# Collapse instability and staccato decay of oscillons in various dimensions

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#### In memoriam

Gy. Fodor, P. Forgács, <u>Z. Horváth</u>, and Á. Lukács: *Small amplitude quasi-breathers and oscillons*, Phys. Rev. D78: 025003, 2008.

Gy. Fodor, P. Forgács, <u>Z. Horváth</u>, and M. Mezei: *Computation of the radiation amplitude of oscillons*, Phys. Rev. D79: 065002, 2009.

Gy. Fodor, P. Forgács, <u>Z. Horváth</u>, and M. Mezei: *Radiation of scalar oscillons in 2 and 3 dimensions*, Phys. Lett. B674: 319-324, 2009.

Gy. Fodor, P. Forgács, <u>Z. Horváth</u>, and M. Mezei: *Oscillons in dilaton-scalar theories*, JHEP 2009(8): 106, 2009.



Zalán Horváth (1943-2011)

# Outline of talk

- 1. Introduction
- 2. Oscillons and quasi-breathers
- 3. Decay mechanisms of oscillons
- 4. Collapse instability in dimension D
- 5. Staccato decay in dimension D
- 6. Conclusions and outlook

#### Introduction

Oscillons: metastable scalar field lumps with extraordinarily long life-time

Nonlinear scalar field theory in D+1 dimensions

$$\mathcal{L}=rac{1}{2}\left(\partial_t\phi
ight)^2-rac{1}{2}\left(
abla\phi
ight)^2-m{V}(\phi)$$

Assume:  $V(\phi)$  has a minimum in  $\phi = 0$  Mass scale:  $m^2 = V''(0)$ 

Energy:

$$\mathcal{E} = \int d^D x \left\{ rac{1}{2} \left( \partial_t \phi 
ight)^2 + rac{1}{2} \left( 
abla \phi 
ight)^2 + V(\phi) 
ight\}$$

Equation of motion:

$$\left(\partial_t^2 - \nabla^2\right)\phi = -V'(\phi)$$

Assuming spherical symmetry:

$$\left(\partial_t^2 - \partial_r^2 - \frac{D-1}{r}\partial_r\right)\phi = -V'(\phi)$$
$$\mathcal{E} = \frac{2\pi^{D/2}}{\Gamma(D/2)}\int dr r^{D-1}\left\{\frac{1}{2}\left(\partial_t\phi\right)^2 + \frac{1}{2}\left(\partial_r\phi\right)^2 + V(\phi)\right\}$$

# Example: $\phi^4$ theory with D = 3Potential:

$$V(\phi) = rac{1}{4}\phi^2 \left(\phi - 2
ight)^2 \Rightarrow m^2 = 2$$

Initial condition:

$$\phi(r,0) = C \exp\left(-\frac{r^2}{r_0^2}\right) \quad \dot{\phi}(r,0) = 0$$

Time evolution (our simulation with C = 2,  $r_0 = 2.39$ )





#### Numerical time evolution

Gy. Fodor: "A review on radiation of oscillons and oscillatons", arXiv: 1911.03340 [hep-th].

$$\partial_t^2 \phi = \left(\partial_r^2 + \frac{D-1}{r}\partial_r\right)\phi - V'(\phi)$$

$$\Downarrow \quad r = \frac{2R}{\kappa(1-R^2)} \quad \kappa = 0.05 \quad (\text{ maps } r \in [0,\infty) \text{ to } R \in [0,1) )$$

$$\partial_t^2 \phi = \frac{\kappa^2 \left(1-R^2\right)^3}{2\left(1+R^2\right)} \left[\frac{\left(1-R^2\right)}{2\left(1+R^2\right)}\partial_R^2 \phi - \frac{R\left(3+R^2\right)}{\left(1+R^2\right)}\partial_R \phi + \frac{D-1}{2R}\partial_R \phi\right] - V'(\phi)$$

Numerics: method of lines with Courant factor  $\Delta t/\Delta R = 1$ , with 4th order Runge-Kutta in t. Interval extended to  $R \in [-1, 1]$ ,  $\partial_R \phi|_{R=0} = 0$  enforced by  $\phi(R) = \phi(-R)$ . Choice of  $\kappa$ : tunes resolution of oscillon core vs. radiation tail. Protection against short-wavelength instabilities: add dissipation term

$$\mathcal{D} = K\left(\partial_R^6 \Phi\right) (\Delta R)^5$$

Independence of results on choice of K must be always checked.

# ${\sf Quasi-breathers}$

Oscillons slowly radiate and lose energy Can be compensated by adding a small standing wave tail: **quasi-breathers** 

For most of their life, the tail part is exponentially small. Typical radiation rate of small amplitude oscillons:

$$rac{dE}{dt} \sim rac{A}{\epsilon^{D-1}} \exp\left(-rac{B}{\epsilon}
ight)$$

where  $\epsilon$  is the amplitude.

As a result, the oscillon core evolves essentially adiabatically via a succession of quasi-breather core configurations.



The core and the tail part of a quasi-breather.

Figure from Gy. Fodor: *A review on radiation of oscillons and oscillatons*, arXiv: 1911.03340 [hep-th].

#### Constructing quasi-breathers numerically

Quasi-breather: exactly periodic solution

Ansatz:  

$$\phi(t, r) = \sum_{m=1}^{N} \phi_m(r) \cos(m\omega t) \qquad N: \text{ mode truncation}$$

$$\left(\partial_t^2 - \partial_r^2 - \frac{D-1}{r} \partial_r\right) \phi = -V'(\phi)$$

$$\downarrow$$

$$\left(n^2 \omega^2 + \partial_r^2 + \frac{D-1}{r} \partial_r\right) \phi = \frac{\omega}{\pi} \int_0^{2\pi\omega} dt V'\left(\sum_{m=1}^N \phi_m(r) \cos(m\omega t)\right) \cos(n\omega t)$$

 $\rightarrow$  coupled equations for the functions  $\phi_n(r)$ . Boundary conditions:

$$\begin{split} \partial_r \phi_n(r)|_{r=0} &= 0\\ \phi_n(r \to \infty) \sim r^{(1-D)/2} \exp\left(-\sqrt{1-n^2\omega^2}r\right) & n\omega < 1: \text{ localised modes}\\ \phi_n(r \to \infty) \sim r^{(1-D)/2} \sin\left(\sqrt{n^2\omega^2 - 1}r + \varphi_n\right) & n\omega > 1: \text{ radiation modes} \end{split}$$

Sine-Gordon theory in D=1: exact breathers

Sine-Gordon action

$$S = \int dt dx \left\{ \frac{1}{2} \left( \partial \phi \right)^2 + \mu \left( \cos \beta \phi - 1 \right) \right\} \qquad \phi_0^{(n)} = \frac{2\pi}{\beta} n \qquad m = \sqrt{\mu}\beta$$

Breather solution: exactly periodic, no radiation!

$$egin{aligned} \phi_B(t,x) &= rac{4}{eta} \arctan\left(rac{\epsilon\cos\left(m\omega t
ight)}{\omega\cosh\left(m\epsilon x
ight)}
ight) \ M_B(\epsilon) &= rac{16\sqrt{\mu}}{eta}\epsilon \qquad 0 < \epsilon < 1 \end{aligned}$$

Frequency-amplitude relation

$$\omega = \sqrt{1 - \epsilon^2}$$

For a general D and  $V(\phi)$ , it is always possible to define the amplitude parameter  $\epsilon$  such that this relation is exact, and perform a small amplitude expansion of the oscillon and quasi-breather solutions:

Gy. Fodor, P. Forgács, Z. Horváth, and Á. Lukács: *Small amplitude quasi-breathers and oscillons*, Phys. Rev. D78: 025003, 2008.

# Mechanism of oscillon longevity

 $\omega <$  1: fundamental mode is localised  $\rightarrow$  only higher harmonics radiate, but: they are suppressed!

Oscillon with core size  $d(\sim \epsilon^{-1}) \gg \lambda_{rad}$ : geometric decoupling

There are also windows of longevity due to destructive interference between outgoing radiation from higher harmonics.

Note: for even potentials

$$V(\phi) = V(-\phi)$$

only odd harmonics exist!



Oscillon fundamental frequency  $\omega$ 

Axion theory (sine-Gordon in D = 3)

Figure from D. Cyncynates and T. Giurgica-Tiron, Phys. Rev. D103: 116011, 2021.

Sine-Gordon in D = 1: complete destructive interference of radiation - no decay!

#### Sudden collapse a.k.a. energetic death Energy stored in quasi-breather core

$$E = \frac{2\pi^{D/2}}{\Gamma(D/2)} \int_0^{R_{\text{core}}} dr r^{D-1} \left\{ \frac{1}{2} \left( \partial_t \phi \right)^2 + \frac{1}{2} \left( \partial_r \phi \right)^2 + V(\phi) \right\}$$

Radius of core  $R_{core}$ : dominant localised vs radiation mode amplitudes are equal.



Core energy as a function of  $\omega$  for  $V(\phi) = \frac{1}{4}\phi^2 (\phi - 2)^2$ , D = 3

Figure from Gy. Fodor: *A review on radiation of oscillons and oscillatons*, arXiv: 1911.03340

#### Sudden death: $\omega \approx 1.368 \approx 0.967 m$



#### Staccato decay in D = 1

P. Dorey, T. Romańczukiewicz and Y. Shnir: *Staccato radiation from the decay of large amplitude oscillons*, Phys. Lett. B806: 135497, 2020.

Model:

$$\mathcal{L} = \frac{1}{2} (\partial \phi)^2 - V_{\lambda}(\phi)$$
$$\ell_{\lambda}(\phi) = (1 - \lambda)(1 - \cos \phi) + \frac{\lambda}{8\pi^2} \phi^2 (\phi - 2\pi)^2$$
$$\phi(0, x) = 4 \arctan\left(\frac{\epsilon \cos(\omega t)}{\omega \cosh(\epsilon x)}\right)$$
$$\dot{\phi}(0, x) = 0$$
$$\epsilon = \sqrt{1 - \omega^2}$$

Set coupling  $\lambda = 0.0025$ , start evolution with  $\omega = 0.1$ 



## Staccato decay in D = 1

Staccato decay proves to be a robust feature in 1+1 D, e.g. in two-kink collisions



P. Dorey, T. Romańczukiewicz and Y. Shnir, Phys. Lett. B806: 135497, 2020.

#### Our research plan

- Construct quasi-breathers with given initial frequency in spatial dimension  $D\in\mathbb{R}$ 

This is necessary to

- Control initial frequency
- Optimize life-time of oscillon, minimize shape modes
- $\blacktriangleright$  Determine energy-frequency diagram, sudden collapse as a function of D
- Study whether staccato decay persists for D > 1

#### Constructing quasi-breathers

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} \left( \partial \phi \right)^2 - V_{\lambda}(\phi) \\ V_{\lambda}(\phi) &= (1 - \lambda)(1 - \cos \phi) + \frac{\lambda}{8\pi^2} \phi^2 \left( \phi - 2\pi \right)^2 \\ \text{Ansatz:} \qquad \phi(t, r) &= \sum_{n=1}^{N} \phi_n(r) \cos \left( n \omega t \right) \end{aligned}$$

Initialising solution:

$$D = 1: \quad \phi_n(0) \approx \phi_n^{sG}(0)$$
$$= \frac{\omega}{\pi} \int_0^{2\pi/\omega} 4 \arctan\left(\frac{\epsilon \cos(\omega t)}{\omega}\right) \cos(n\omega t)$$
$$D > 1: \quad \phi_n^{D+\delta D}(0) \approx \phi_n^D(0)$$

Cutting off the tail of radiation modes:

$$\phi_n(r) \to \phi_n(r) \cdot e^{-\sqrt{n^2 \omega^2 - 1(r - r_0)}}$$
  $r_0 = 10/\sqrt{1 - \omega^2}$ 

B. C. Nagy and G. Takács, Phys. Rev. D 104: 056033, 2021.



#### Sudden collapse in D spatial dimensions

B. C. Nagy and G. Takács, Phys. Rev. D 104: 056033, 2021.

D > 2: there is a minimum  $\Rightarrow$  critical frequency  $\omega_c$ 

D	2.25	2.50	2.75	3.00
$\omega_c$	0.9591	0.9429	0.9381	0.9395

The disappearance of the minimum at D = 2 was previously derived in the small amplitude expansion by. Gy. Fodor et al in 2008.



#### Sudden collapse: numerical time evolution



# Starting with a supercritical frequency $\omega > \omega_c$



It is possible to switch between these behaviours by fine-tuning initial data G. Fodor, P. Forgács, P. Grandclément, and I. Rácz, Phys. Rev. D 74 (2006) 124003.

#### Reproducing staccato decay in D = 1 with our methods





Staccato decay in D > 1?



- oscillon decay speeds up (note time scale!
- staccato steps become indistinct (although first burst at ω = m/2 is still observable at D = 1.20)

#### Time passed until staccato steps in D > 1



- ▶ The time passed until staccato steps decreases with D
- When it is close to  $2\pi/\omega$ : staccato steps become unobservable
- Unobservability happens before reaching the next physical dimension D = 2, even for  $\lambda = 0$  (sine-Gordon/axion theory)

Sadly, it looks likely that staccato decays are confined to D = 1. Exceptional stability islands (destructive interference of radiation!) might help, but: known cases are not suitable - all have  $\omega > m/2$ .

# Conclusions and outlook

- 1. Oscillons: long lived metastable configuration of interacting scalar field
  - a) Small amplitude: analytic results for life time etc.
    - Frequency:  $\omega = m\sqrt{1-\epsilon^2} \epsilon$ : amplitude parameter
  - b) Large amplitude: interesting, novel dynamics
- 2. D > 2: critical frequency
  - a)  $\omega_0 < \omega_c$ : slow radiative evolution followed by sudden decay
  - b)  $\omega_0 > \omega_c$ : immediate decay, can be suppressed by fine-tuning
- 3.  $D \leq 2$ : no critical frequency, slow radiative evolution
  - a) D=1: staccato decay, very robust feature when  $\omega_0 < m/2$
  - b) D > 1: staccation decay disappears well below D = 2
- 4. Ongoing work
  - a) Decay rate in quantum theory Small amplitudes: analytic results
     M.P. Hertzberg: *Quantum radiation of oscillons*, Phys. Rev. D82: 045022.
     Question: what happens for large amplitudes?
  - b) Staccato decay at the quantum level?

Work was done in collaboration with Botond Nagy (MSc student)



Results appeared in:

B. C. Nagy: Oszcillonok dinamikája, TDK paper, Oct. 2020.

B. C. Nagy and G. Takács: *Collapse instability and staccato decay of oscillons in various dimensions*, Phys. Rev. D 104: 056033, 2021.

# Thank you for your attention!