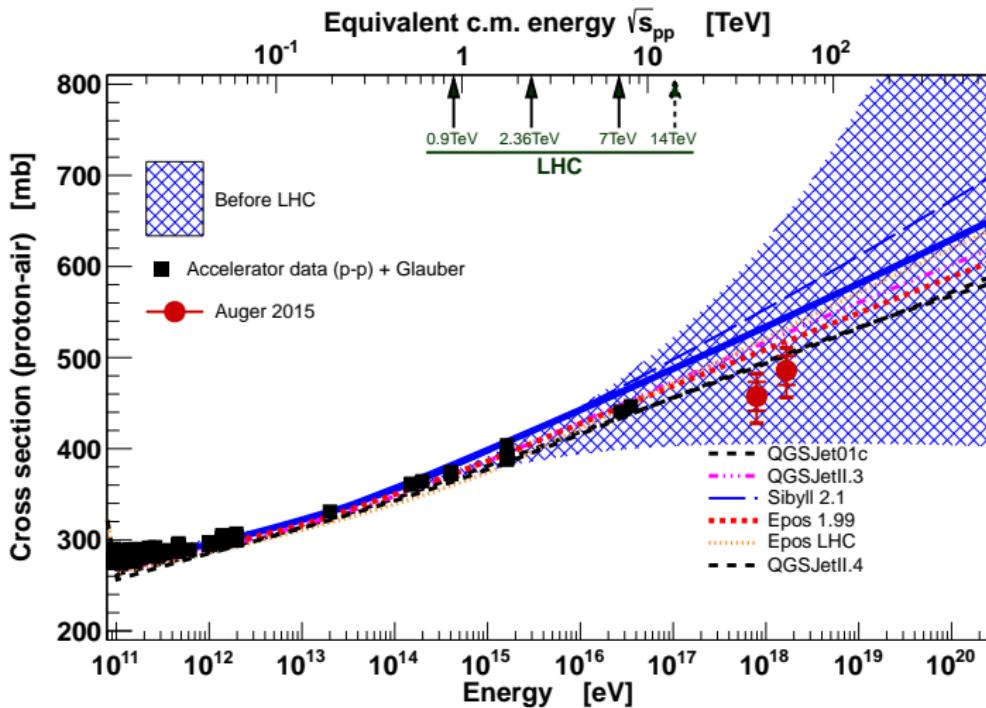
Hadronic energy



JHEP 08 (2017) 046

- Model performance tested in crucial phase-space domain
- Very sensitive to MPI, diffraction, low-x structure
- Represents a significant part of the overall energy-flow

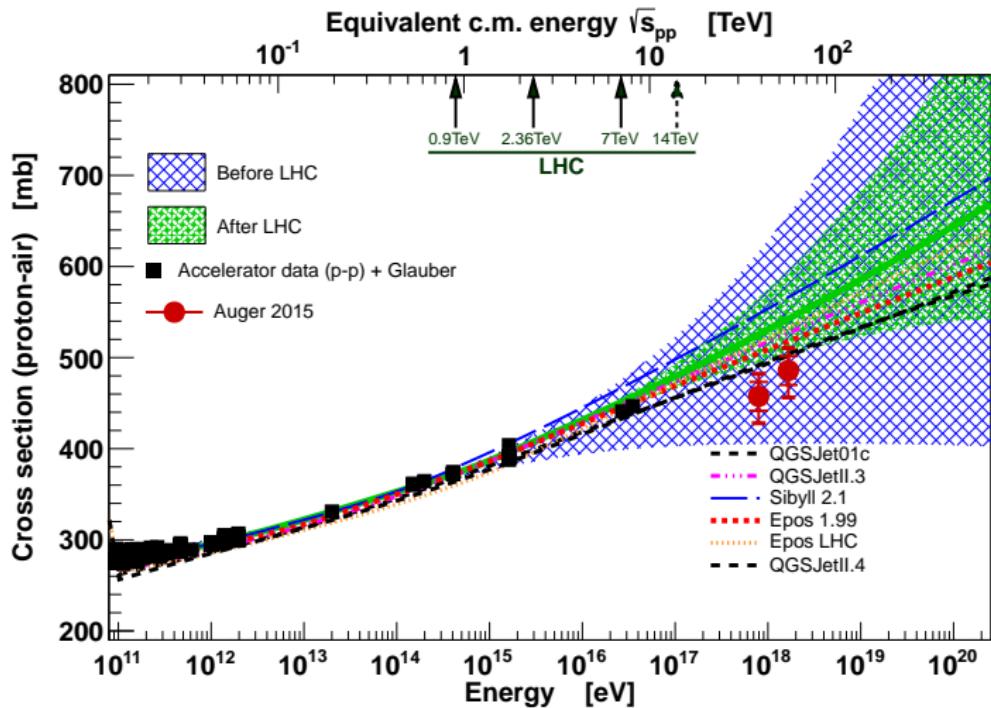
Proton-Air cross section, with Tevatron data



compare to Nucl.Phys.Proc.Suppl. 196 (2009) 335

Large uncertainties due to nuclear effects from pp to p-air

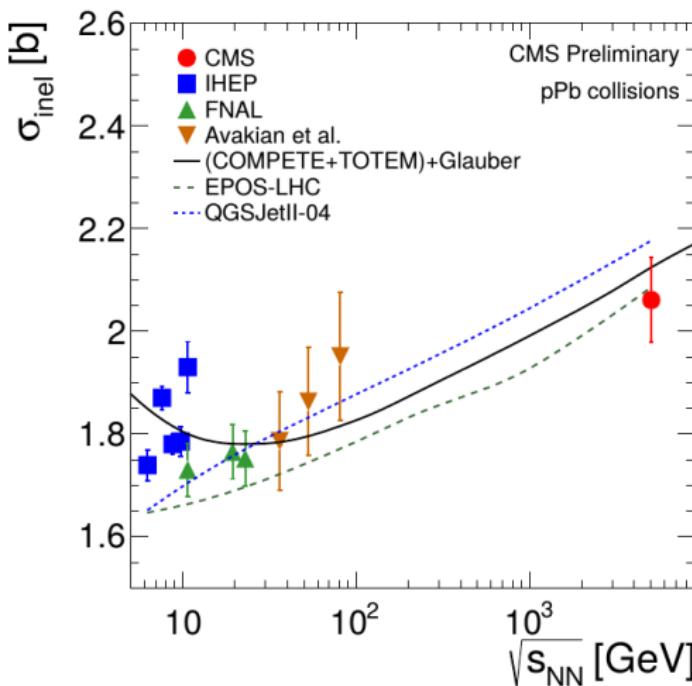
Proton-Air cross section, with LHC data



⇒ Sign of a clear relevant improvement

Large uncertainties due to nuclear effects from pp to p-air

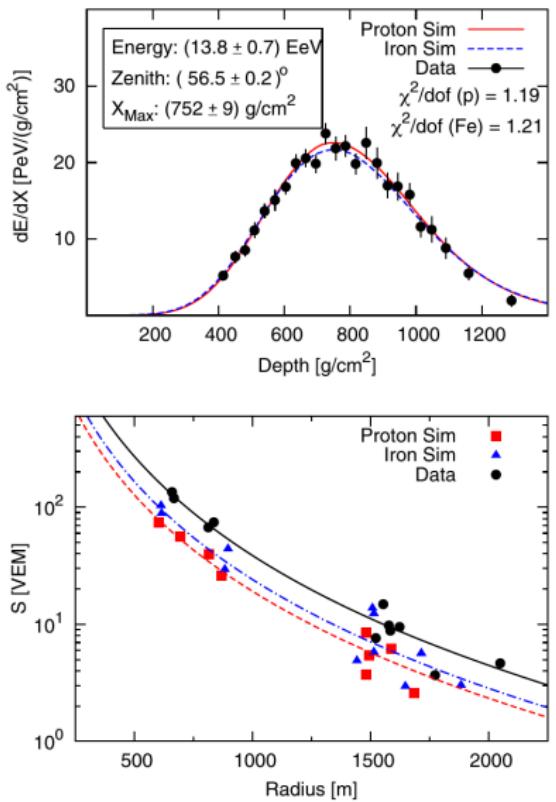
Inelastic proton-lead cross section at 5.02 TeV



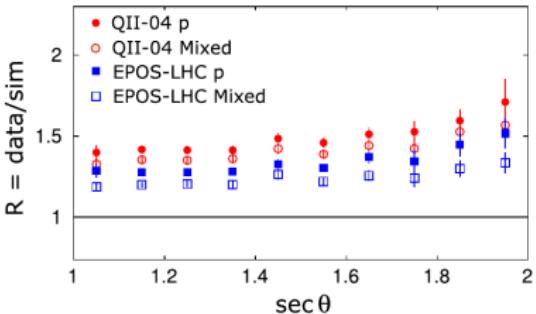
PLB 759 (2016) 641

- Direct test of Glauber model (and related physics) at LHC
- proton-oxygen will be much more powerful.

Hadronic interactions, Auger

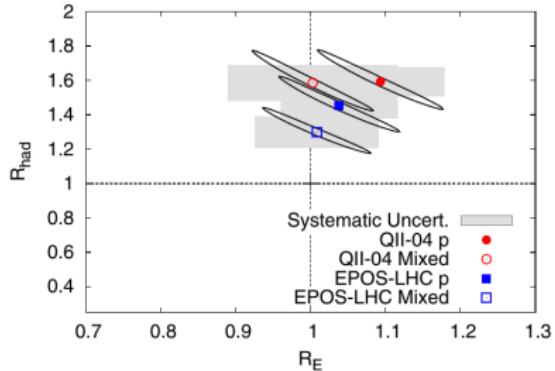
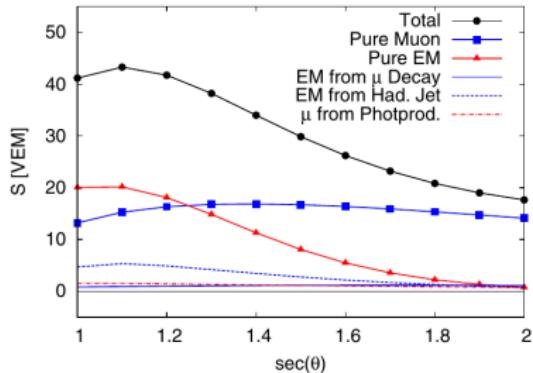


Attempt of consistent description of longitudinal and lateral shower data.



PRL 117, 192001 (2016)

Hadron/Muon component in data is too large



PRL 117, 192001 (2016)

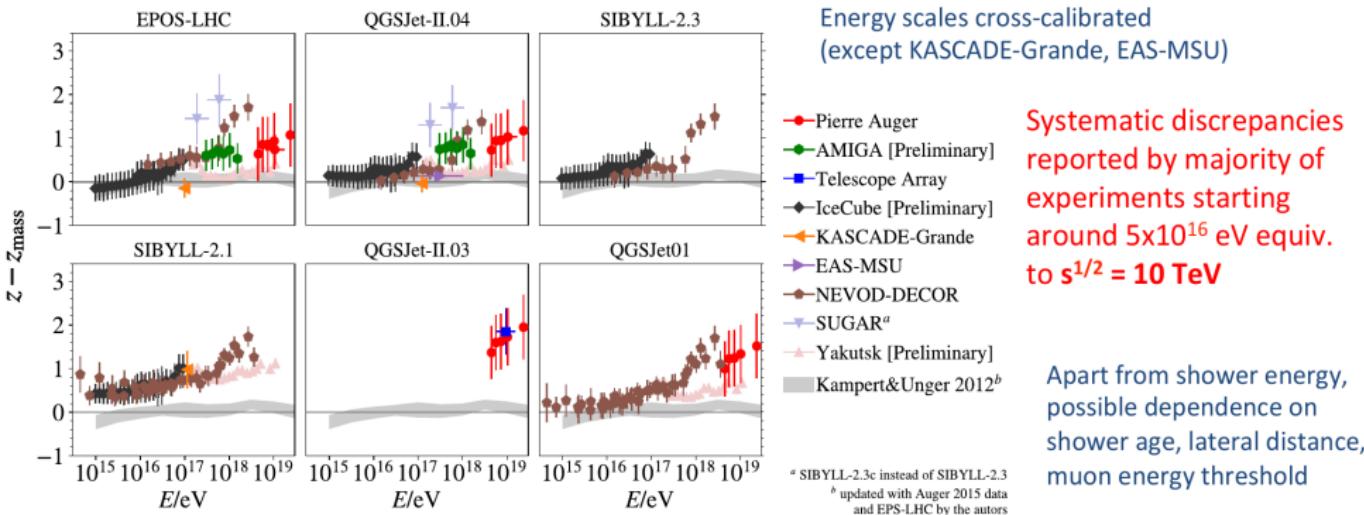
- Scale E.M. and had. part of MC showers with R_E and R_{had} to fit data:

$$S_{\text{resc}}(R_E, R_{\text{had}}) = R_E S_{\text{EM}} + R_{\text{had}} R_E^\alpha S_{\text{had}}$$

- While $R_E = 1$ is possible and mostly consistent with data
- R_{had} is significantly above 1
- None of the models/assumptions reproduces data

Compilation of muon measurements

- WHISP report at UHECR 2018 conference, Oct 8-12 2018
- Comprehensive compilation of muon measurements from air shower experiments

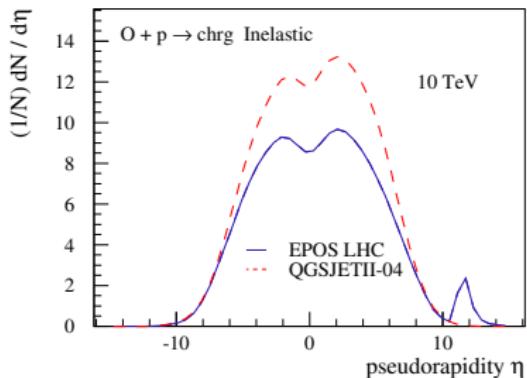
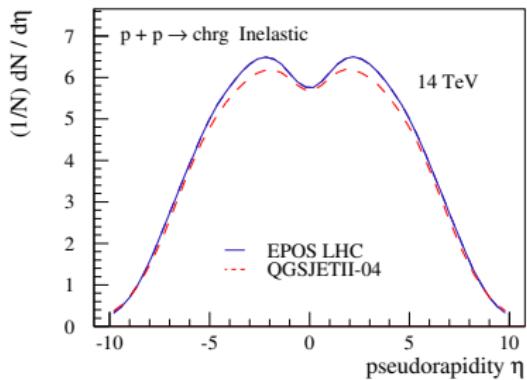


Not simple to fix:

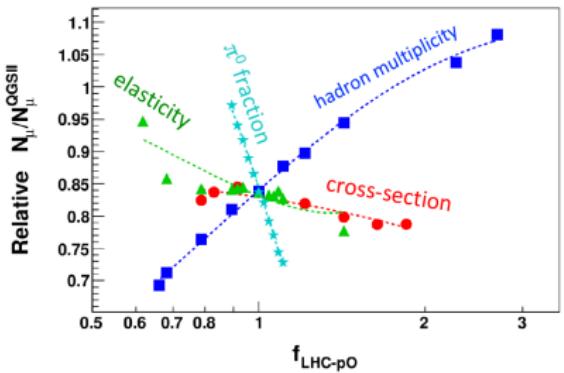
- No simple key measurement
- Need to accurately know and extrapolate several features
- Focus on precise measurements of features in reference systems, use models to predict interactions in unavailable target systems and phase space

Proton-Oxygen Data at LHC: Very Relevant

- Asymmetric heavy-ion run with proton-oxygen nuclei
- LHC Run3 $\sqrt{s_{NN}^{pO}} = 10 \text{ TeV}$
(Proton beam at 7 TeV)
- Oxygen very close to atmospheric material of extensive air shower production (nitrogen)
- Impact on model predictions :

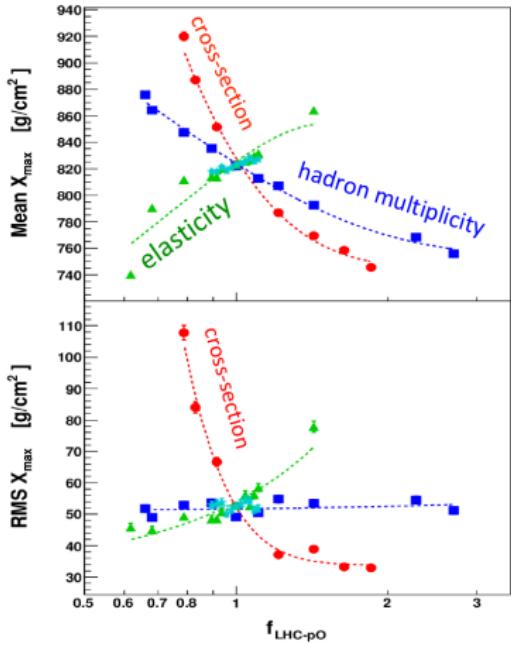


Potential of proton-oxygen data



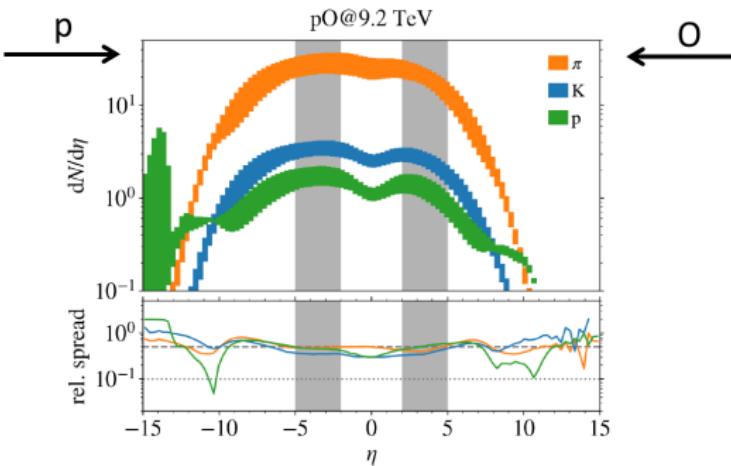
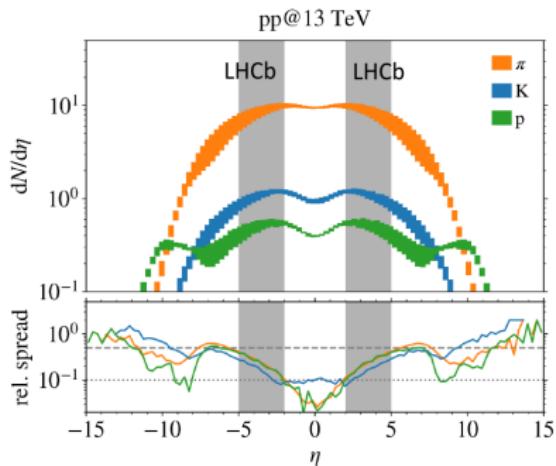
Remove uncertainties from nuclear effects.

Increase accuracy of air shower simulations.



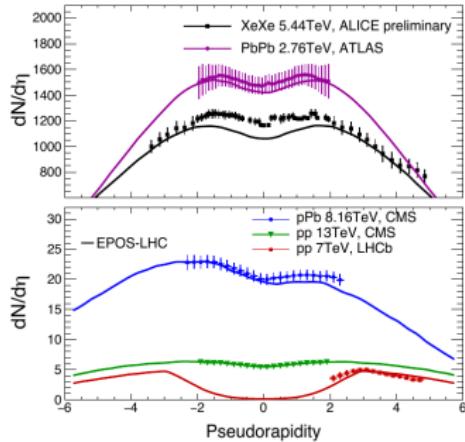
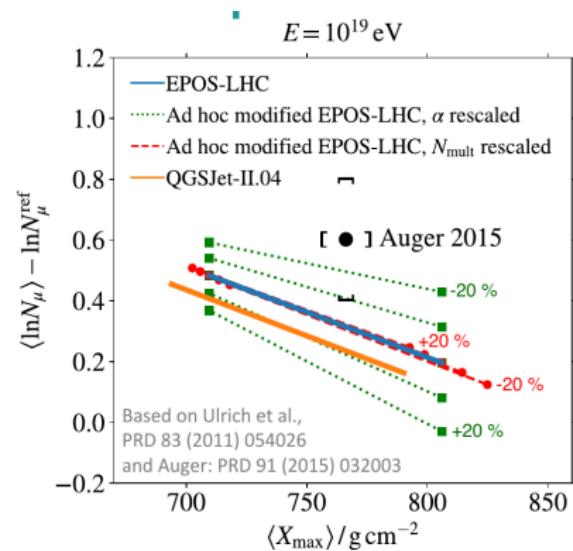
Extrapolation as one problem

Shown is spread between: EPOS-LHC, QGSJetII.04, SIBYLL-2.3



Models mostly tuned to p+p data at $|\eta| < 2$: p+p 10 % model spread, p+O 50 % model spread

Impact of LHC measurements

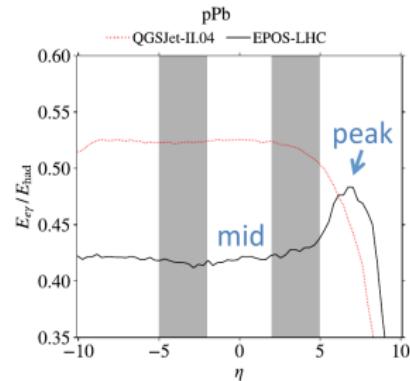
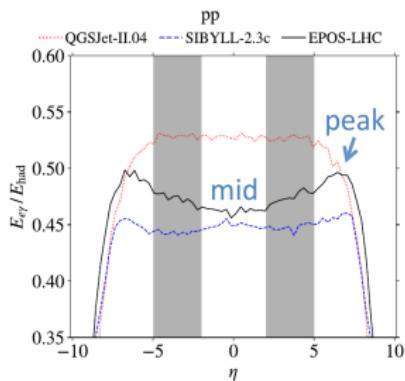
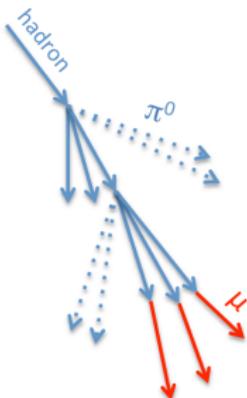


ALICE Xe-Xe arXiv:1807.09061; ATLAS Pb-Pb arXiv:1504.04337; CMS p-Pb arXiv:1710.09355v2; CMS p-p arXiv:1507.05915v2; LHCb p-p arXiv:1402.4430

- X_{\max} sensitive to: **inelastic cross-section**, hadron multiplicity
- N_μ sensitive to: **energy fraction lost to π^0** , hadron multiplicity
- **Nuclear modification in forward-produced hadrons** expected and important

Energy flow: e/γ vs. hadrons

- Hadronic energy “lost” to π^0 s cannot produce muons in late shower
- “Energy loss” described by observable $E_{e\gamma}/E_{\text{hadrons}}$

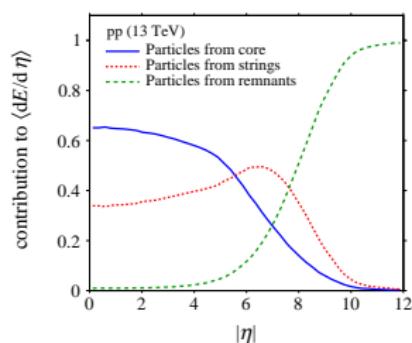


- Model predictions differ by **15 %** and in **shape**: only EPOS has forward peaks
- Translates to about **20 % shift** in N_μ → **high impact on Muon Puzzle**

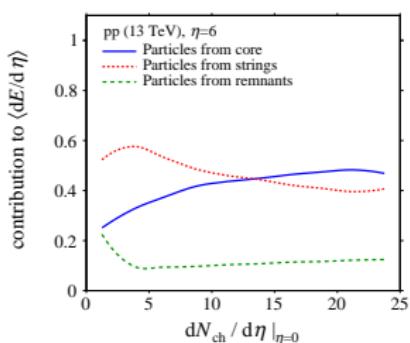
Mechanisms for different EM-ratios

Secondaries are produced in different mechanisms in different phase space regions:

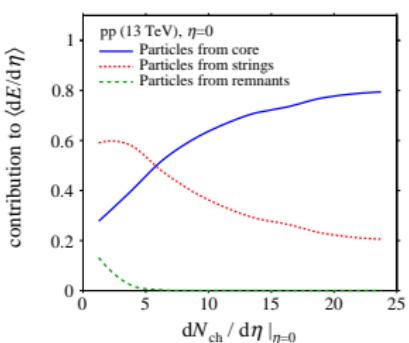
vs η



vs N_{mult} ($\eta = 0$)



vs N_{mult} ($\eta = 6$)



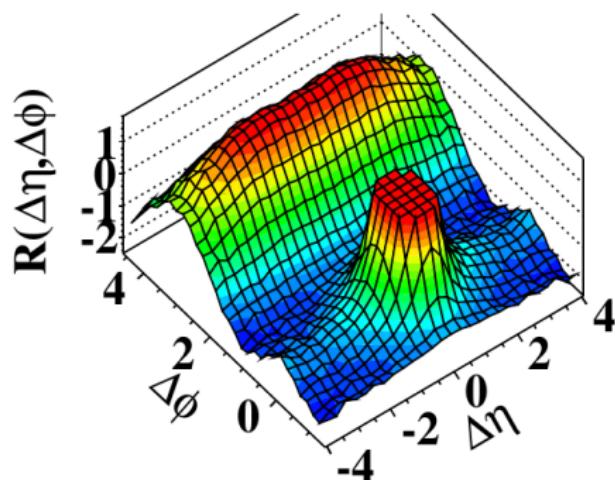
arXiv:1902.09265 [hep-ph]

And: each mechanism has its own EM-ratio ($R = E_{\text{e.m.}}/E_{\text{had}}$).

Collective effects in small systems

Observation of Long-Range, Near-Side Angular Correlations in Proton-Proton Collisions at the LHC:

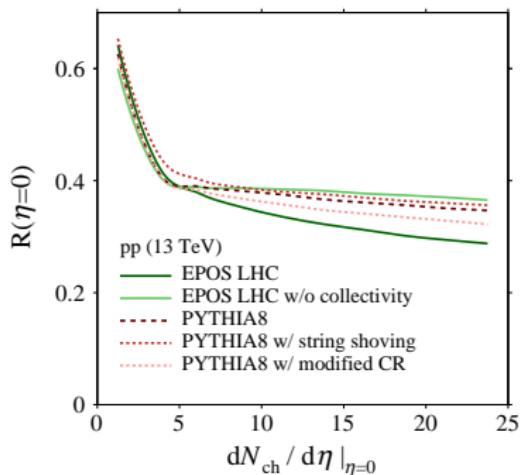
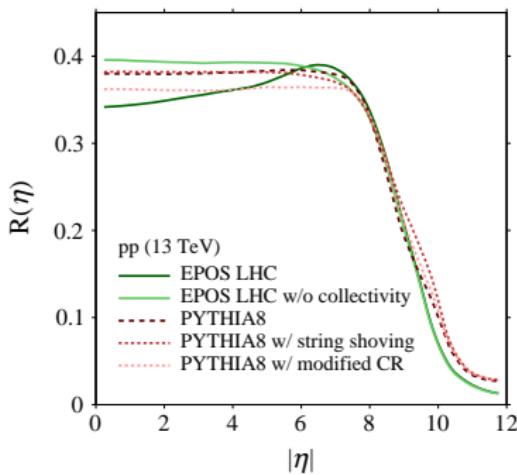
(d) CMS $N \geq 110, 1.0\text{GeV}/c < p_T < 3.0\text{GeV}/c$



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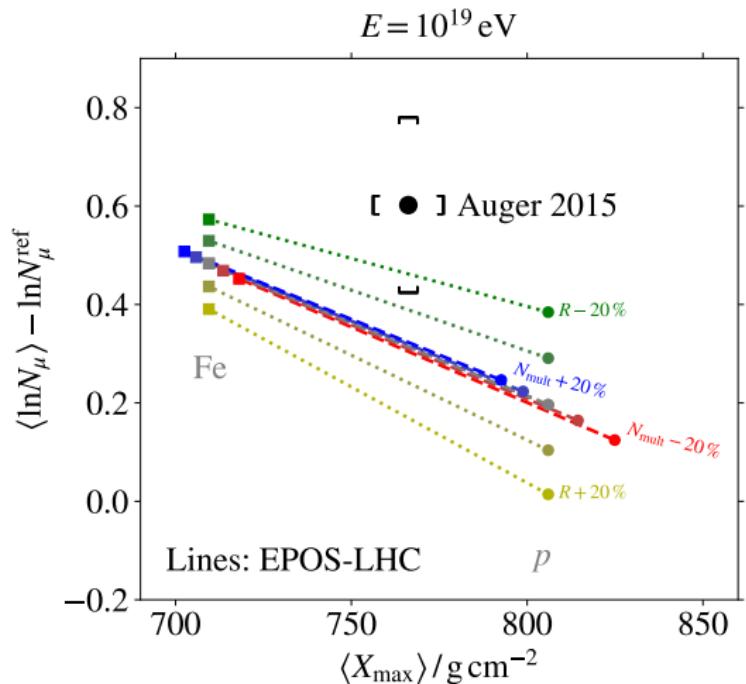
A rare discovery at LHC!

Potential (further) signatures at LHC



Could be measured at LHC → shape analyses

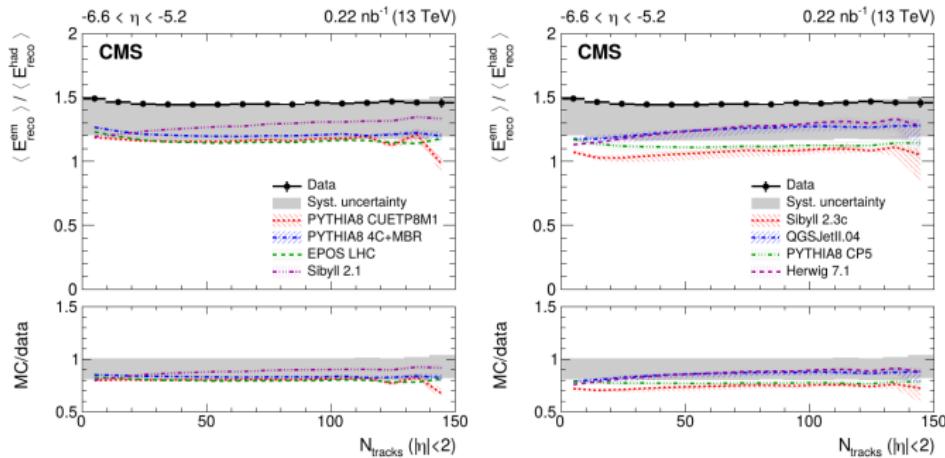
Relation to muon mystery



Orthogonal to many other effects → important for muon mystery

Measurement at CMS/CASTOR

Measurement of the average very forward energy as a function of the track multiplicity at central pseudorapidities in proton-proton collisions at 13 TeV:



CMS, arXiv:1908.01750 [hep-ex], submitted

LHC shopping list

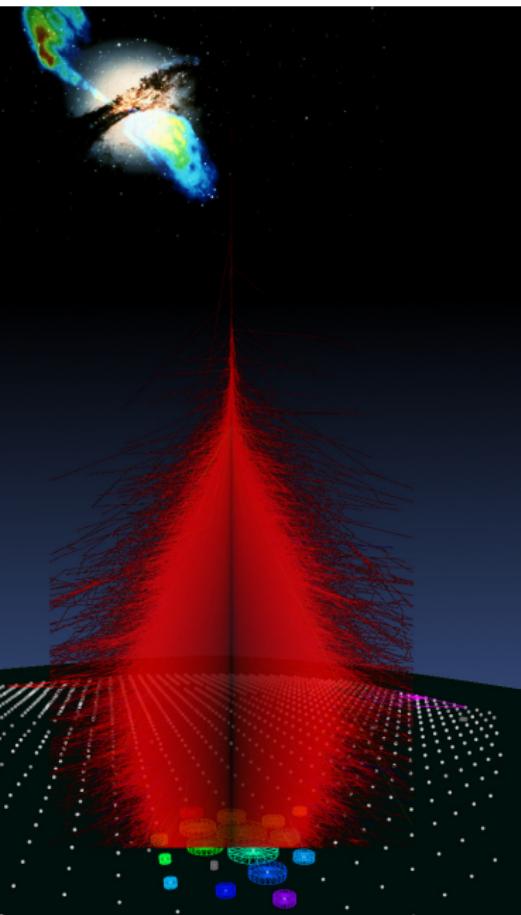
At 13 TeV

- Total and elastic cross section, elastic slope, imaginary to real part of amplitude (rho)
- Soft diffraction, rapidity gaps, characteristics of “pomeron” exchange
- Mini-jets, PID spectra, correlations
- PDF constraints (W^\pm , Z , ...)
- LHCf zero-degree π^0 and neutrons
- Underlying event with CASTOR, CMS+TOTEM $dN/d\eta$

Finally

- LHC energy scan: 0.9 GeV → large x_F -range
- Nuclear scan: proton-lead, proton-oxygen
- Maximum acceptance, dedicated detector components
- Pion as beam particle

Summary



LHC impact on UHECR interpretation

- Models are constraint much better
- Predictions tend to converge (overall)
- Different observables have very **different sensitivity** to interaction uncertainties
- Muon mystery may be related to a **few percent** of unaccounted physics per collision
- Models so far tuned at LHC roughly to a level of **about 10 %**
- **Forward detectors** ultimately of paramount importance
- Important for air showers: energy AND **phase space** !
- Nuclear effects are another important source of uncertainties. **Need oxygen in LHC.**